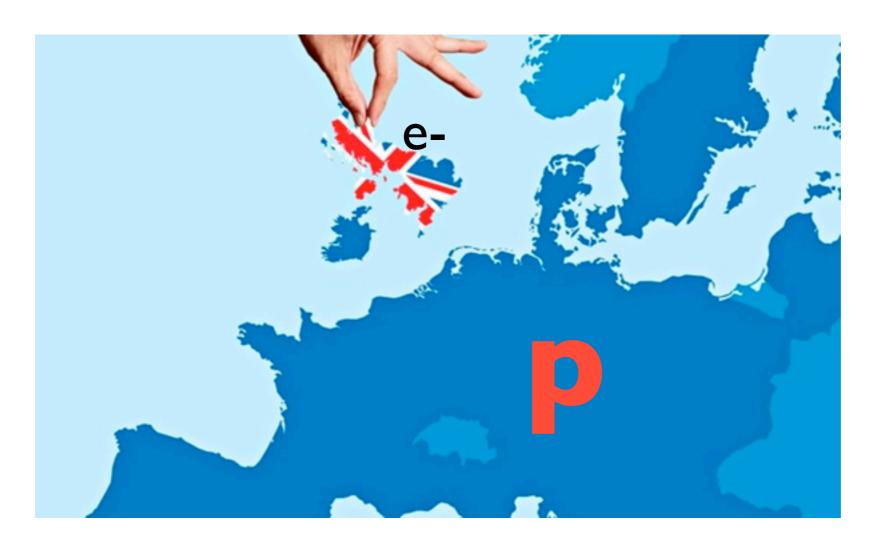


## Reionization



When?





Source?



# Looking forward to a good weekend



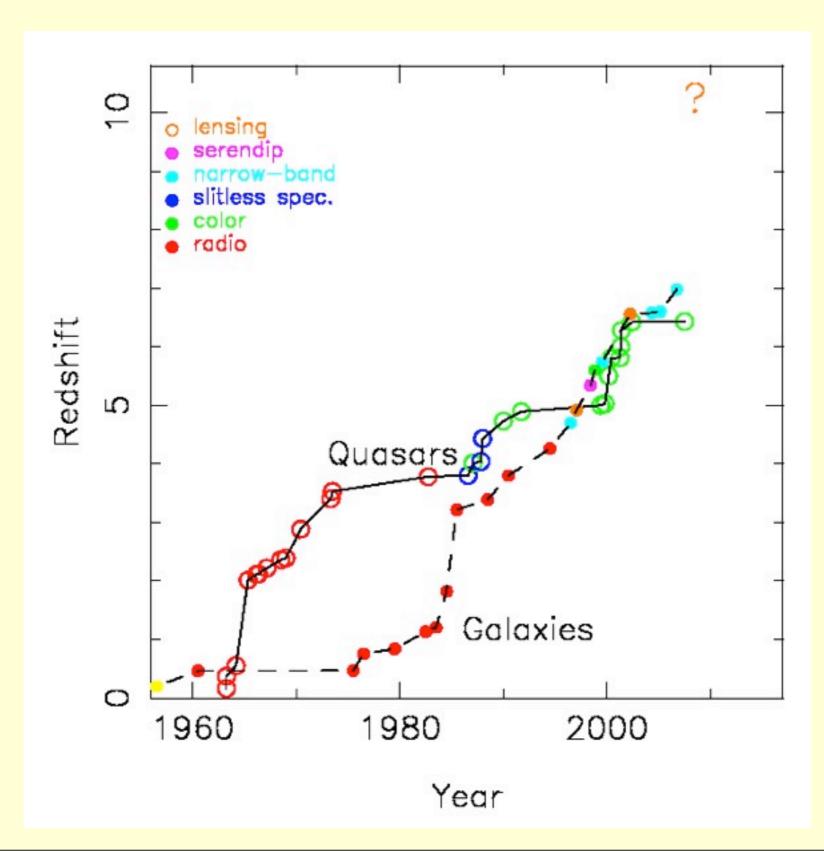
# Thank you!

## **Local Organising Committee**

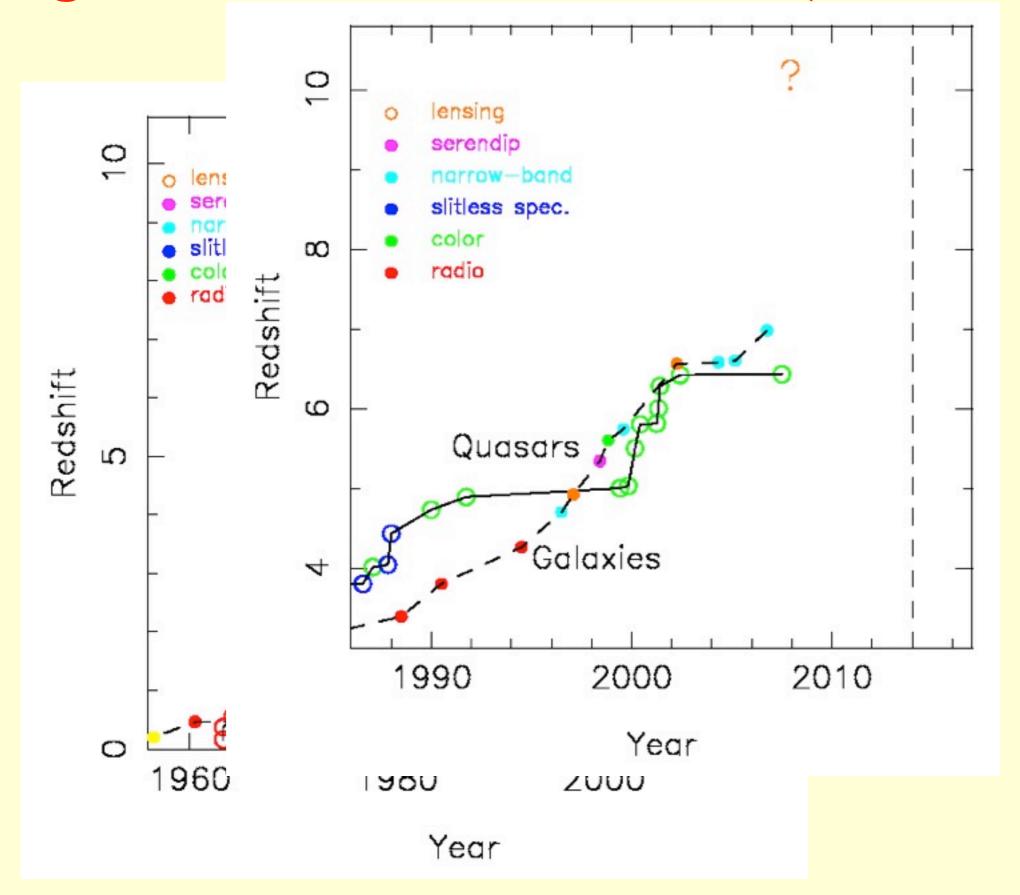
- Frederick Davies (MPIA)
- Christina Eilers (MPIA)
- Emanuele Farina (MPIA)
- Carola Jordan (MPIA)
- Chiara Mazzucchelli (MPIA)
- Bram Venemans (MPIA)

- for I40-character soundbite of each talk:
   #DarkAgesHD on twitter
- reflection of what have changed in the last decade in our understanding of reionzation
- results of yesterday's survey
- discussion

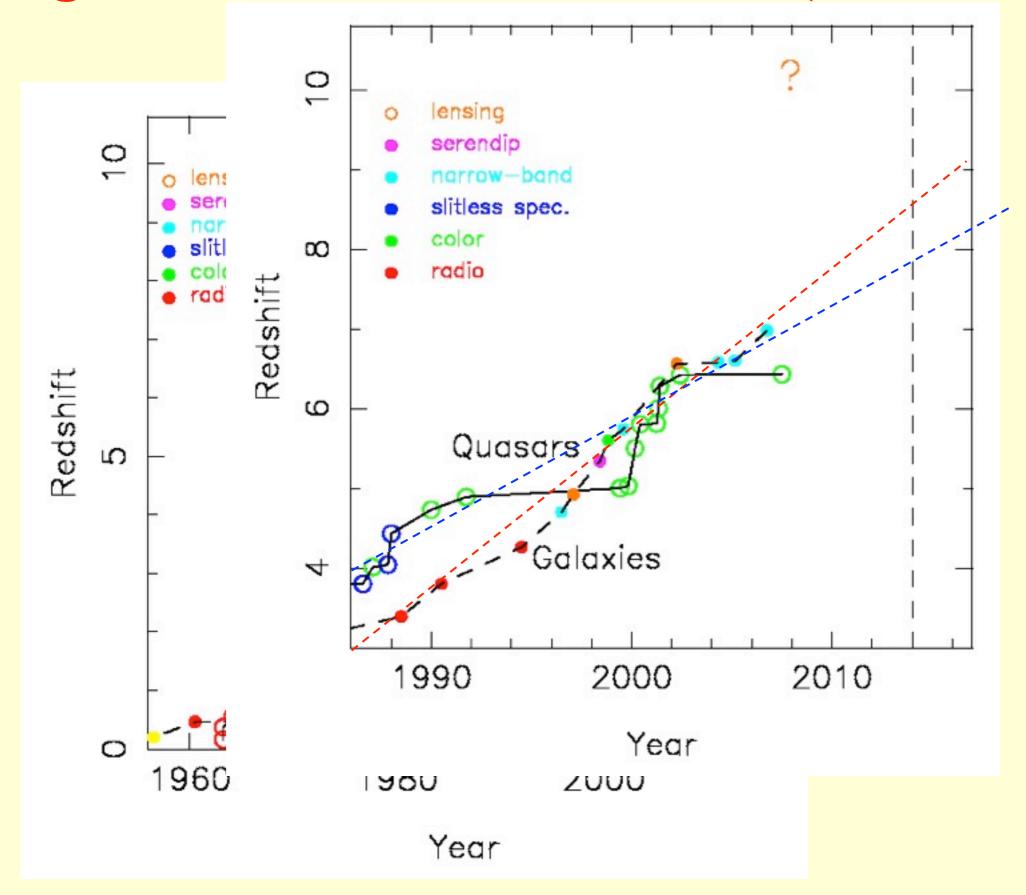
# **Highest Redshift Prediction (XF 2007)**



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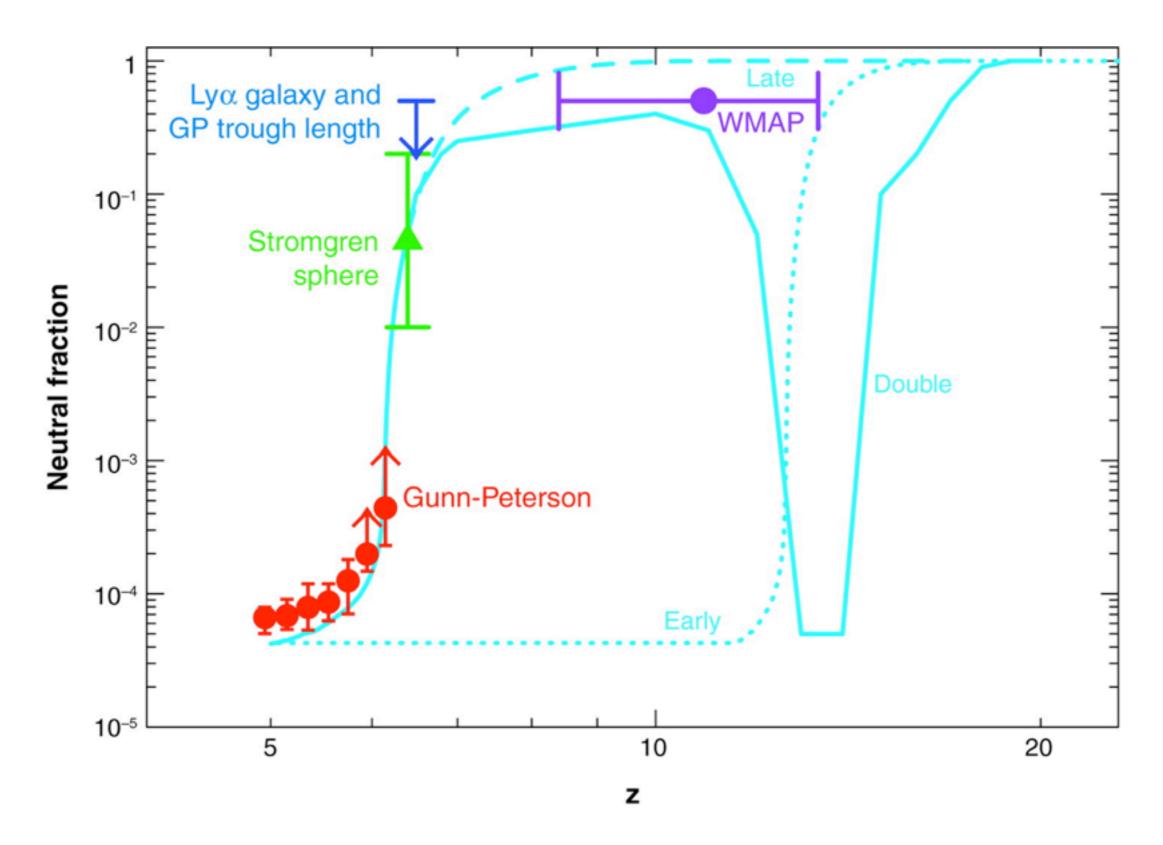


#### From Fan, Carilli and Keating (2006) ARAA:

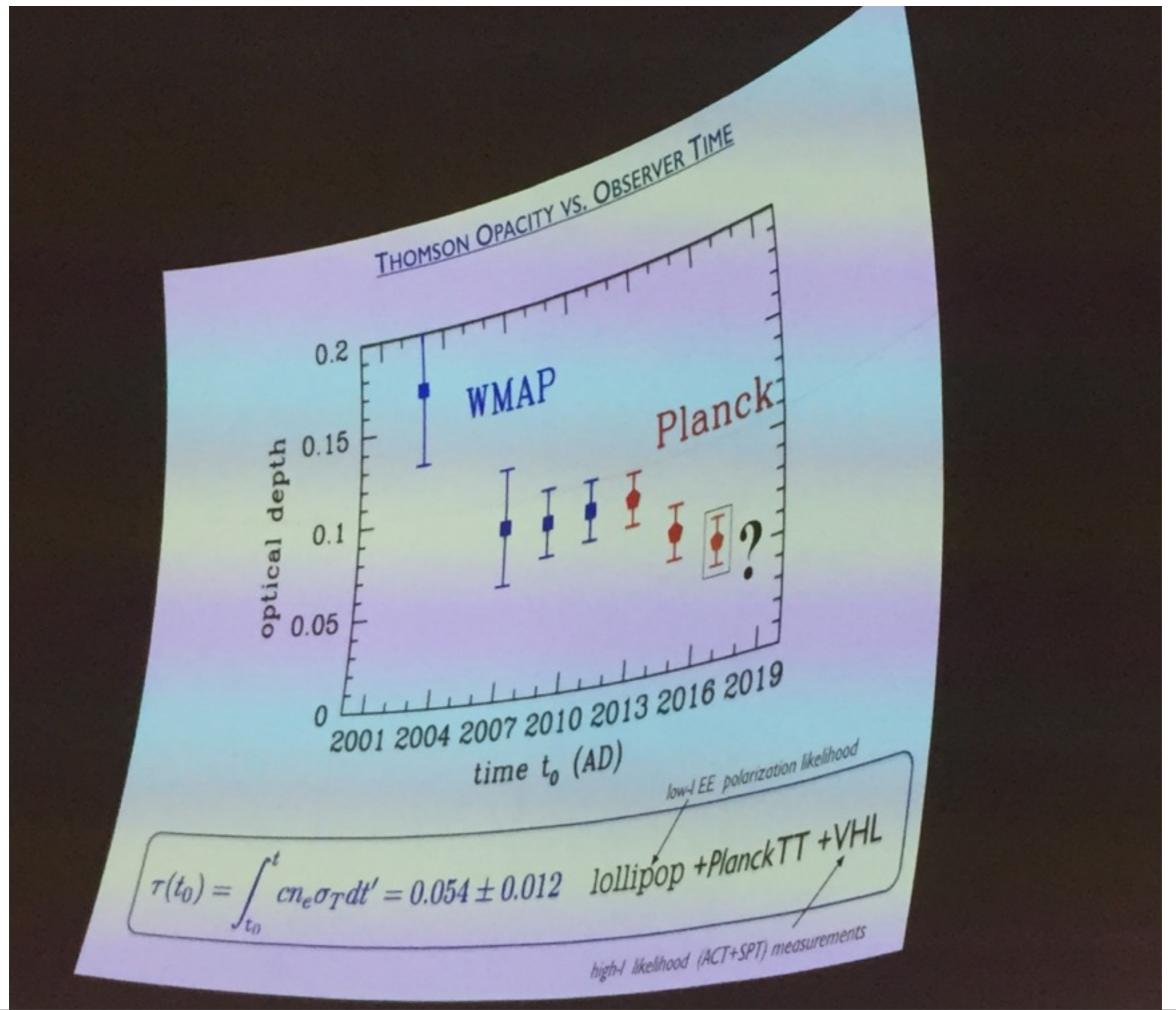
- Observations have set the first constraints on the epoch of reionization (EoR), corresponding to the formation epoch of the first luminous objects.
  - Studies of Gunn-Peterson (GP) absorption indicate a rapid increase in the neutral fraction of the intergalactic medium (IGM) from  $x_{HI}$  < 10<sup>-4</sup> at z ≤ 5.5, to  $x_{HI}$  > 10<sup>-3</sup>, perhaps up to 0.1, at z~6,
  - while the large scale polarization of the cosmic microwave background (CMB) implies a significant ionization fraction extending to higher redshifts,  $z_{\sim}11 \pm 3$ .
  - These results, as well as observations of galaxy populations, suggest that reionization is a process that begins as early as  $z_14$ , and ends with the "percolation" phase at  $z_6$  to 8.
  - Low luminosity star-forming galaxies are likely the dominant sources of reionizing photons.
  - Future low-frequency radio telescopes will make direct measurements of HI 21-cm emission from the neutral IGM during the EoR, and measurements of secondary CMB temperature anisotropy will provide details of the dynamics of the reionized IGM.

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  - These results, as well as observations of galaxy populations, suggest that reionization is a process that begins as early as  $z_14$ , and ends with the "percolation" phase at  $z_6$  to 8. simpler than we thought?
  - Low luminosity star-forming galaxies are likely the dominant sources of reionizing photons. AGN are back!
  - Future low-frequency radio telescopes will make direct measurements of HI 21-cm emission from the neutral IGM during the EoR, and measurements of secondary CMB temperature anisotropy will provide details of the dynamics of the reionized IGM. still waiting...



Ran X, et al. 2006.
Annu. Rev. Astron. Astrophys. 44:415–62

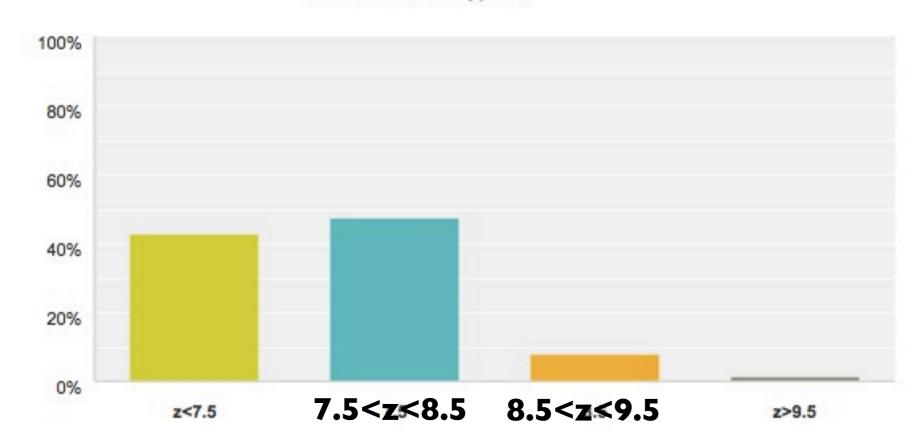


# The reionization vote

- 63/87 participants voted
- 72% voter turnout (same as Brexit)

# When did the universe become 50% ionized?

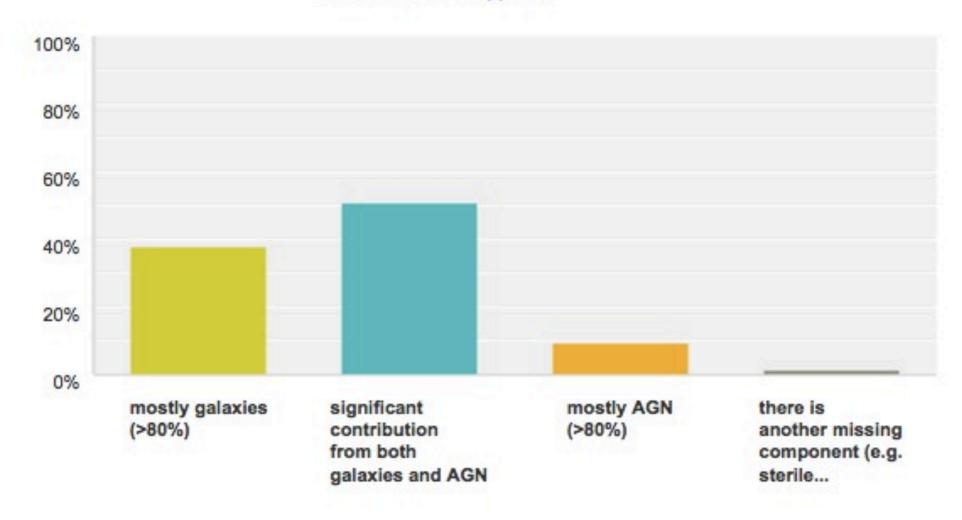
Answered: 63 Skipped: 0



Answer Choices	Responses	*
z<7.5	42.86%	27
7.5	47.62%	30
8.5	7.94%	5
z>9.5	1.59%	1
otal		63

#### What are the main sources of reionization?

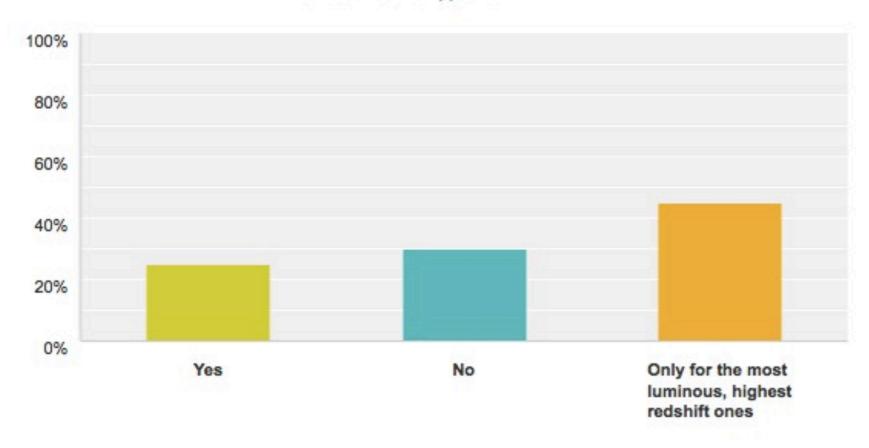
Answered: 63 Skipped: 0



Answer Choices		Responses	
	mostly galaxies (>80%)	38.10%	24
	significant contribution from both galaxies and AGN	50.79%	32
	mostly AGN (>80%)	9.52%	6
	there is another missing component (e.g. sterile neutrinos)	1.59%	1
ota			63

#### Are DCBHs high-z SMBH seeds?

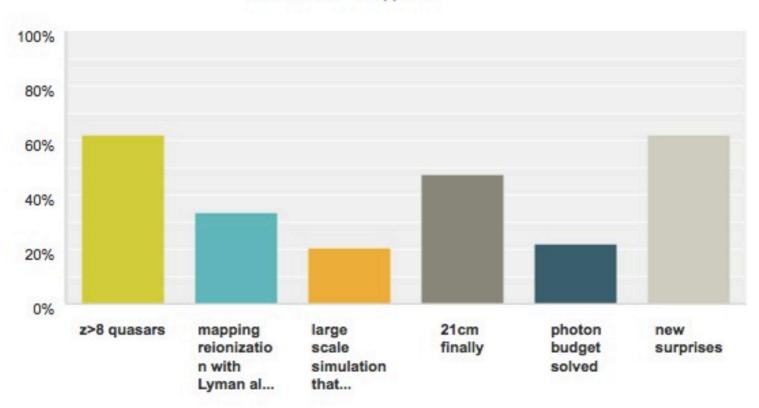
Answered: 60 Skipped: 3



Ans	wer Choices	· ×	Responses	7
9"	Yes		25.00%	15
	No		30.00%	18
283	Only for the most luminous, highest redshift ones		45.00%	27
Tota	al		<u> </u>	60

# What area will likely have exciting breakthrough in understanding EoR in the next decade? (multiple choices ok)

Answered: 63 Skipped: 0



Ans	swer Choices	Responses	1
¥	z>8 quasars	61.90%	39
	mapping reionization with Lyman alpha emitter topology	33.33%	21
~	large scale simulation that captures all relevant physics	20.63%	13
~	21cm finally	47.62%	30
•	photon budget solved	22.22%	14
~	new surprises	61.90%	39
Tota	al Respondents: 63		

## Things that I heard

- Fabian Walter: Why haven't we find z~8 quasars?
- Masami Ouchi: no serious problem of ionizing photons anymore; now worry about too many ionizing photons
- **Piero Madau:** We have been misled by CMB polarization results on reionization in the last decade
- Tiziana di Mateo: the end of dark ages is bright.