

Young Quasar Jets Revealed by Dynamic Radio Surveys



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Extragalactic Jets – June 14-18, 2021

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CrossMark

Quasars That Have Transitioned from Radio-quiet to Radio-loud on Decadal Timescales Revealed by VLASS and FIRST

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Kunal Mooley¹⁵ , Dipanjan Mukherjee¹⁶ , Wendy Peters¹⁰, Leonid Petrov¹⁷, Emil Polisensky¹⁰, Wiphu Rujopakarn^{18,19} ,
Mark Whittle⁴, and Mattia Vaccari^{20,21} 

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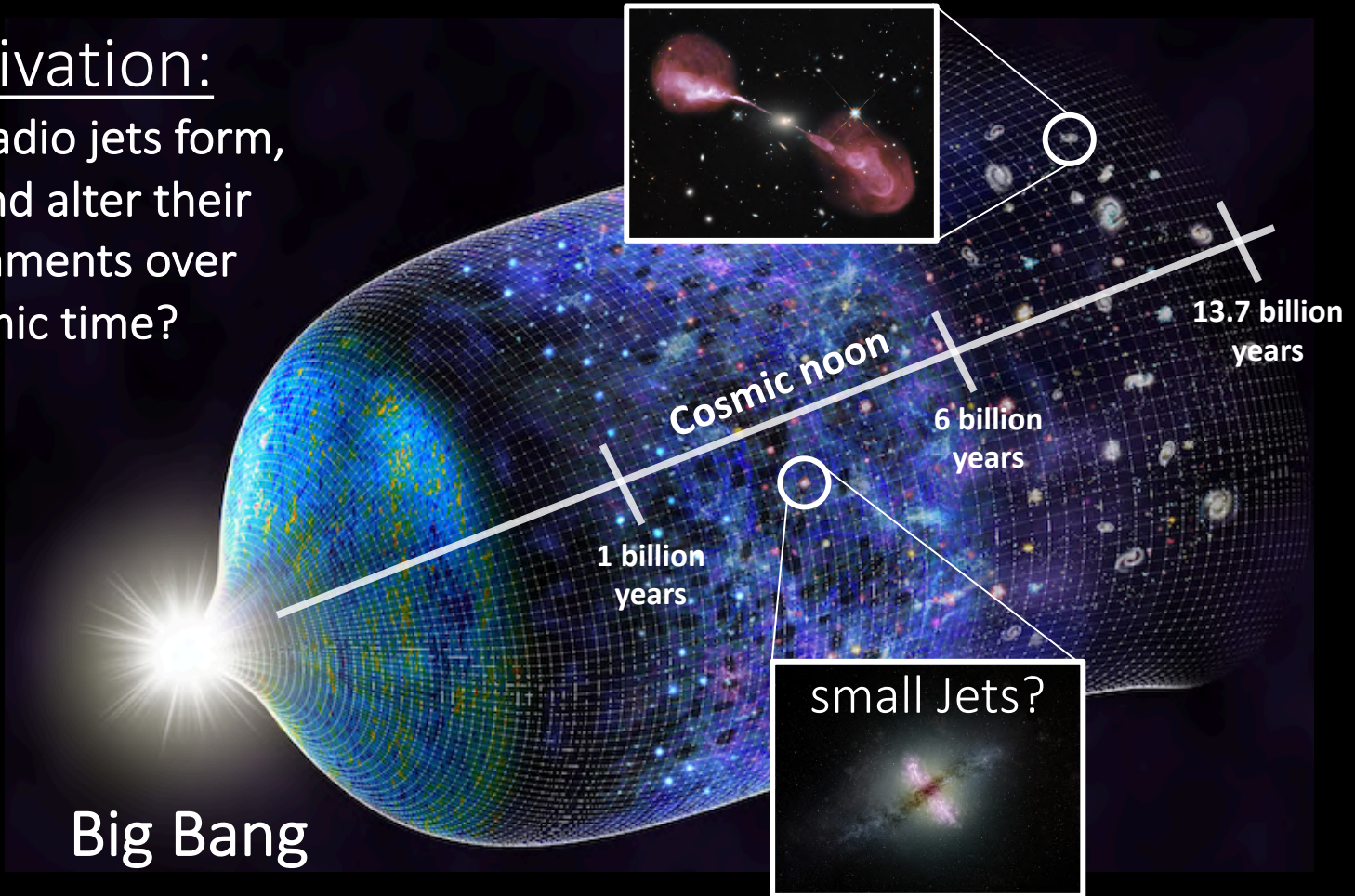
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Extragalactic Jets – June 14-18, 2021

Motivation:

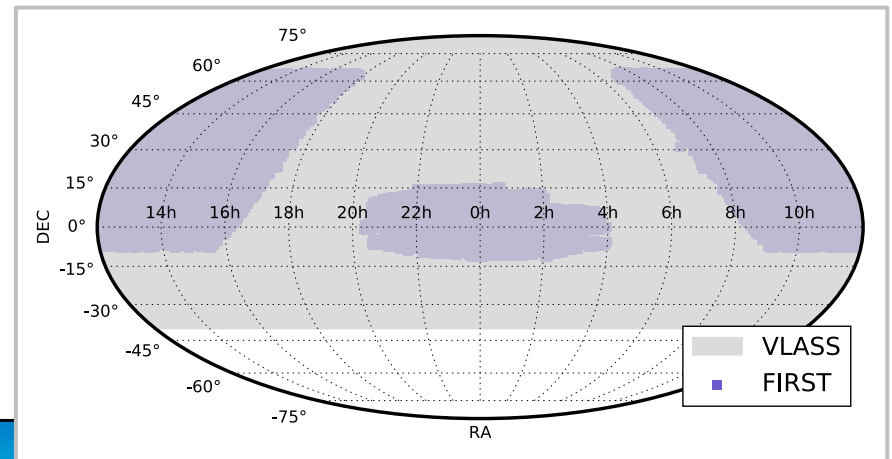
How do radio jets form, grow, and alter their environments over cosmic time?



Catching Young Jets with Multi-epoch Radio Data

	ν (GHz)	θ_{FWHM}	Dates	Epochs
FIRST	1.5	5"	1993-2011	1
VLASS	2-4	2.5"	2017-2019 [†]	3

[†]Epoch 1



References:

Becker et al. 1995
Helfand et al. 2015
Lacy et al. 2020



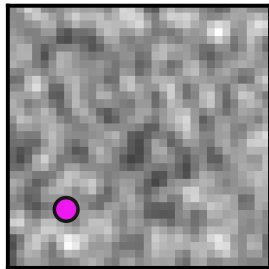
Selection Criteria

1

Initial Sample:

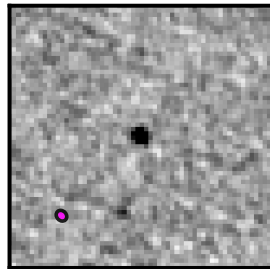
2000 variable sources
over 3440 deg² of VLASS

FIRST



1999

VLASS

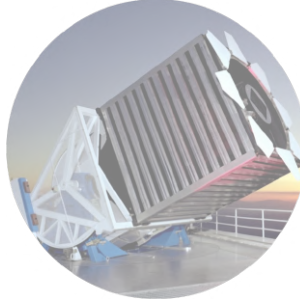


2019

2

Quasar Cross-matching:

SDSS



WISE

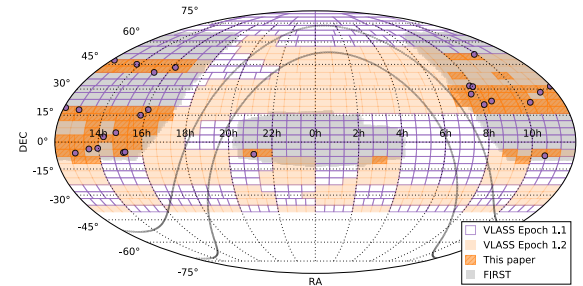


Nyland et al. 2020

3

Final sample:

26 newly radio-loud
quasars ($S_{VLASS} > 3$ mJy)

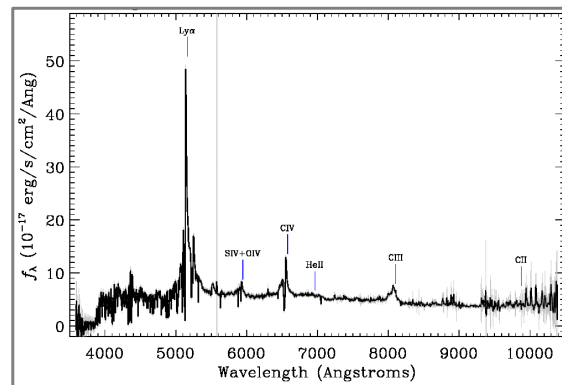
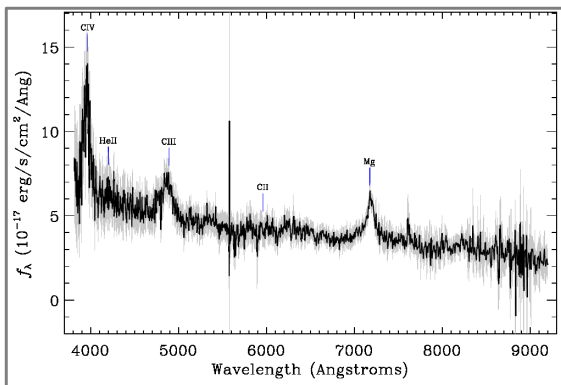
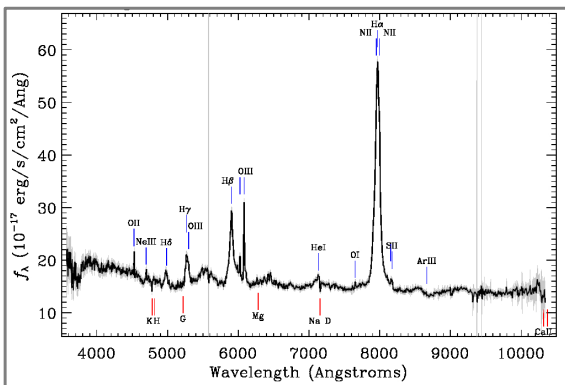
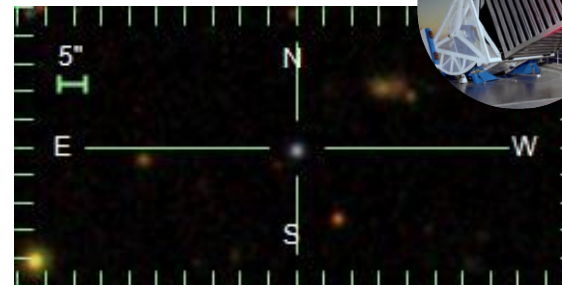
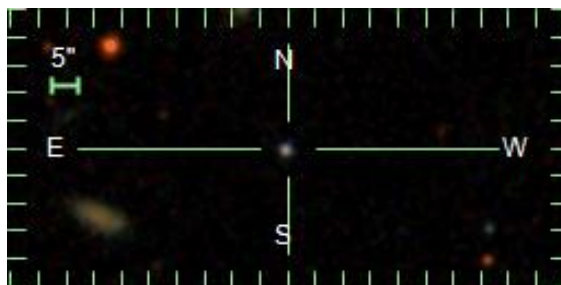
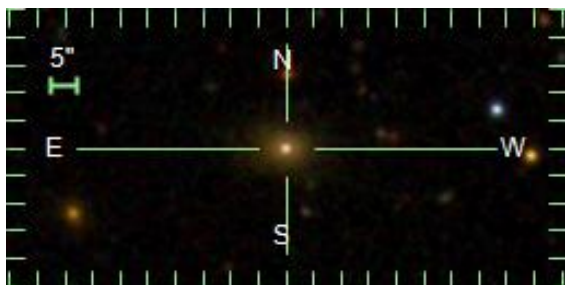
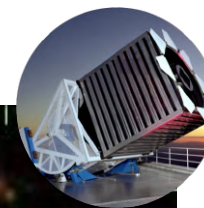


Example SDSS Images and Spectra

$z = 0.2$

$z = 1.6$

$z = 3.2$



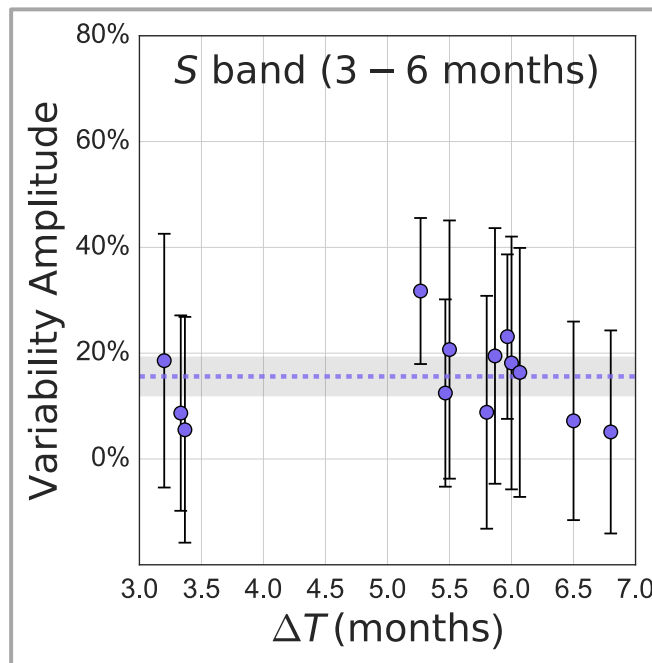
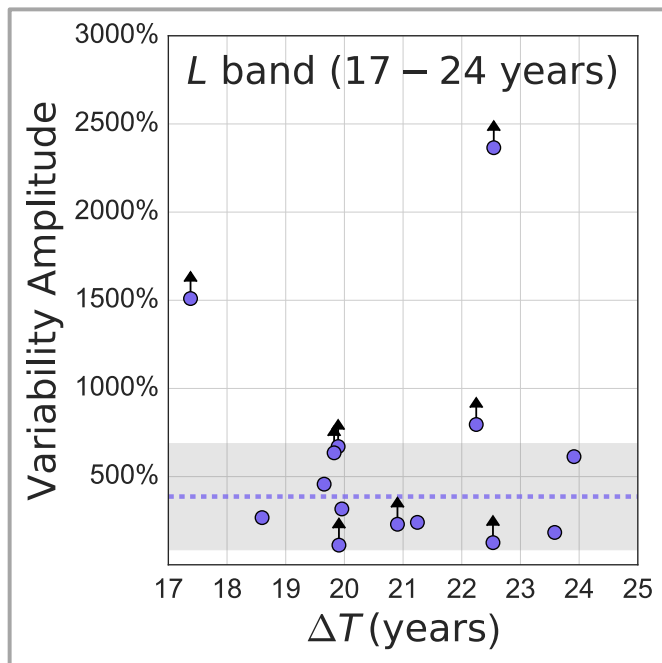
Simultaneous 1-18 GHz VLA Follow-up

- Single-band **variability**
- Radio **spectral shapes**

19A-422 (PI – Hallinan)
20B-329, 20B-459 (PI – Nyland)



Single-band Radio Variability

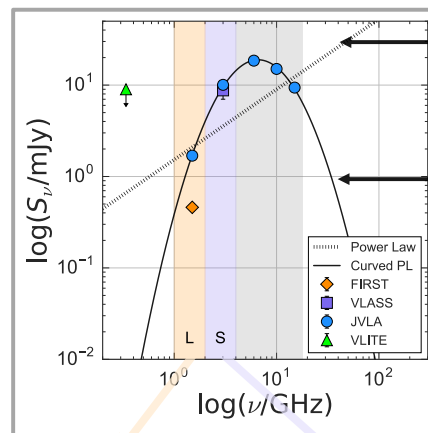
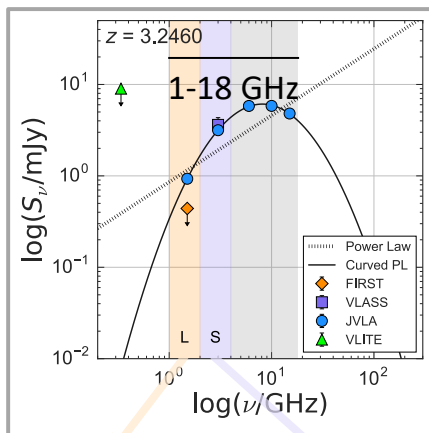


Variability amplitudes and timescales \rightarrow intrinsic radio AGN variability

Nyland et al. 2020

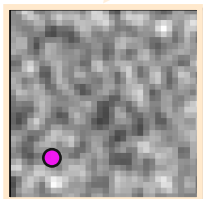
Radio Spectral Shapes

Nyland et al. 2020

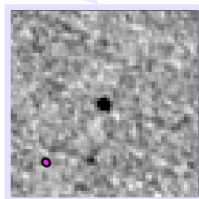


Power-law model doesn't fit

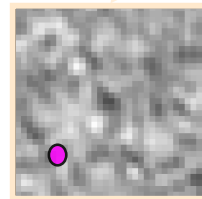
Curved power-law model fits!



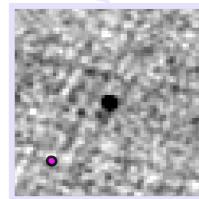
FIRST



VLASS



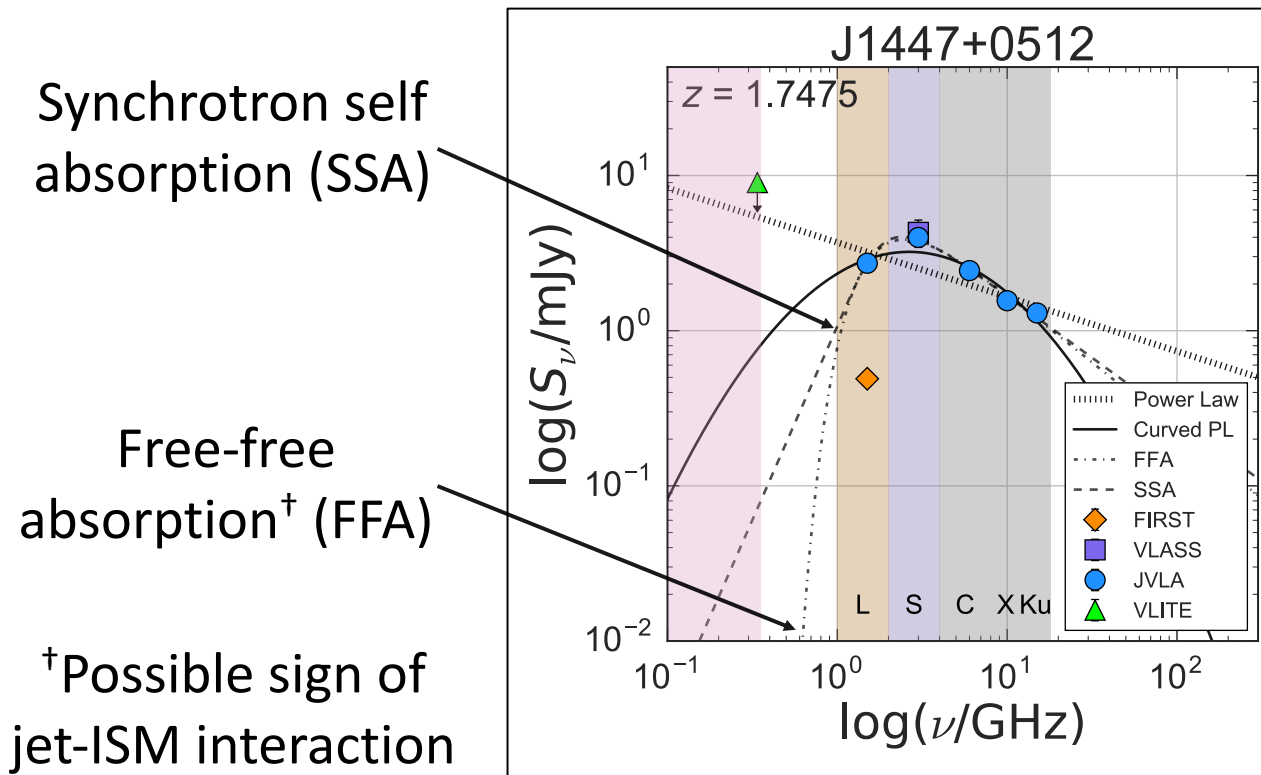
FIRST



VLASS

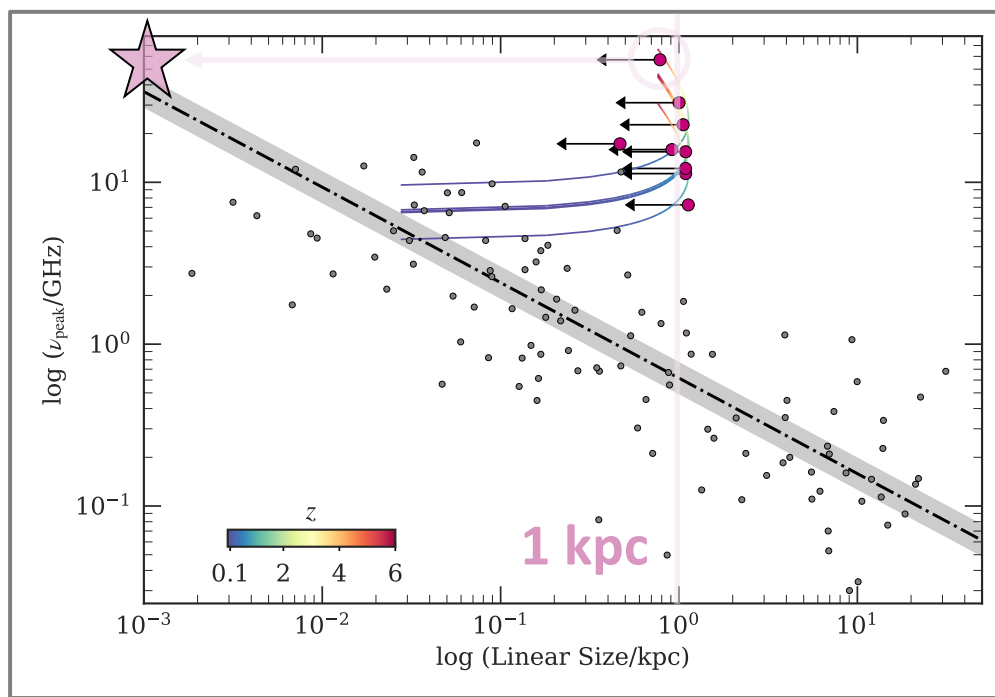
Peaked spectra consistent with absorbed and compact jets

Origin of the Absorption?



Deep sub-GHz data needed to separate SSA/FFA

Size Constraints from Turnover-size Relation



Consistent
with sub-kpc
jets launched
decades ago

Radio AGN Life Stages

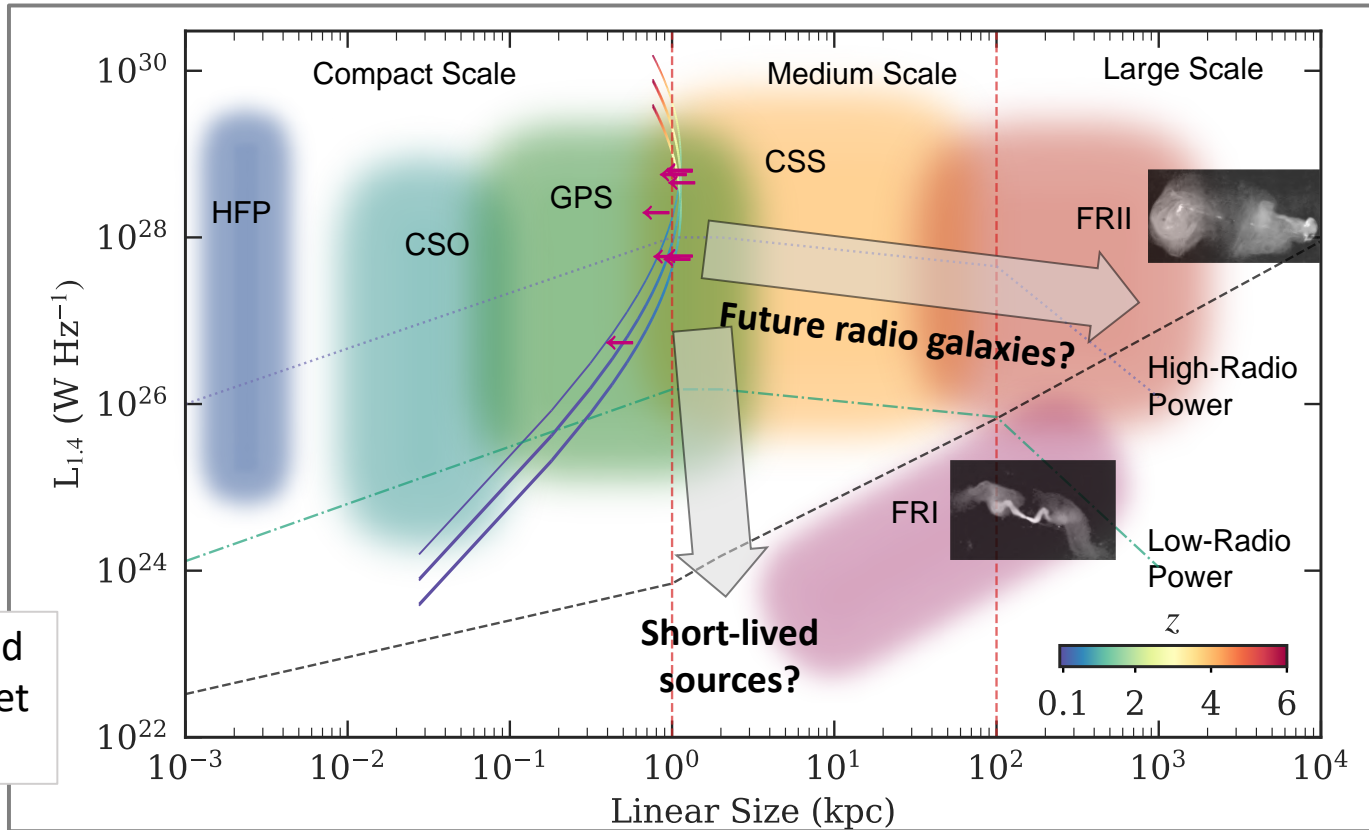
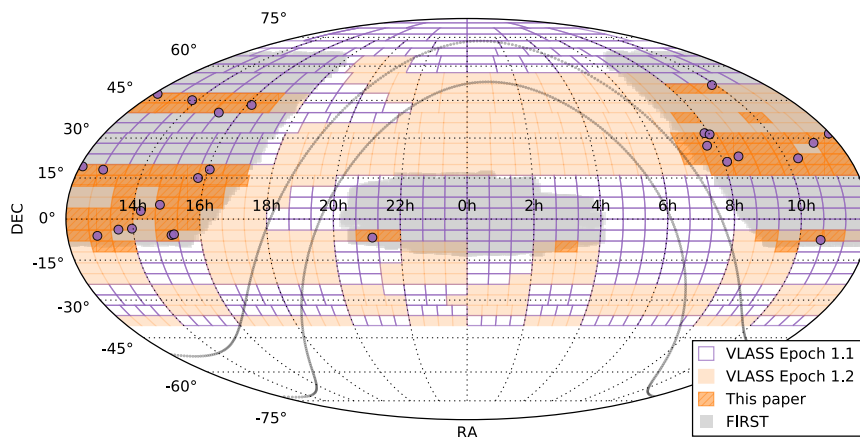


Fig adapted from Patil et al. 2020

Implications for Galaxy Evolution

Sky density:

$4 \times 10^{-3} \text{ deg}^{-2} \rightarrow$ period of occurrence = 10^5 yr



Nyland et al. 2020

Episodic, short-lived jets
common at $z = 1-3 \rightarrow$
Jet-ISM feedback?

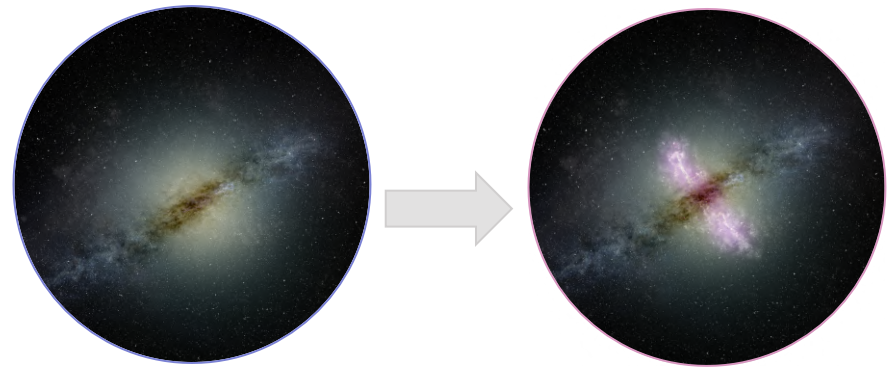


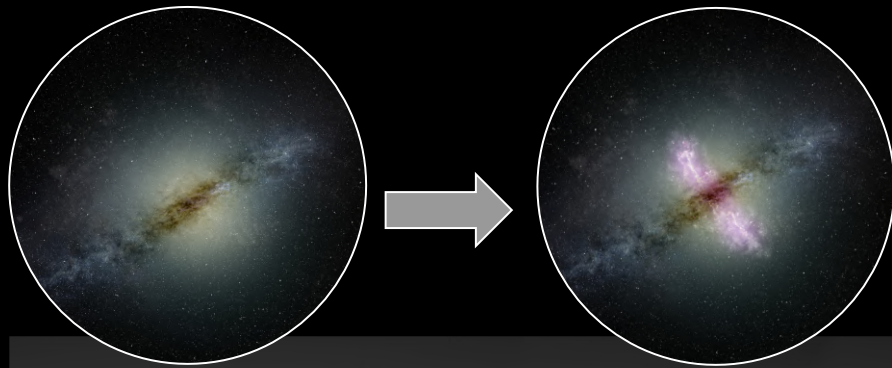
Image credit: Sophia Dagnello; NRAO/AUI/NSF

On-going and Future Work

- Continued radio SED monitoring (VLA)
- Milliarcsecond-scale imaging (VLBA)
- X-ray accretion state/morphology (new Cycle 22 *Chandra* data)
- Optical variability, host properties (new ground-based data, *HST*?)
- ISM content and conditions (ALMA?)



Summary



Nyland et al. 2020

Radio jets may “switch-on”
over *human timescales*

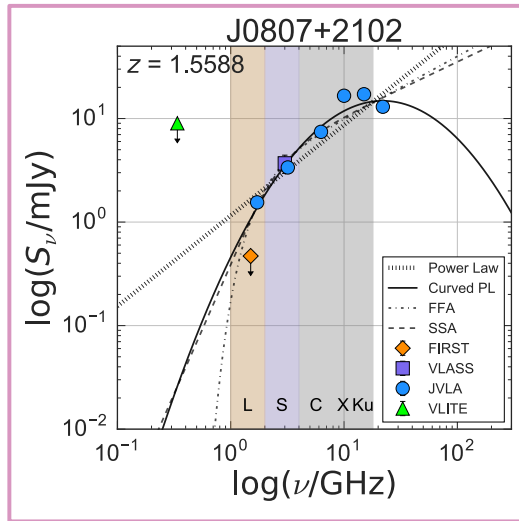
Multi-epoch radio surveys
catch newborn jets!

Short-lived jets common at
 $z=1-3 \rightarrow$ jet-ISM feedback?

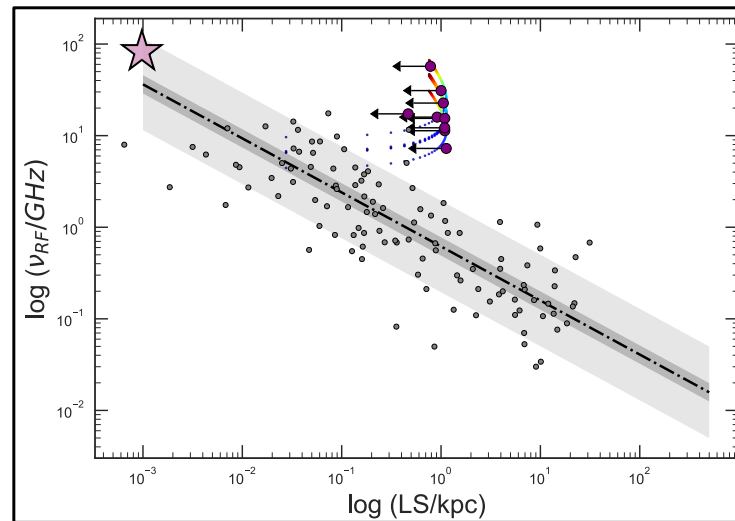
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Extra Slides

Turnover-Size Relation: Age Constraint Example



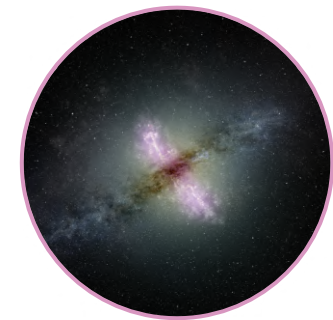
Spectral peak: 15 GHz
(= 40 GHz in restframe)



Size estimate: ~1-10 pc

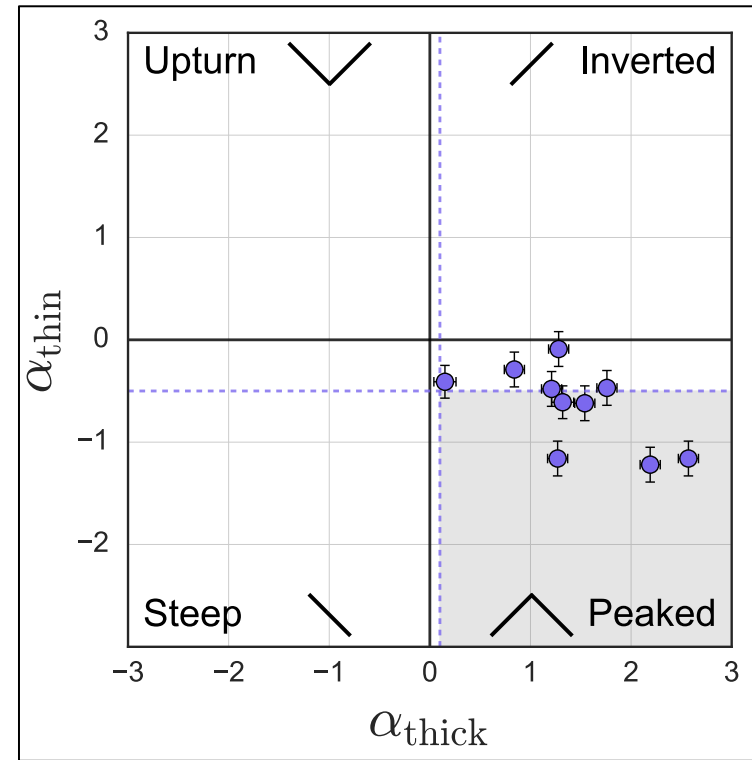
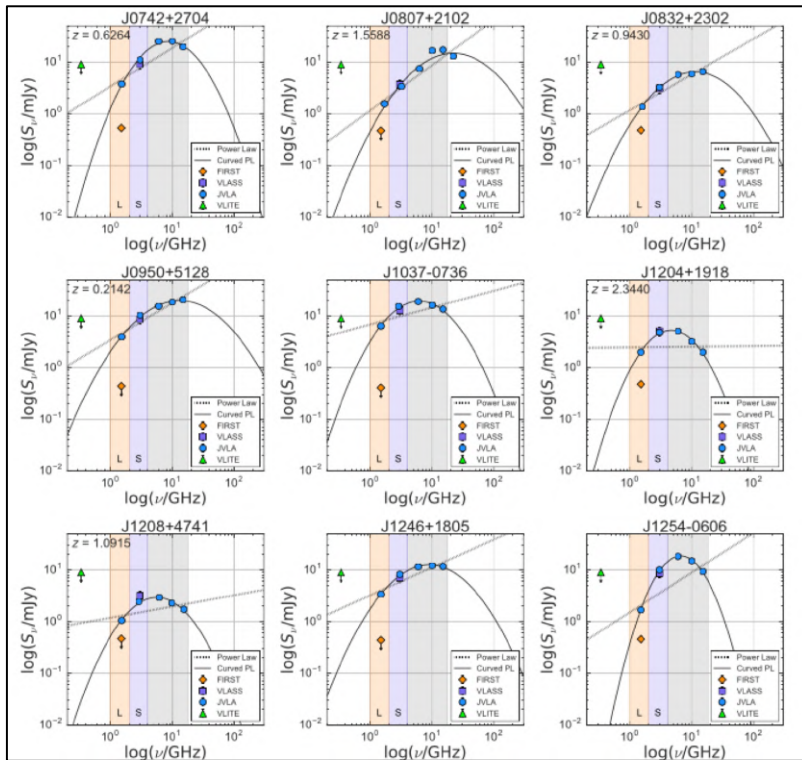
Velocity: 0.1c (assumed)

Age: $t=d/v \sim 30-300$ yr

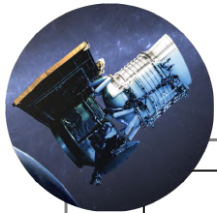


Young Jet!

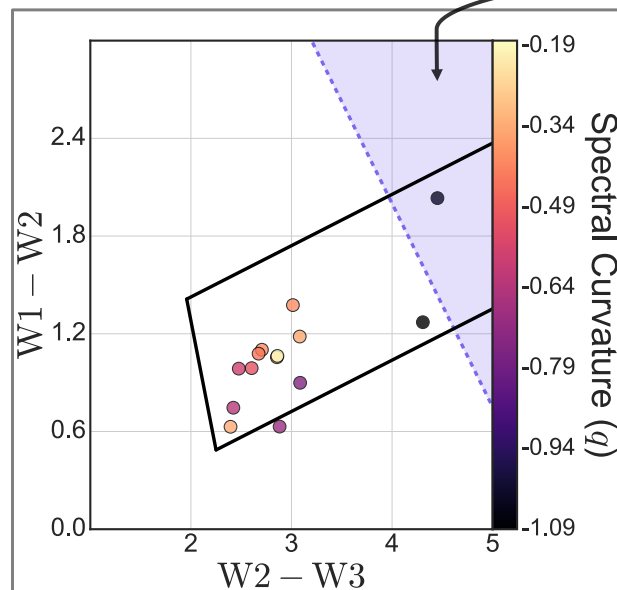
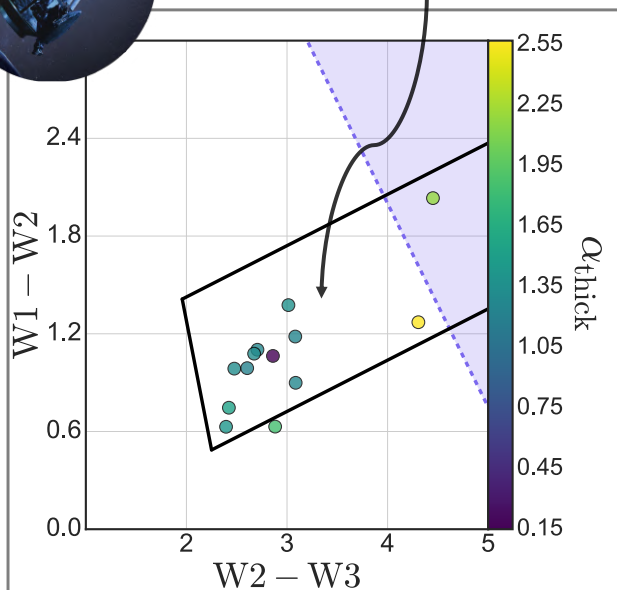
Radio Spectral Shapes



Connection with Quasar Reddening (and Mergers?)



Infrared AGN “wedge” (Matteos et al. 2012)



Nyland et al. 2020

Heavily obscured,
luminous AGN
(Lonsdale et al. 2015)

Free-free absorption
in dusty quasars with
compact jets?

More information on
host galaxies needed!