## Detection of thermal SiO emission from a Massive Dense Cold Core

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## The massive dense cold core G333.125-0.562

The core is located in the G333 giant molecular cloud complex, at a distance of 3.6 kpc. It

- is undetectable at wavelengths shorter than 70 μm
- has compact 1.2 mm dust continuum emission (Garay et al., 2004 ApJ)
- is cold, massive and dense
- has 6.6 GHz methanol maser (Ellingsen et al., 1996 MNRAS; Ellingsen, 2005 MNRAS)

## Why is it special?

From the UNSW multi-molecular line survey of the G333 GMC-complex with the Mopra Telescope (see "The Mopra DQS Survey" poster - Cunningham *et al.*), we have detected thermal SiO (a shock tracer) emission in this cold core (Figure 1), along with a range of other molecules (Figure 2). The **wider line** of the SiO (Figure 1 inset) suggests there is an embedded object driving an outflow.



on the integrated emission of CS (grey scale) towards the warm starforming core IRAS16172-5028 and the cold core G333.125-0.562. Note the prominence of SiO cf CS in the cold core. The cross marks the 6.7-GHz methanol maser. (Lo *et al.*, 2007 MNRAS in press)

Figure 1) Integrated emission of SiO (red contours) overlaid



6<sup>h</sup>21<sup>m</sup>45<sup>s</sup> 21<sup>m</sup>30<sup>s</sup> 21<sup>m</sup>15<sup>s</sup> 21 Right Ascension (J2000)

Some derived physical parameters: from NH<sub>3</sub>  $T_{kin} = 13 \text{ K}$ from the SED  $T_{dust} = 18 \text{ K}$  $M_{gas} = 1.8 \times 10^3 \text{ M}_{\odot}$  $L_{bol} = 8 \times 10^3 \text{ L}_{\odot}$  $n_{H_2} = 1 \times 10^5 \text{ cm}^{-3}$  $N_{H_2} = 4 \times 10^{23} \text{ cm}^{-2}$ 

## **Conclusion**

The cold massive core habours a deeply embedded massive protostellar object, that is driving an outflow. This is occurring at a very early stage of star formation, prior to the creation of an infrared source in the core.

