

Dynamics and emission model of the recollimation shock in BL Lacertae

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Recollimation shock and moving features

VLBA monitoring at 15 GHz (1995-2012)
in the frame of the MOJAVE program

- Beam size: ~ 0.9 mas
- 1 mas = 1.29 pc; 114 epochs

Radio core:

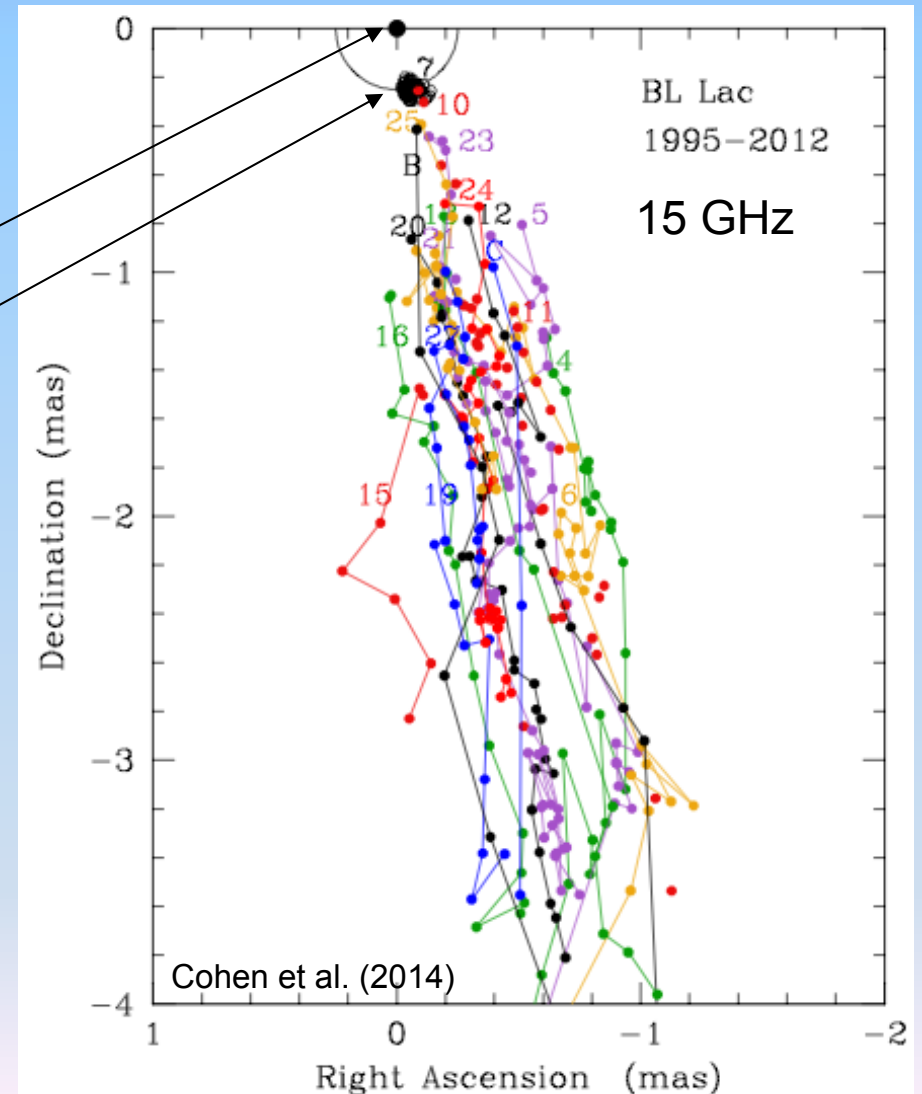
- reference point; flat spectrum,
- compound str. at 43 GHz \rightarrow core shifts at 15 GHz

C7 component:

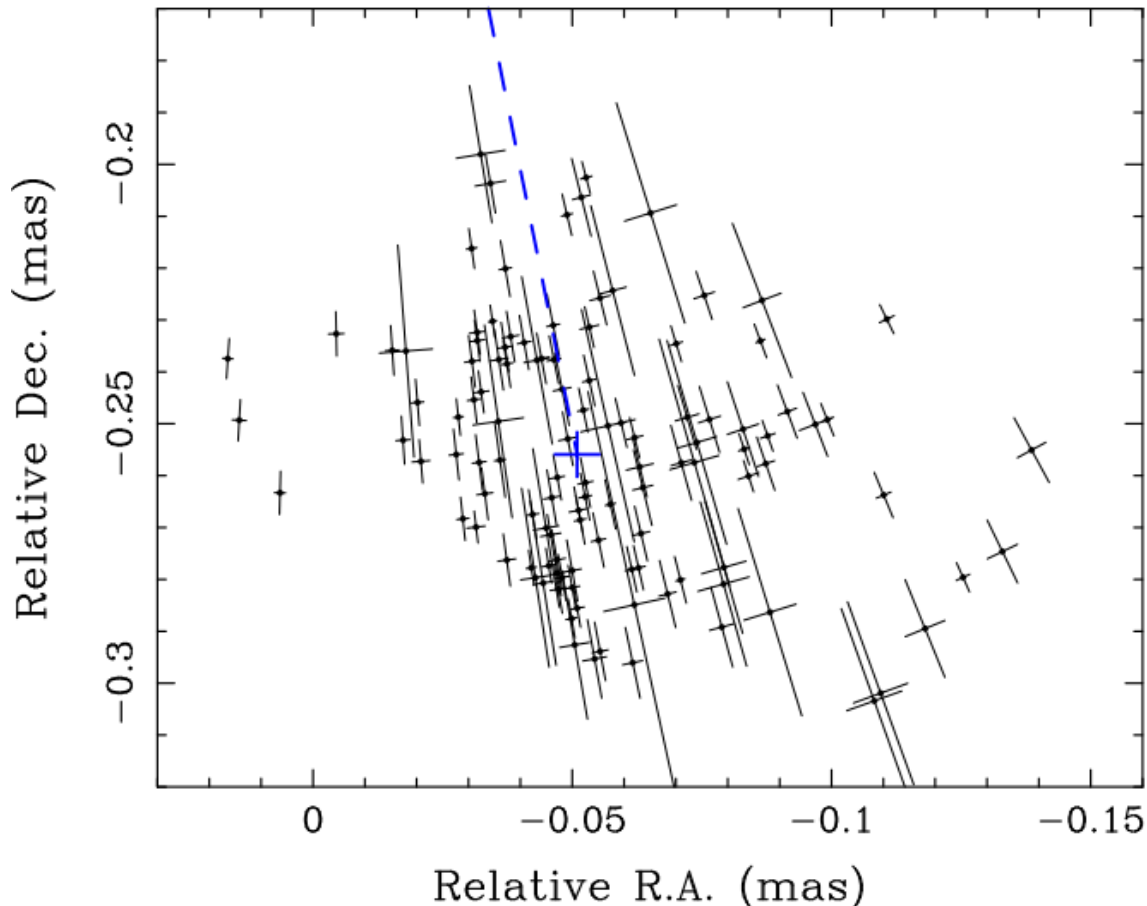
- Quasi-stationary component (recollimation shock)
- tightness of points

19 moving components emanate from C7;
apparent speeds $< 10c$

“Whip” jet model: relativistic, rapidly shaken
RCS generates transverse Alfvén waves
propagating downstream on helical magnetic field
(Cohen et al. 2014, 2015).



RCS scatter



Data: 116 epochs between 1999-2016 (Arshakian et al. 2020)

Data reduction as in Cohen et al. (2014)

Scatter:

- Size ~ 0.1 mas
- Dynamical/geometry reasons.
- Moving of the core as a result of changes of a pressure/density.
- Intrinsic error of RCS.

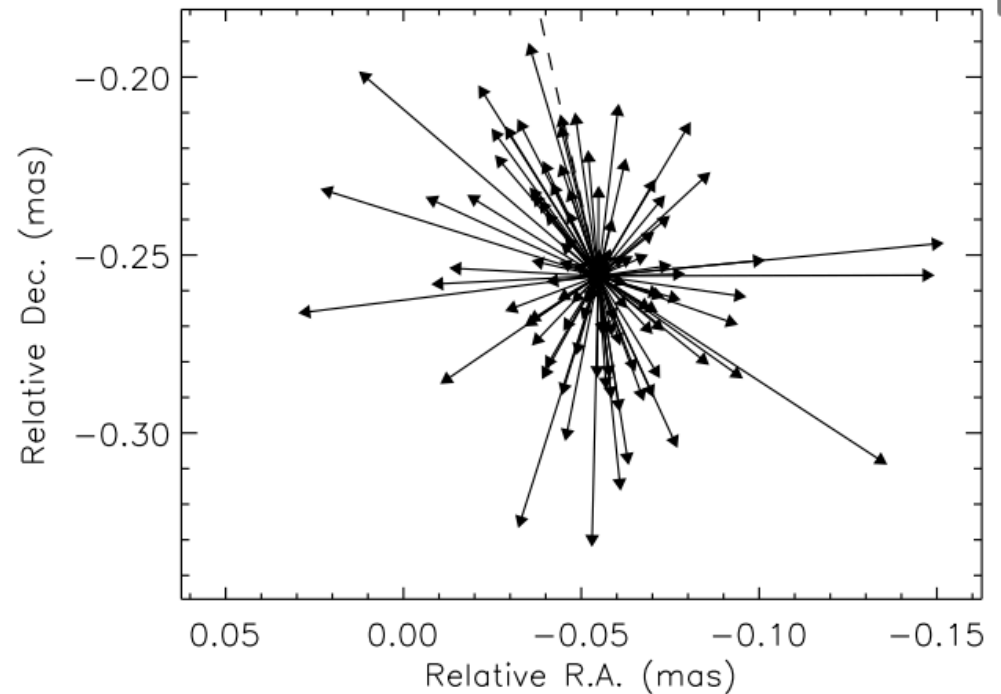
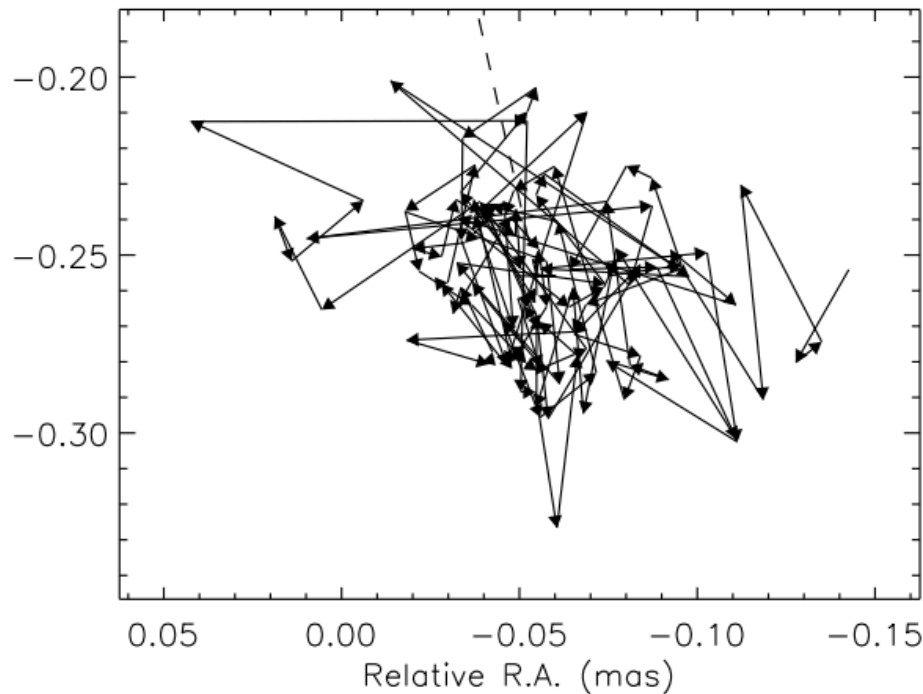
Positional uncertainties:

Median: (4.9×1.6) μ as

Flux leakage between core and

RCS: Typically small (within 10%) but in rare cases can reach to 50%

RCS trajectory



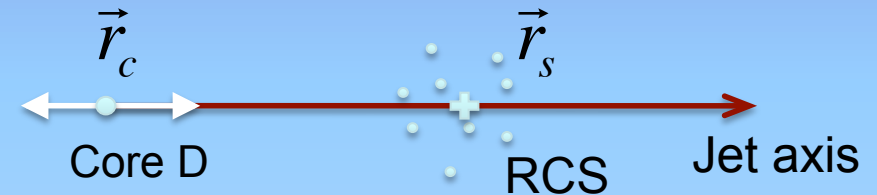
Motion vector – motion of RCS between two consequent epochs.

- **Random orientation** of six long motion vectors (>0.08 mas)
- **Anisotropy**: Motion vectors (<0.08 mas) have a statistically significant preferential orientation along the jet axis
- **Asymmetry**: Length of motion vectors are larger along the jet axis
- Anisotropy and asymmetry - preferential motion of the core along the jet axis

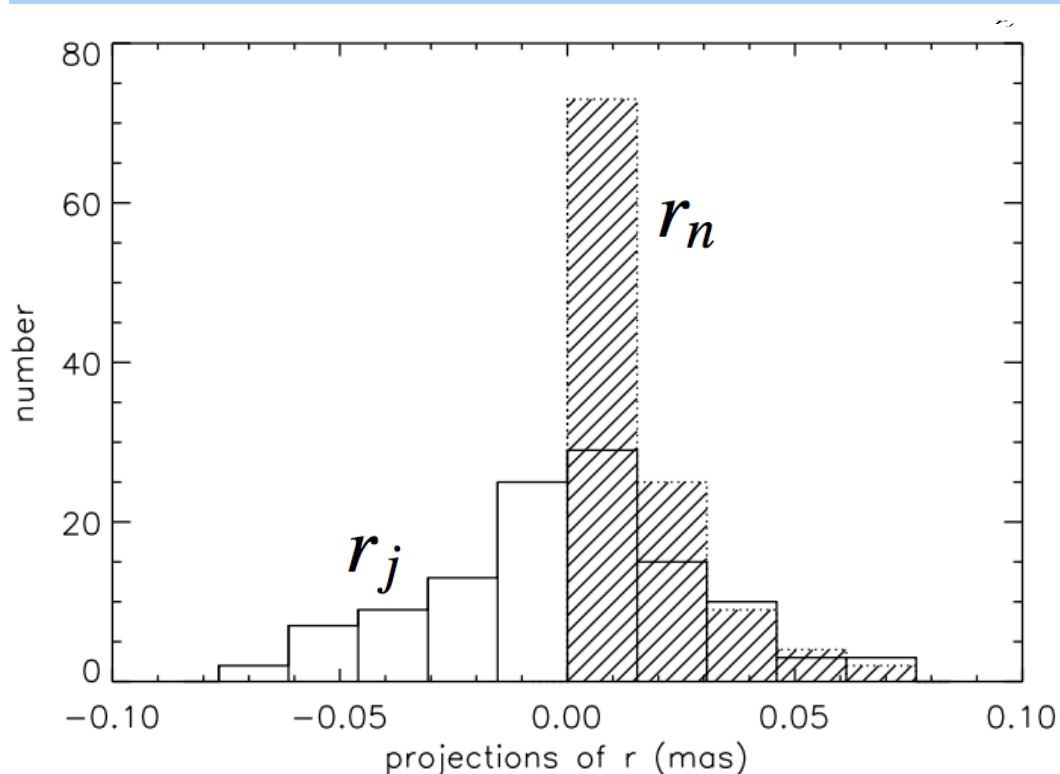
Estimates of core shift and intrinsic motion of RCS

Apparent motion-vector of RCS:

$$\vec{r} = -\vec{r}_c + \vec{r}_s$$



- r_c : The core position shifts in the direction of the jet
- r_s : The RCS moves in random directions within a sphere



$$\text{rms}_c = \sigma_c = \left(\sigma_{r_j}^2 - \sigma_{r_n}^2 \right)^{\frac{1}{2}}$$

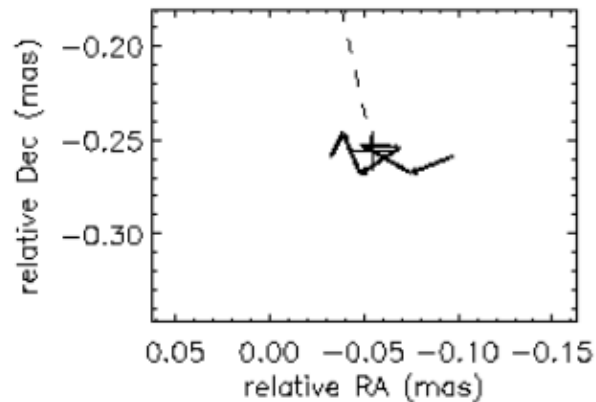
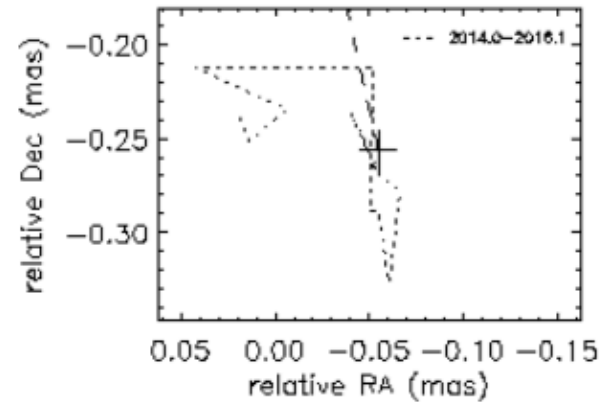
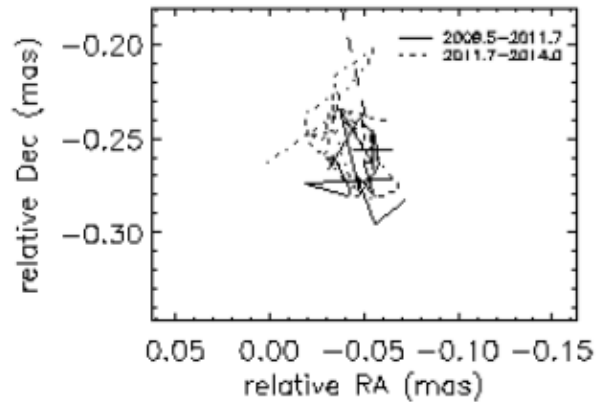
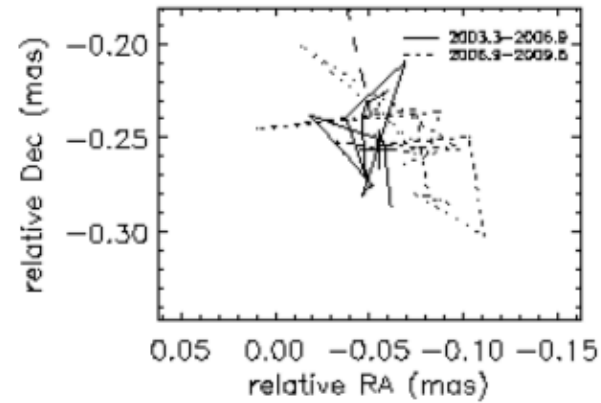
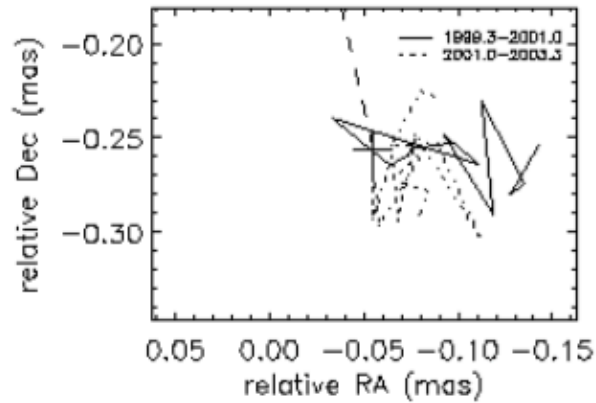
$$\text{rms}_s = \left(s^2 \right)^{\frac{1}{2}} = \left(2\overline{r_n^2} \right)^{\frac{1}{2}}$$

$$\text{rms}_c = 0.025 \pm 0.008 \text{ mas}$$

$$\text{rms}_s = 0.025 \pm 0.009 \text{ mas}$$

Contribution of core shift to the apparent motion of RCS is significant.

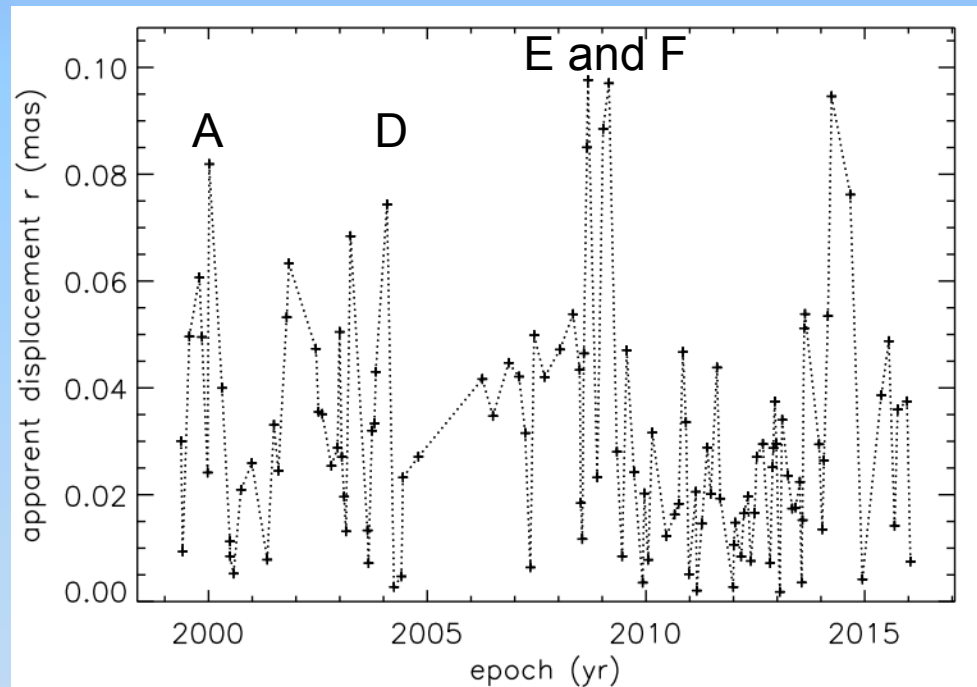
RCS motion



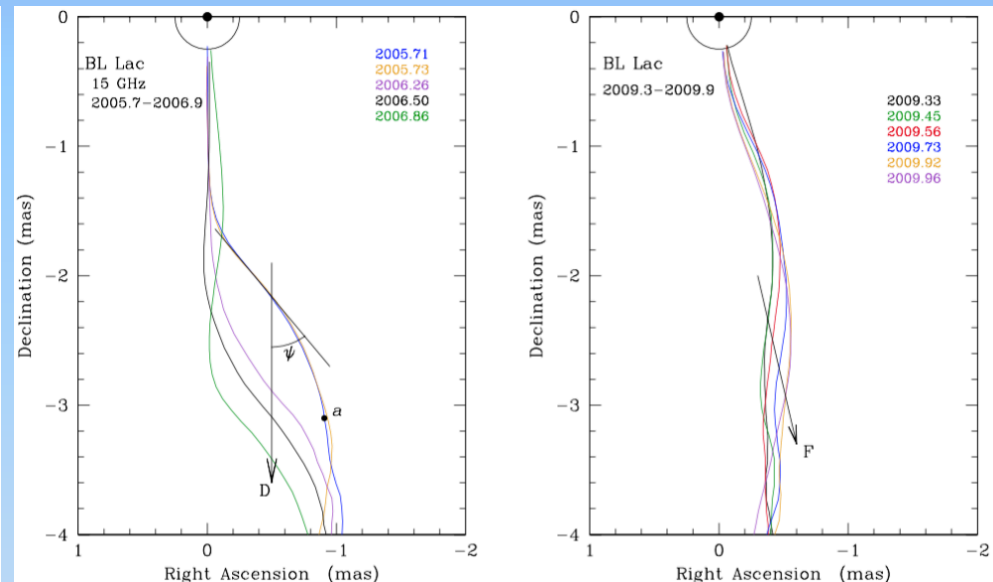
Time scales of weeks-months: erratic motion

Time scales of years: clock-wise motion (chance prob. 3%)?; mean speed - $0.1c$

Link between the motion of RCS and wave excitation

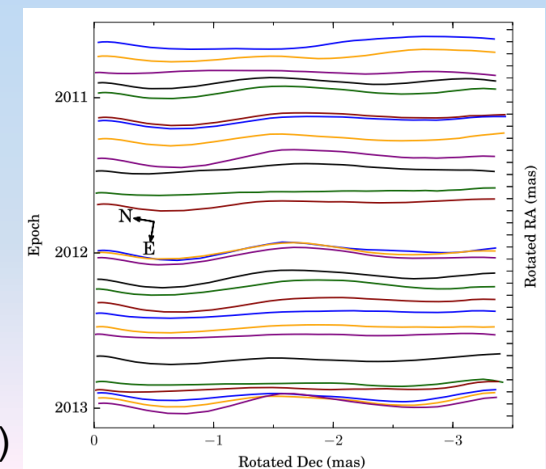


- Relatively large shakings of RCS (≥ 0.07 mas) generates transverse waves A, D, E, and F in the jet.
- Jet stable state 2010-2013: The RCS acts as the nozzle of the jet and generates quasi-sinusoidal waves with amplitudes lower than ≈ 0.02 mas

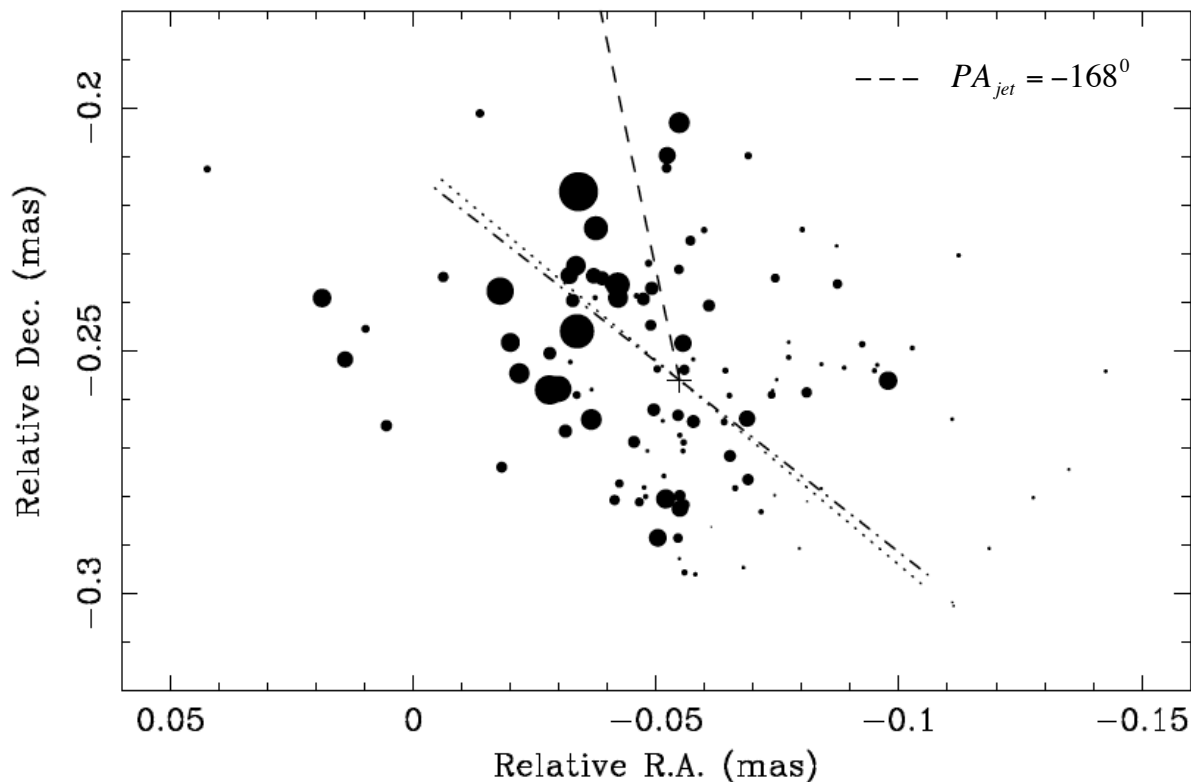


Waves D and F (Cohen et al. 2015)

Quasi-sinusoidal waves
(Cohen et al. 2015)



Distribution of flux density of RCS on sky



- Flux density range: 0.17-4.4 Jy
- Flux distribution is **asymmetric** along and transverse to the jet central axis ($PA_{jet} = -168^{\circ}$, dashed line).
- Statistically significant flux asymmetry with respect to the jet axis: $\alpha_{sym} = 46$ deg.

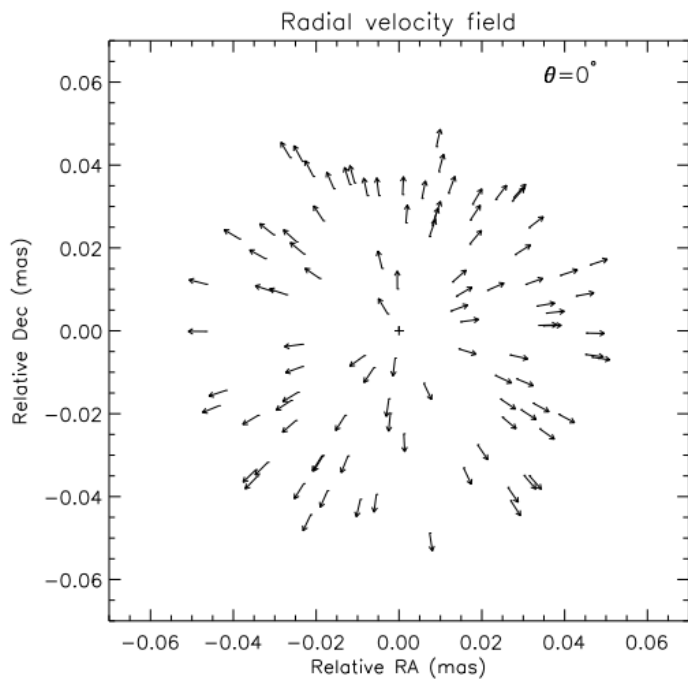
Simulations of RCS emission

Toy model:

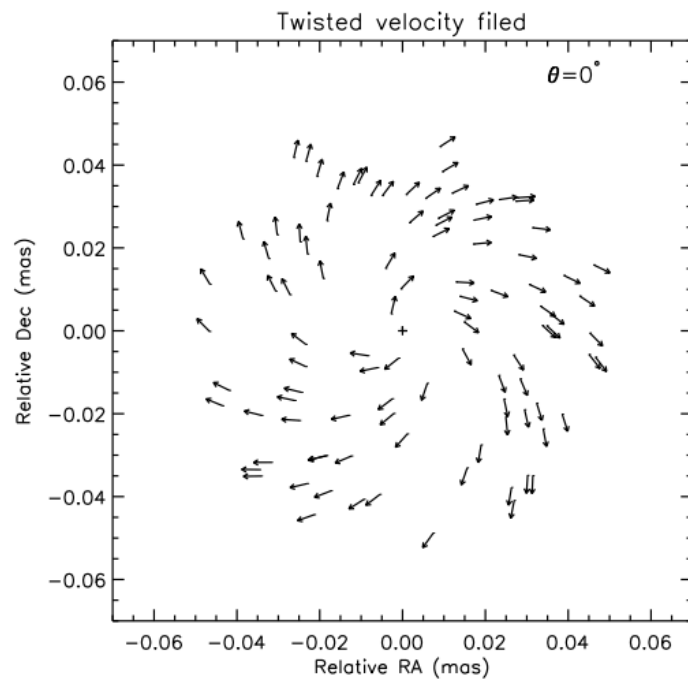
- RCS moves in a plane normal to the jet axis
- RCS drags the jet outflow in a swinging motion: $\theta_{\text{RCS}}(r, \varphi)$ is the viewing angle of the jet outflow at the position of RCS
- S_0 , β are constant
- Beaming is due to variation of θ_{RCS}

Viewing angle of the jet central axis $\theta = 0^\circ$

Radial velocity field



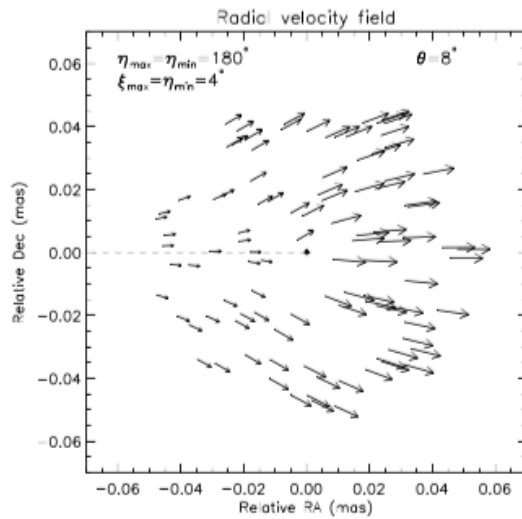
Twisted velocity field



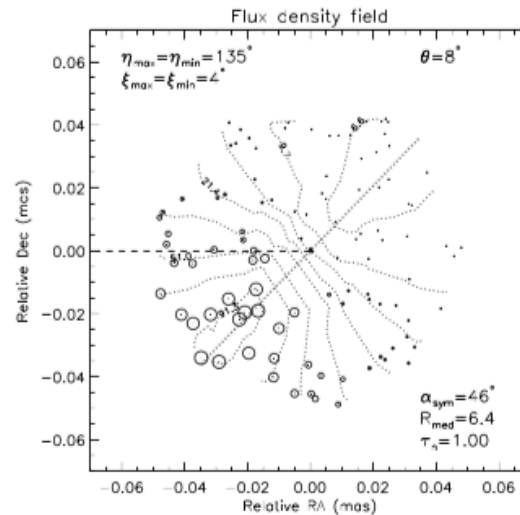
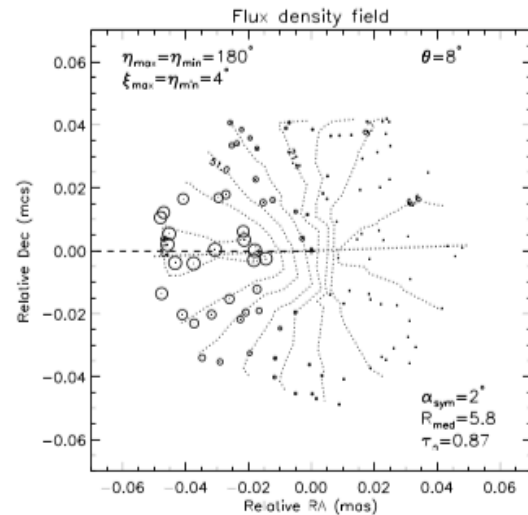
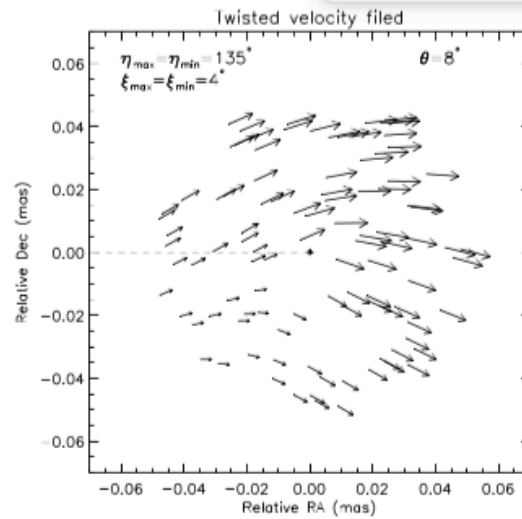
Simulations of RCS emission

Viewing angle of the jet central axis $\theta = 8^\circ$

Radial velocity field



Twisted velocity field



$$S(r, \varphi) = S_0 D(\beta, \theta_{RCS}(r, \varphi))^{2-\alpha}$$

Flux density asymmetry is due to twisted velocity field of the jet at the position of RCS.

Summary

- Vector motions of RCS are asymmetric along the jet axis – evidence of resolution-dependent core shift.
- We developed a statistical tool for estimating the core shift and intrinsic motion of RCS: Projected core motion is comparable to intrinsic motion of RCS
- RCS moves with sub-relativistic speed of about $0.1c$ on time scales of few years
- RCS acts as the nozzle of the jet and its motion generates transverse waves of various amplitudes traveling down the jet (“Whip” model)
- On-sky flux density distribution is asymmetric along and transverse to the jet axis
- Simple model of RCS having the twisted velocity field can account for the observed flux density asymmetry