

# The LOFAR (and X-ray) view of extragalactic jet populations

#### Judith Croston

Thanks to Beatriz Mingo, Judith Ineson, Brendan Webster, Bonny Barkus + LOFAR surveys team

Image from the data presented in Sabater+ 2021 A&A 648 2

### Internal conditions and environments from X-rays

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Ineson, JC+ 2015 MNRAS 453 2682 & 2017 MNRAS 467 1586

#### Morphology and internal conditions





#### LOFAR extragalactic surveys



DR1 Shimwell+ 2019 A&A 622 1

424 deg<sup>2</sup> 300,000 sources 51% with redshifts



5635 deg<sup>2</sup> 4.4 million sources



#### **Resolved jet structures in LoTSS**





#### **Group/cluster environments**

8475 RL AGN matched with SDSS groups/clusters = 900 jet/cluster associations



0.40 0.35 0.30 Cluster match fraction 0.20 0.15 0.10 0.05 radio luminosity 0.00 -22 23 24 25 26 27  $log_{10}$  (150-MHz luminosity / W Hz<sup>-1</sup> ) 0.40 0.35 0.30 match fraction 0.25 0.20 Cluster 0.15 0.10 -\* size 0.05 0.00 101  $10^{2}$ 10<sup>3</sup> log10 (Size / kpc)

#### **Revisiting the FR paradigm with LOFAR**





### **Revisiting the FR paradigm with LOFAR**



#### Mingo et al. 2019 MNRAS 488 2701



#### How does environment affect low-power jets?





LOFAR deep fields dataset: **Tasse+ 2021** A&A 648 1 **Sabater+ 2021** A&A 648 2 **Kondapally+ 2021** A&A 648 2 **Duncan+ 2021** A&A 648 4 Catalogues and images at www.lofar-surveys.org/deepfields.html

Low-power jets are 2-3x more likely to be an FRI than an FR2.

But probability of FR2 morphology increases at low stellar mass.

unrelated to accretion mode



## Accretion mode and morphology in LOFAR deep fields



	HERG (RE)	LERG (RI)
FRII high	24 (35%)	45 (65%)
FRII low	5 (9%)	52 (91%)
FRII all	29 (24%)	96 (76%)
FRI	7 (4%)	153 (96%)
Small/unresolved RLAGN	1225 (8%)	13,215 (92%)

+ see Webster, JC+ 2021 MNRAS 500 4921 for more on small jets in LOFAR surveys

#### Summary



- Deep, representative sample X-ray studies demonstrate a physical difference in FRI and FRII particle content, best explained by entrainment in FRIs.
- LOFAR surveys reveal a new population of low-luminosity FRIIs
- The evolution of low-power jets (i.e. whether FRI or low-luminosity FRII) linked to stellar mass and not accretion mode.
- Accretion mode is decoupled from jet and lobe dynamics and appearance: luminous jets can be RI or RE irrespective of morphology, while low-luminosity jets are typically RI.
- LOFAR catalogues and images, including 6000 morphologically classed radio galaxies, and 900 jet/cluster associations, at www.lofar-surveys.org

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Hardcastle & Croston 2020 NewRevAst, 88 (arXiv:2003.06137)