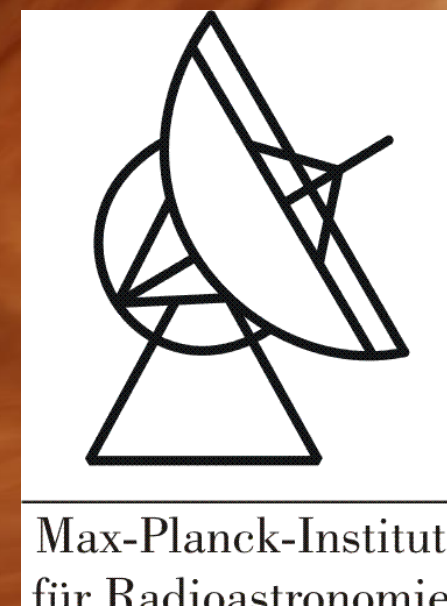


Ray-Tracing in 3D Relativistic Magnetohydrodynamic Jet Simulations

Joana Kramer

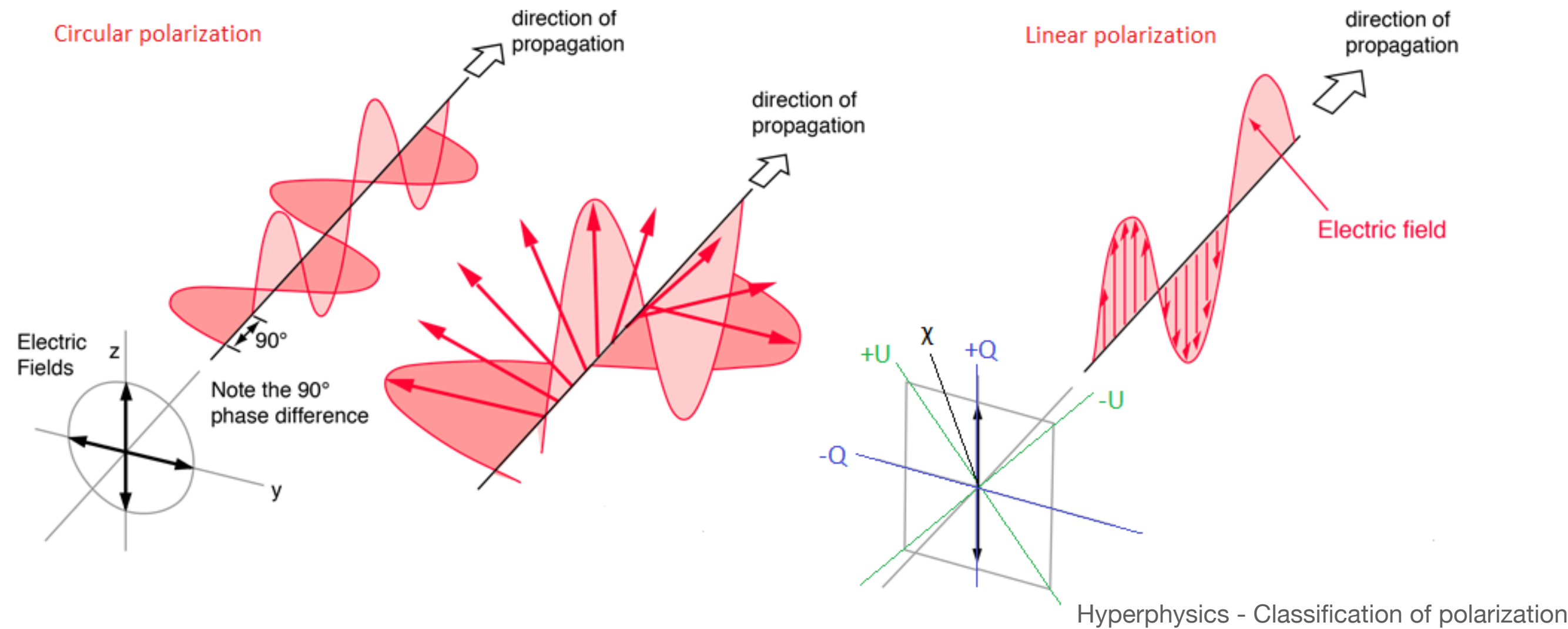
jkramer@mpifr-bonn.mpg.de

Jets2021



IMPRS
astronomy &
astrophysics
Bonn and Cologne

Polarized Emission, *PLUTO*, and *RADMC-3D*



Stokes

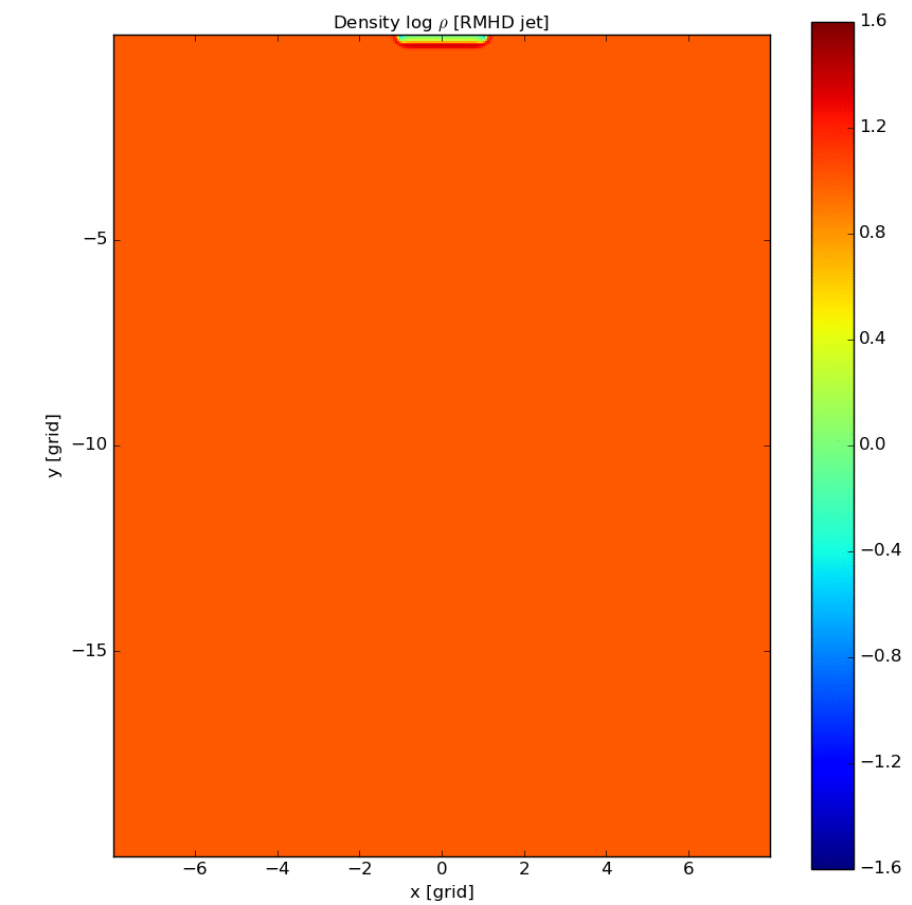
$$I = \epsilon_1^2 + \epsilon_2^2$$

$$Q = \epsilon_1^2 - \epsilon_2^2$$

$$U = 2\epsilon_1\epsilon_2 \cos(\Phi_1 - \Phi_2)$$

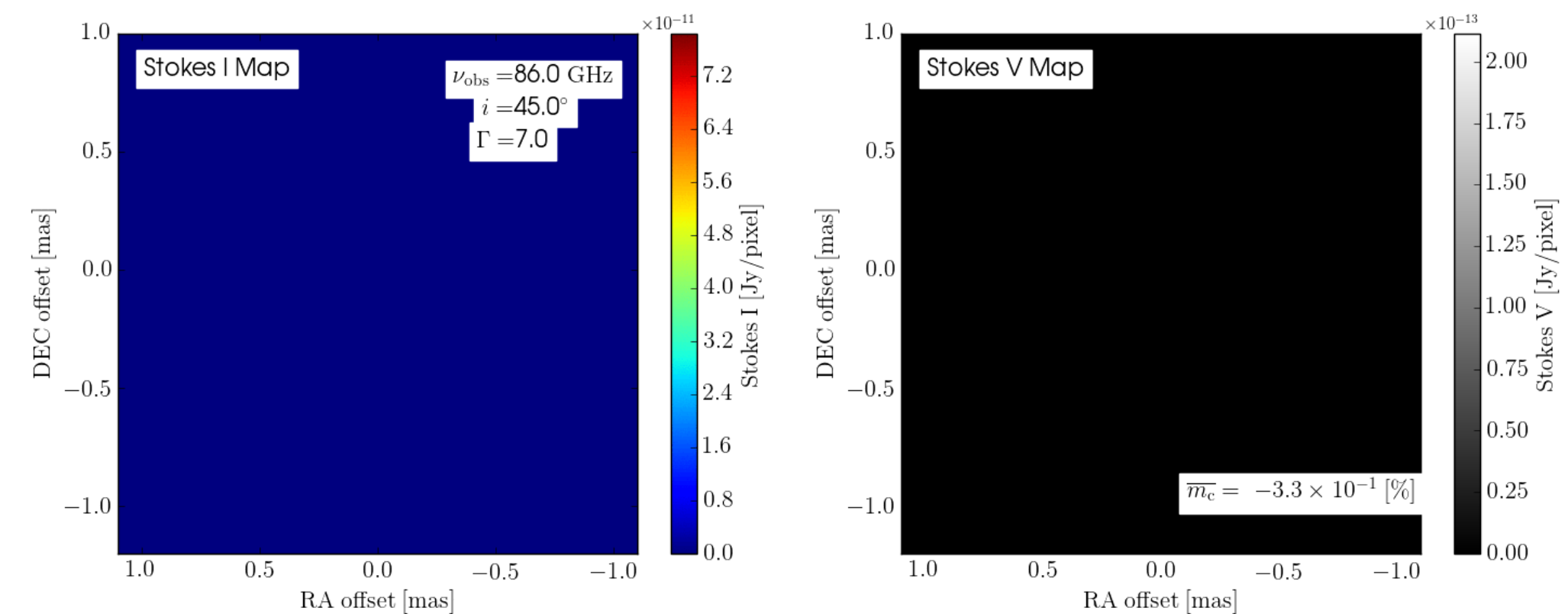
$$V = 2\epsilon_1\epsilon_2 \sin(\Phi_1 - \Phi_2)$$

- **For the first time ever:** we perform a full Stokes analysis including analyzing the circular polarized emission

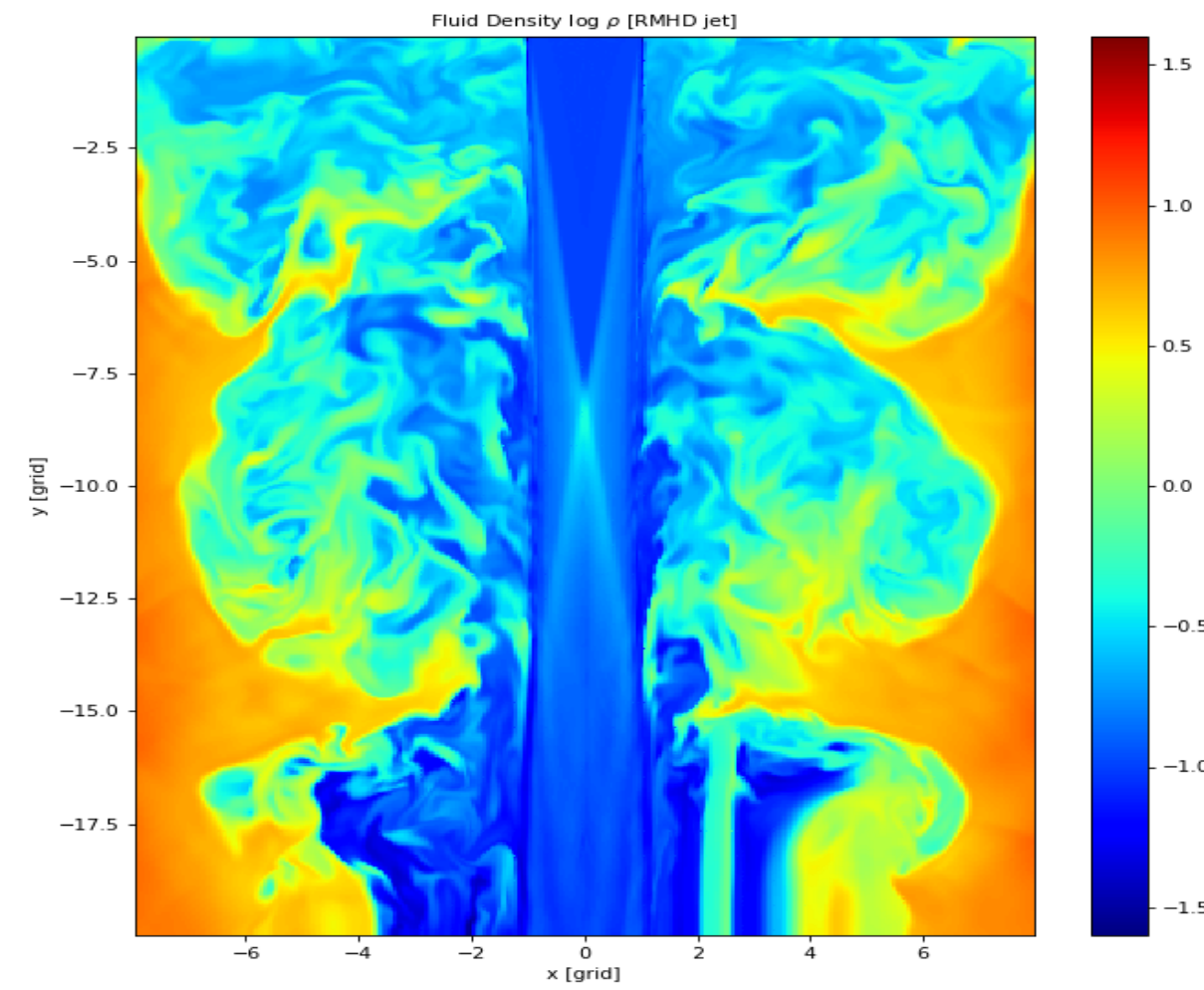


- ***PLUTO*:** solves a time-dependent non-linear system of special relativistic conservation laws (mass, energy momentum) and Maxwell's equations. Mignone et al. 2007, © <http://plutocode.ph.unito.it/documentation.html>

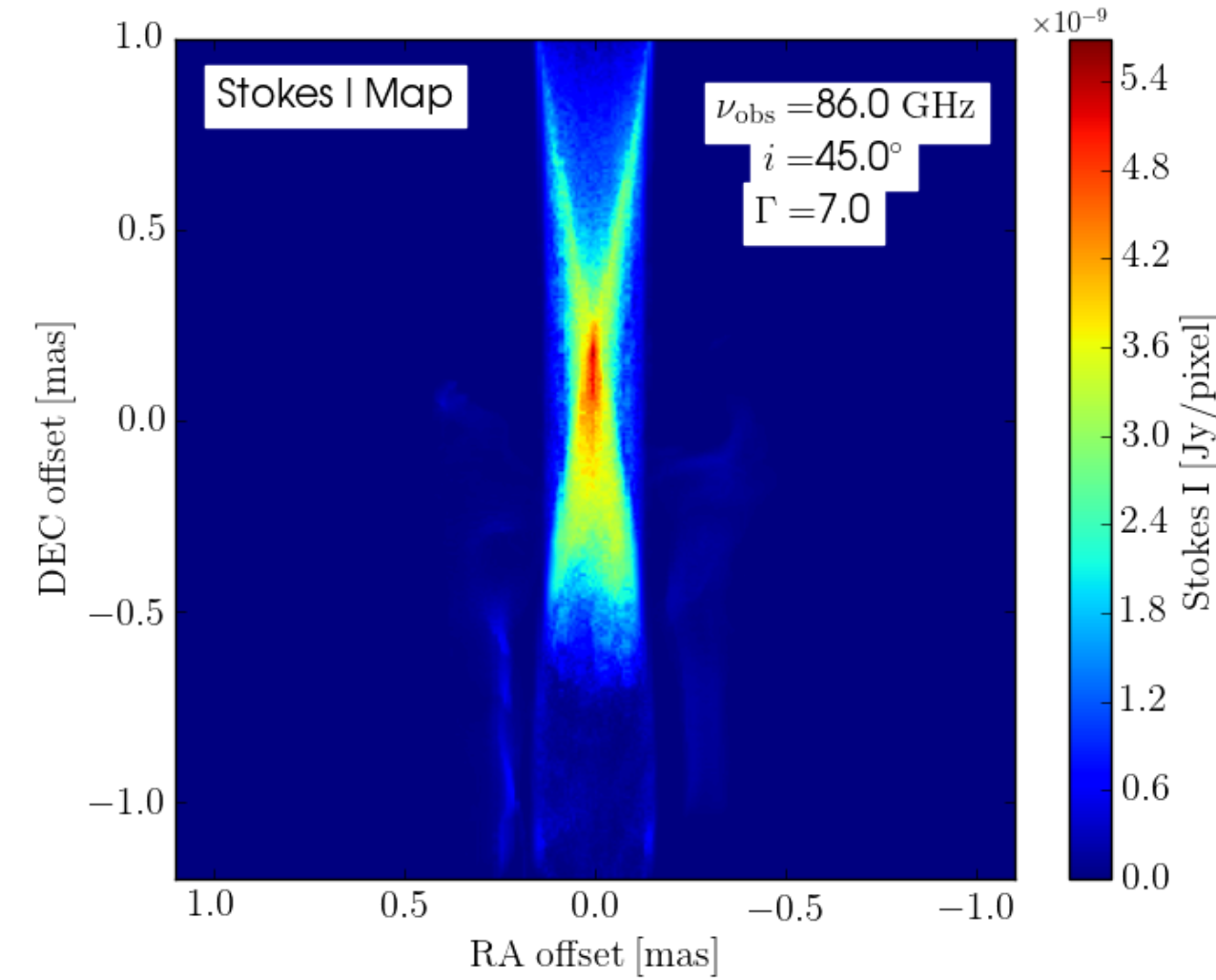
- ***RADMC-3D*:** solves the full Stokes equations of polarized radiative transfer and will create synthetic maps for each Stokes parameter



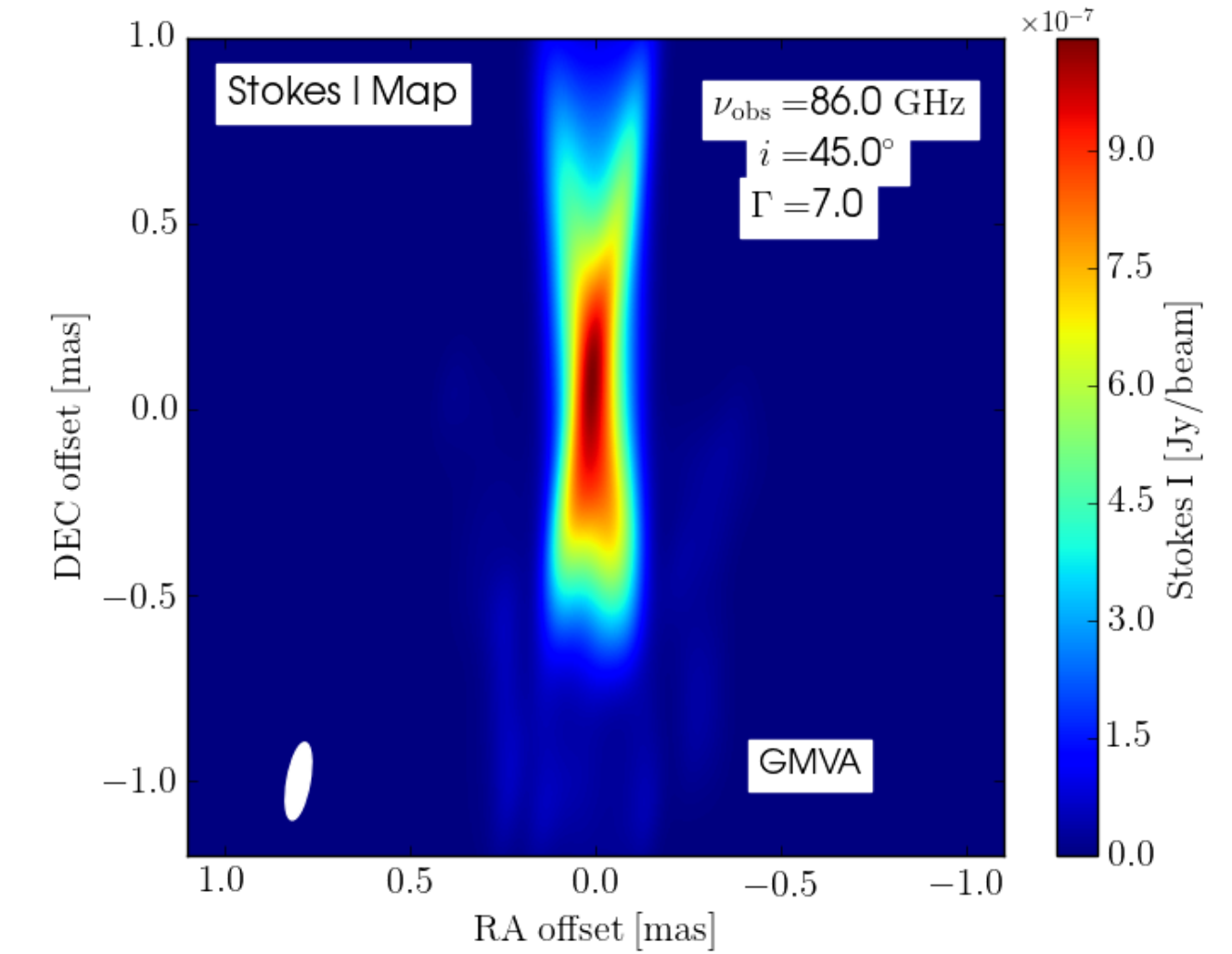
Magnetic Field Morphology and Emission Recipe Study



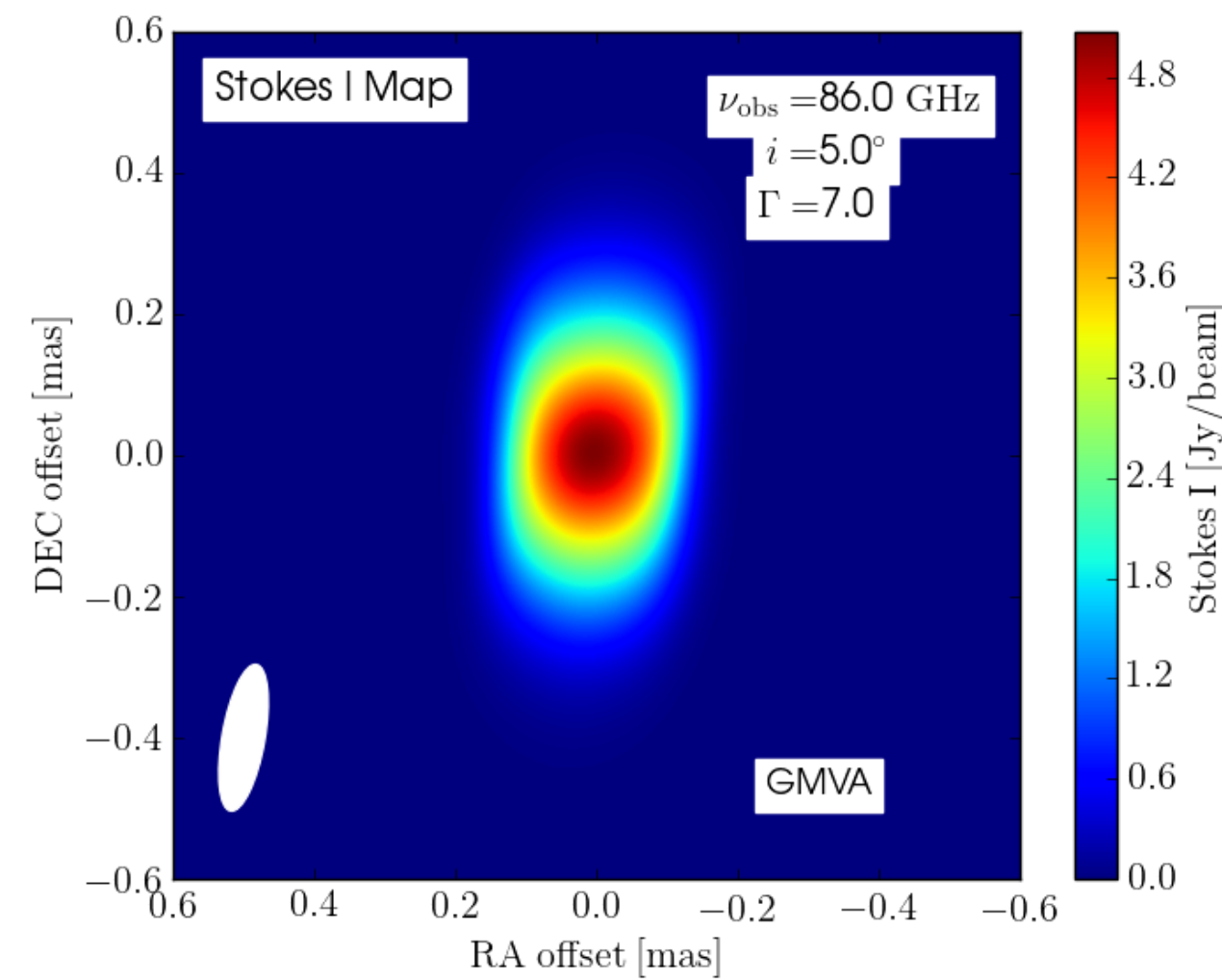
(a)



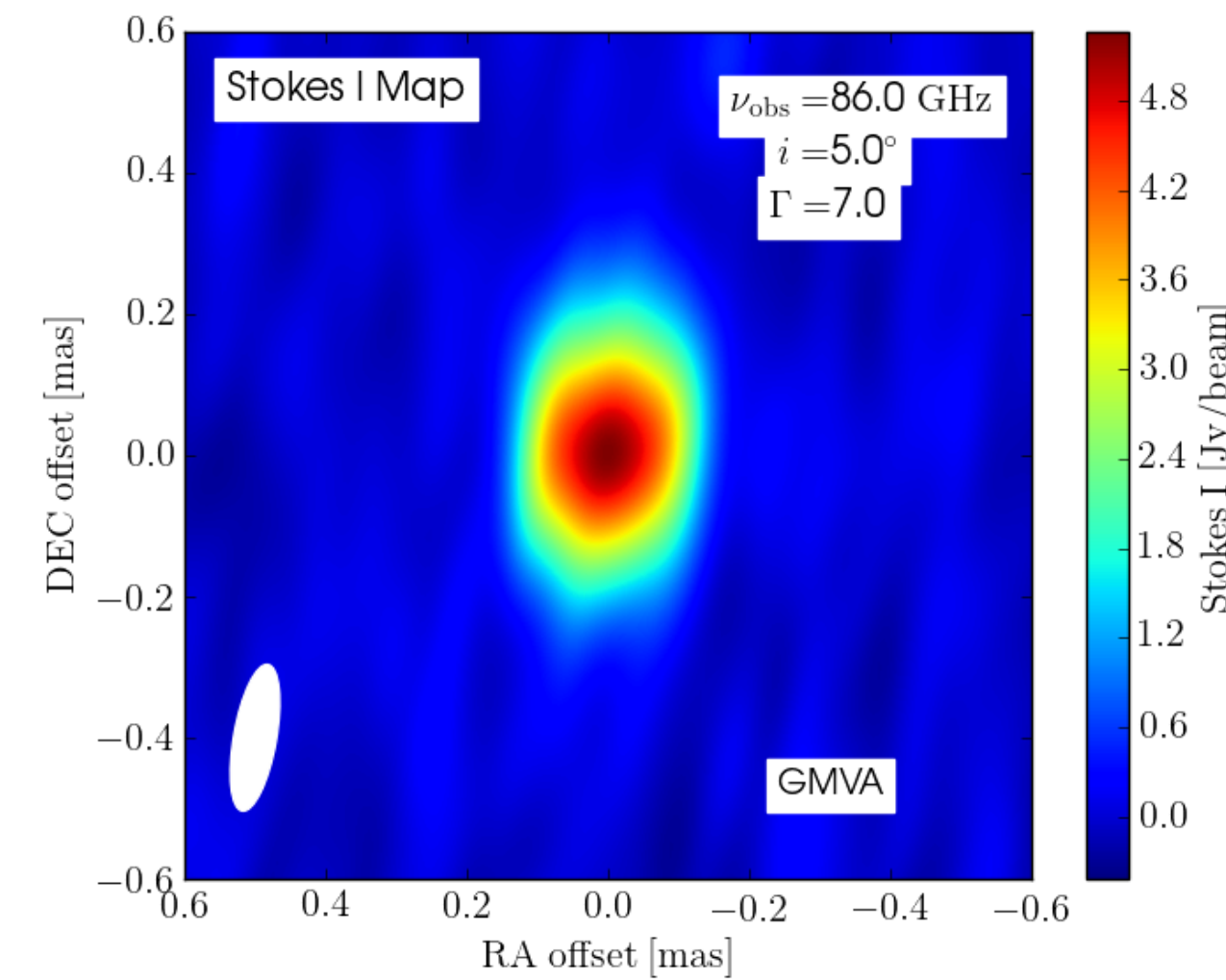
(b)



(c)



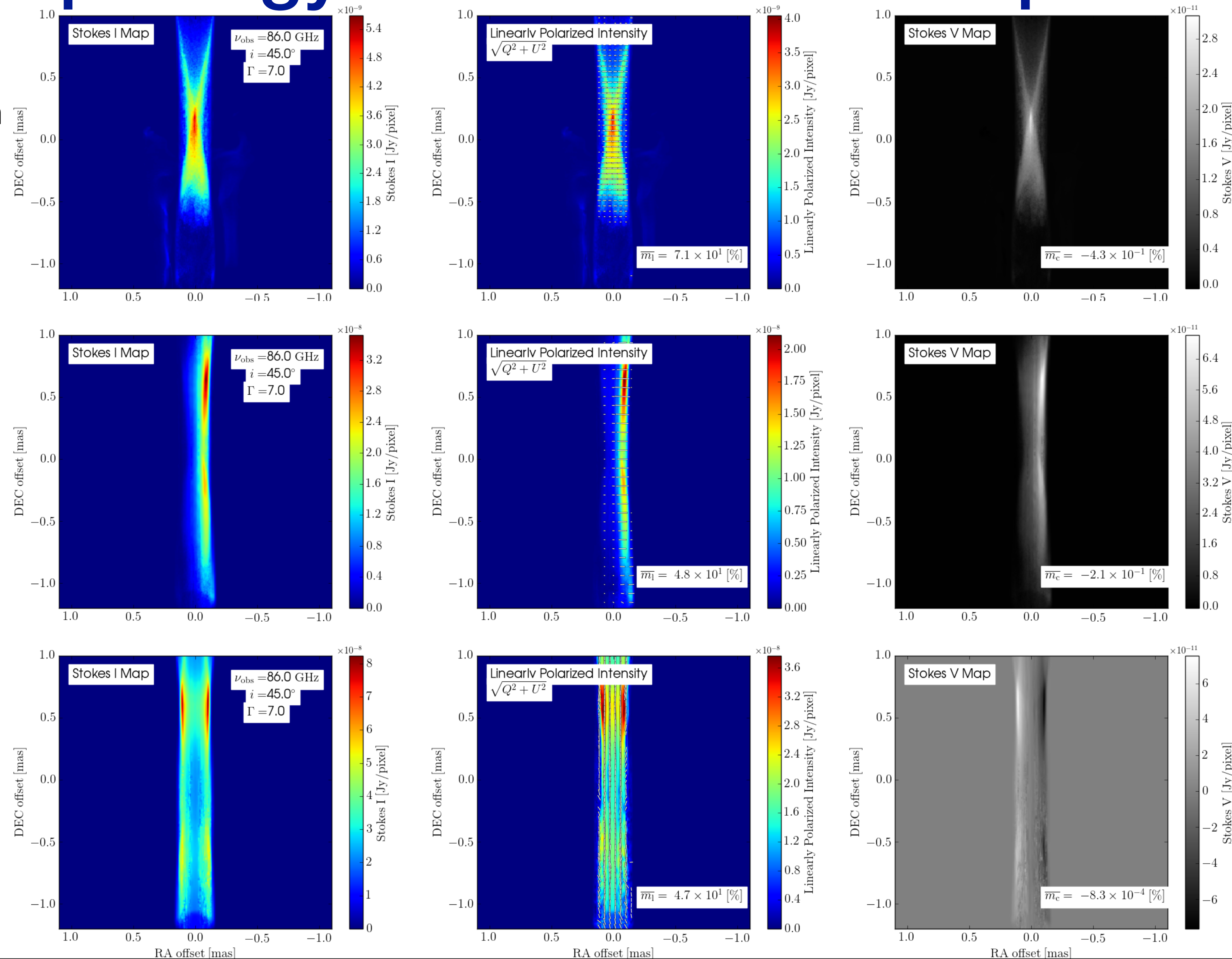
(d)



(e)

Magnetic Field Morphology and Emission Recipe Study

- Resolved linear/circular polarization imaging has the potential to distinguish between a purely poloidal or purely toroidal magnetic field configuration within standing/recollimation shocks.



(i) poloidal

(ii) helical

(iii) toroidal

Stokes I

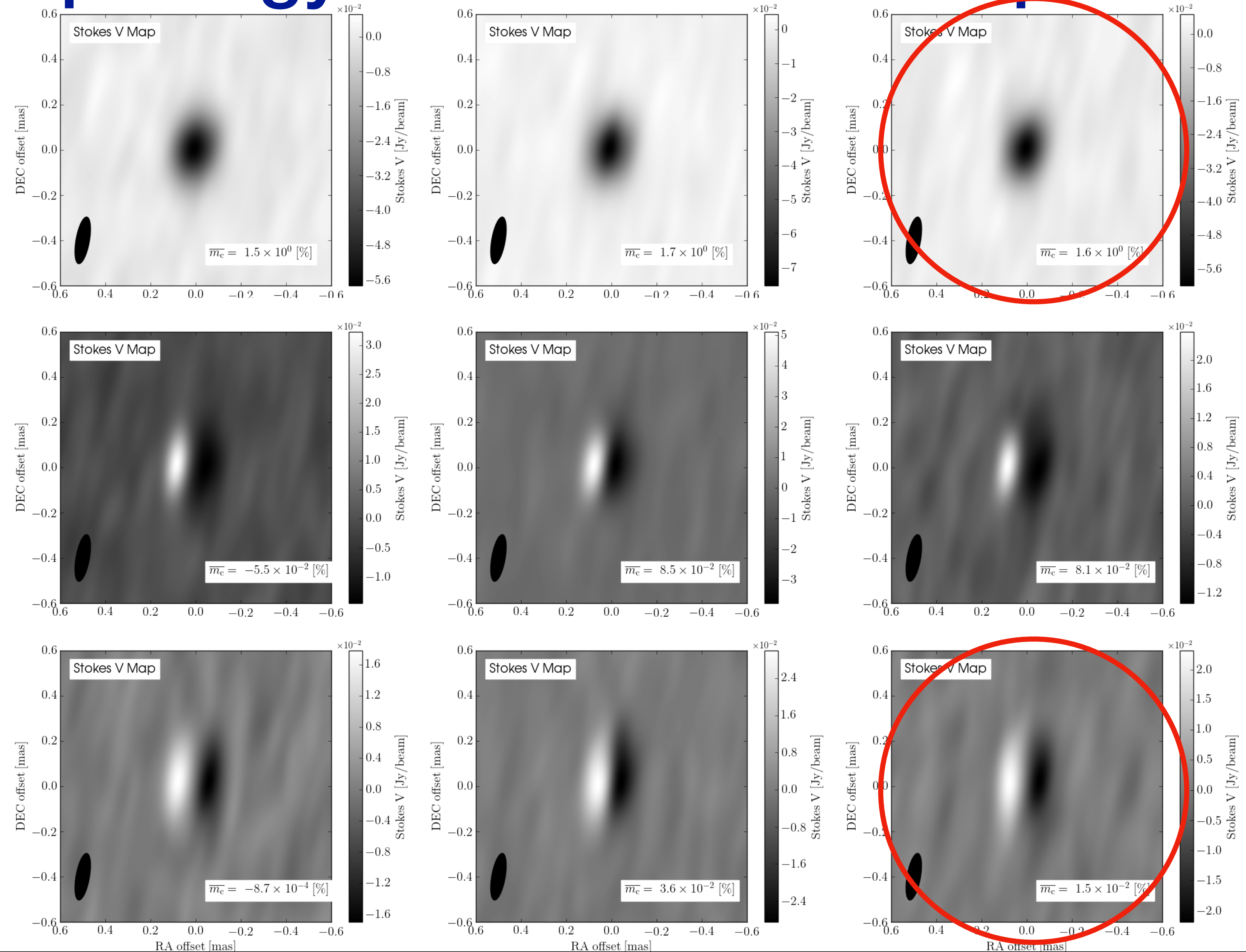
4

Pol. Int.

Stokes V

Magnetic Field Morphology and Emission Recipe Study

- **Poloidal:** highlights the inner structure of the jet (near the recollimation shock)
- **Toroidal:** results in jet **edge-brightening**



(i) poloidal

(ii) helical

(iii) toroidal

$$n_e(\gamma) \propto \rho$$

5

$$n_e(\gamma) \propto p$$

$$n_e(\gamma) \propto B^2$$