

# Jets, disc-winds and oscillations around Kerr black hole

“Extragalactic jets on all scales - launching, propagation, termination”, jointly organized by MPIA and IIT Indore.

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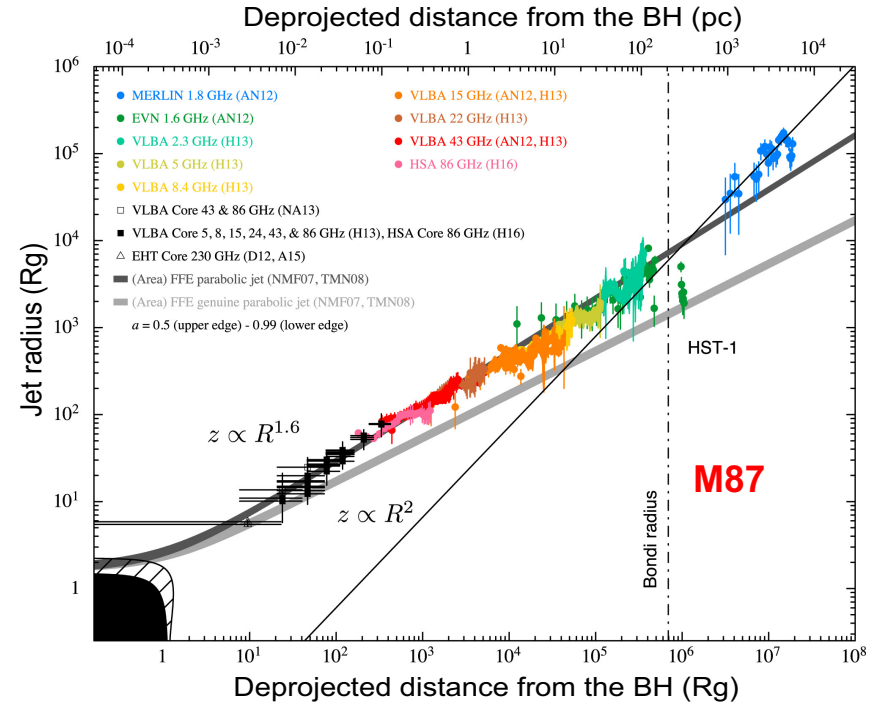
# Motivations:

## Torus based GRMHD simulations:

DeVilliers+2003, Gammie+2003, Noble+2006, DelZanna+2007, Tchekhovskoy+2010, Liska+2018, Nathanail+2020, and more.

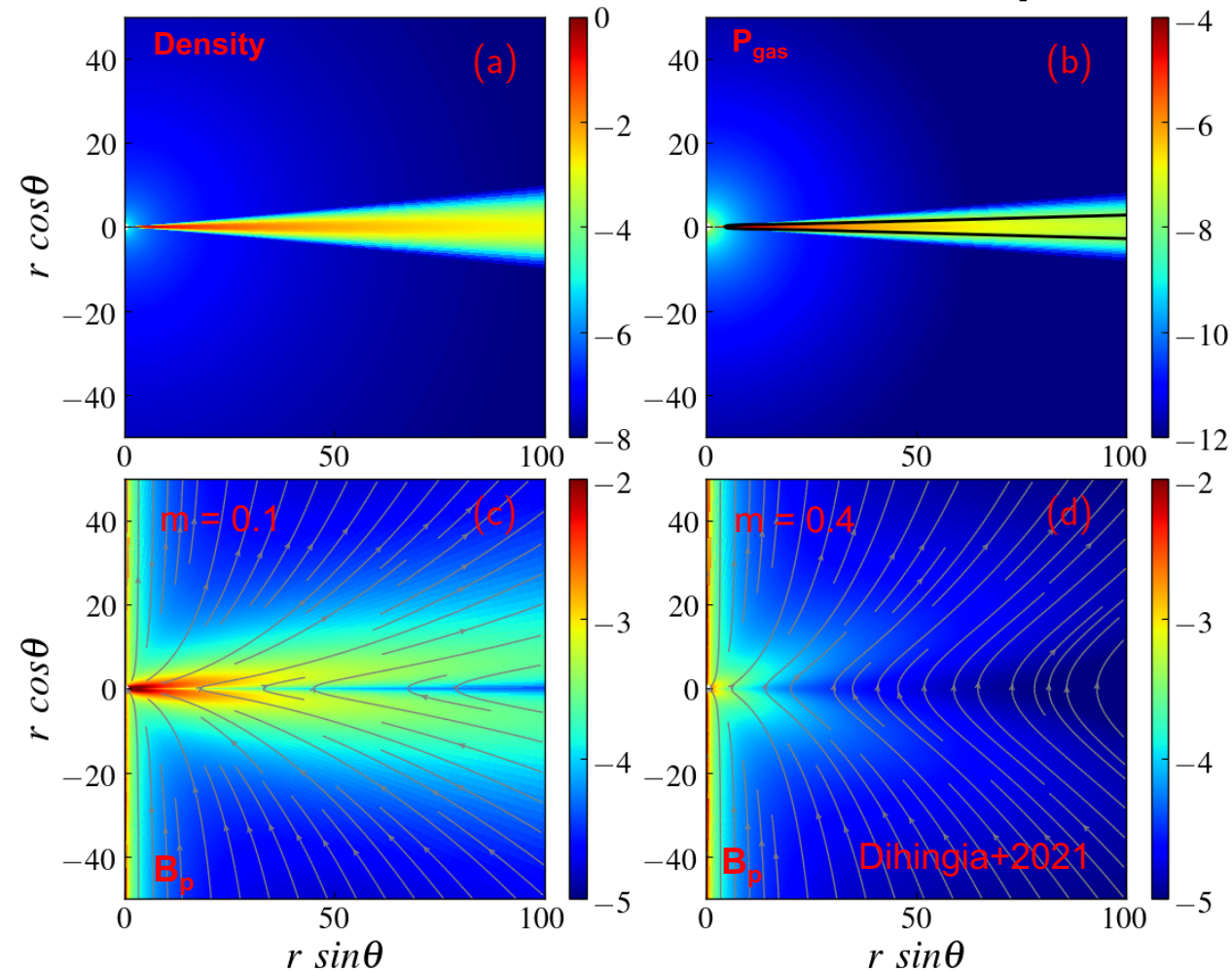
## Disc based GRMHD simulations:

Koide+1999, Qian+2018, Vourellis+2019,2021, Dihingia+2021.



Jet launching radius for M87 is about 5.5 Rg (Nakamura+2018), while for Cyg A it is about 227 Rg (Boccardi+2016)

# Initial setup:



**Initial setup for thin disc around Kerr BH. Using code BHAC**  
(Porth+2017; Olivares+2019)

**Our initial conditions are based on the Novikov & Thorne model** (Novikov & Thorne (1973)).

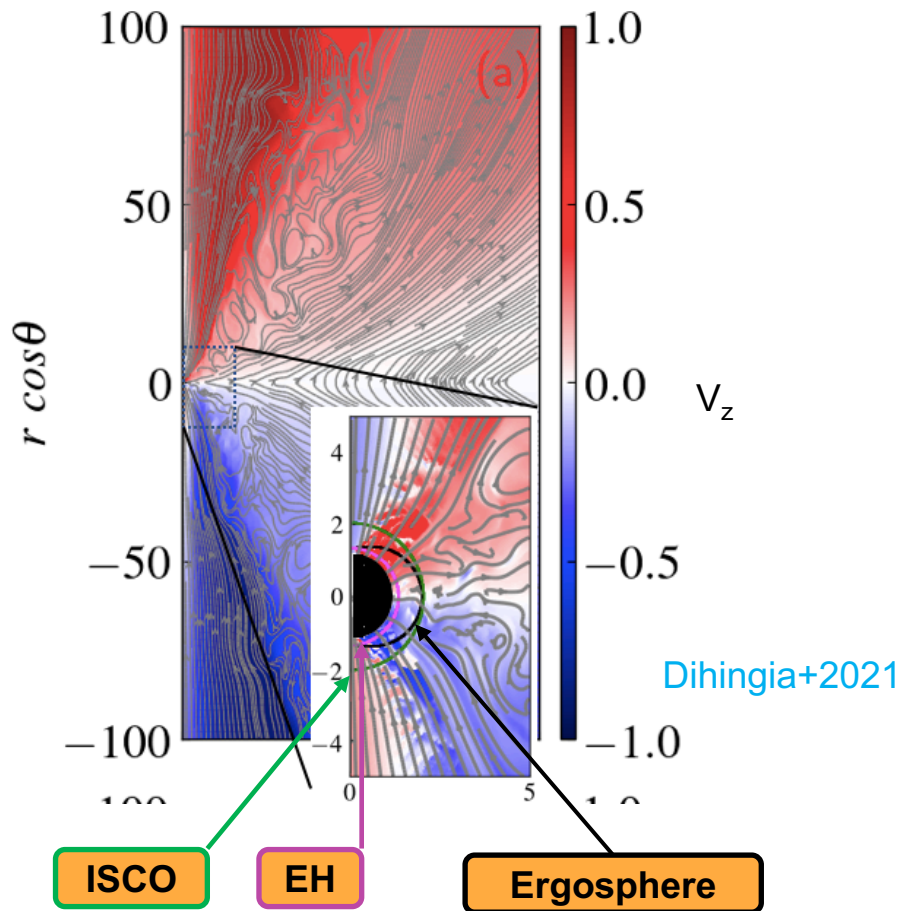
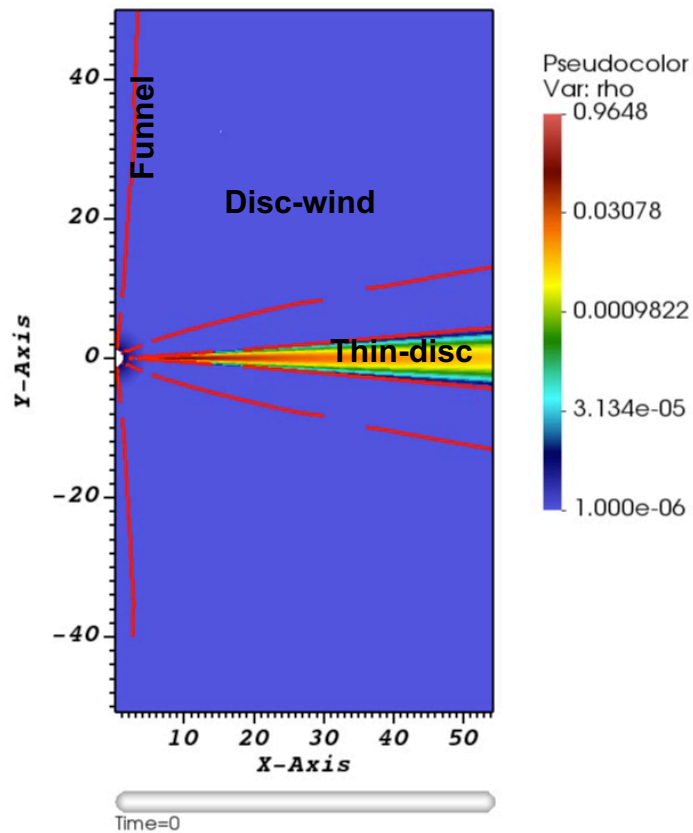
**Vector potential:**

$$\mathcal{A}_\phi \propto (r \sin \theta)^{3/4} \frac{m^{5/4}}{(m^2 + \tan^{-2}(\theta - \pi/2))^{5/8}},$$

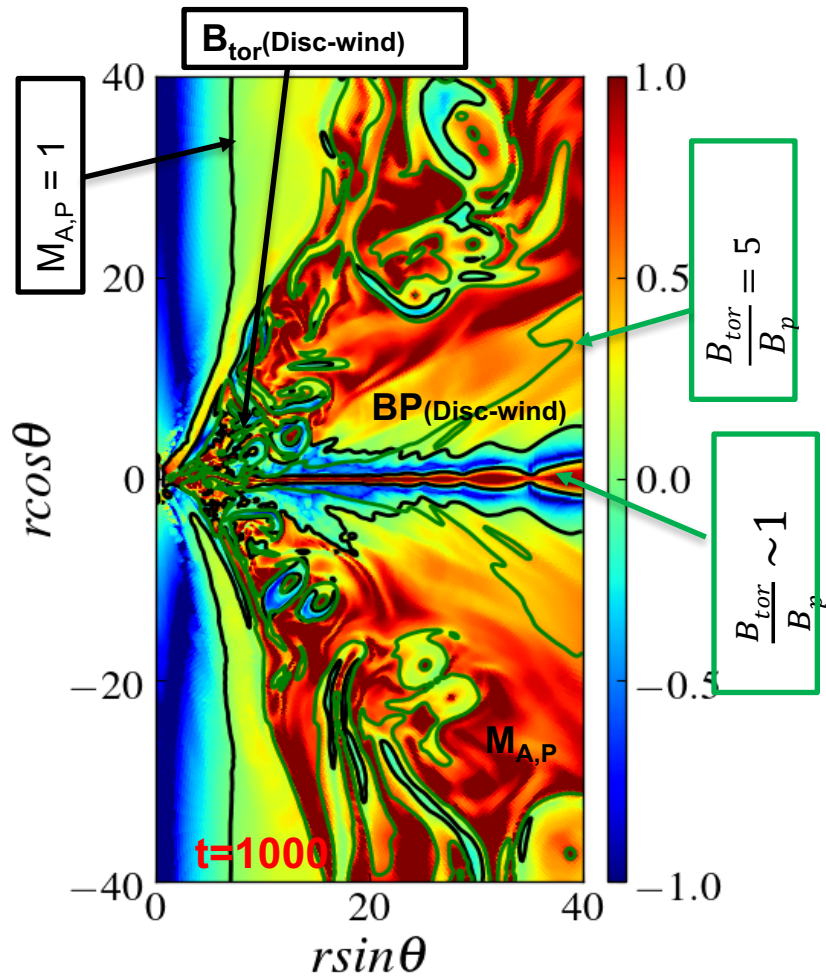
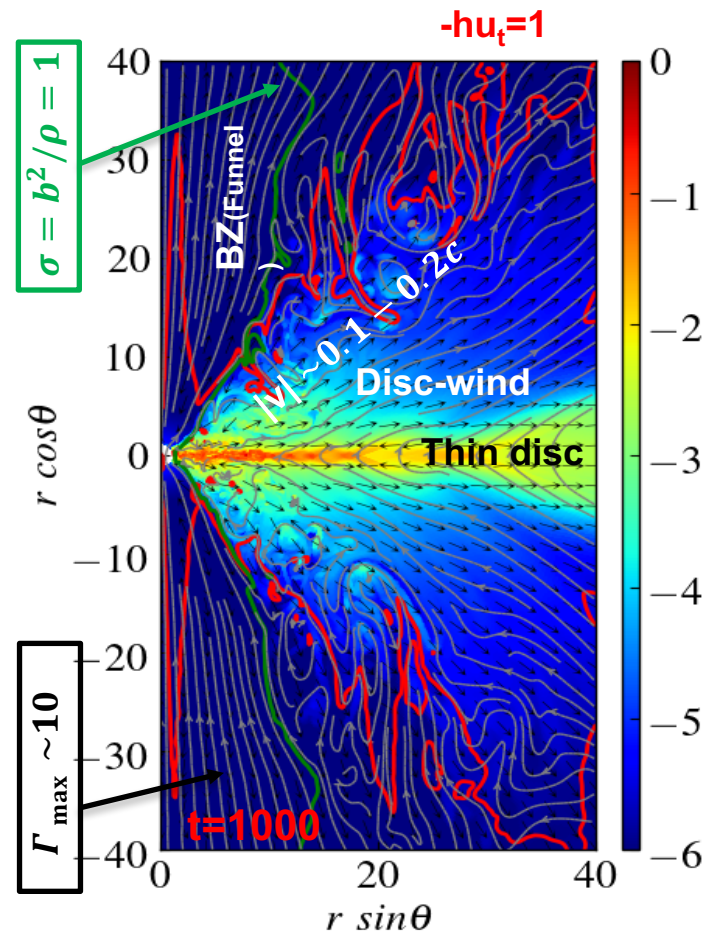
Zanni+2007

# Temporal Evolution

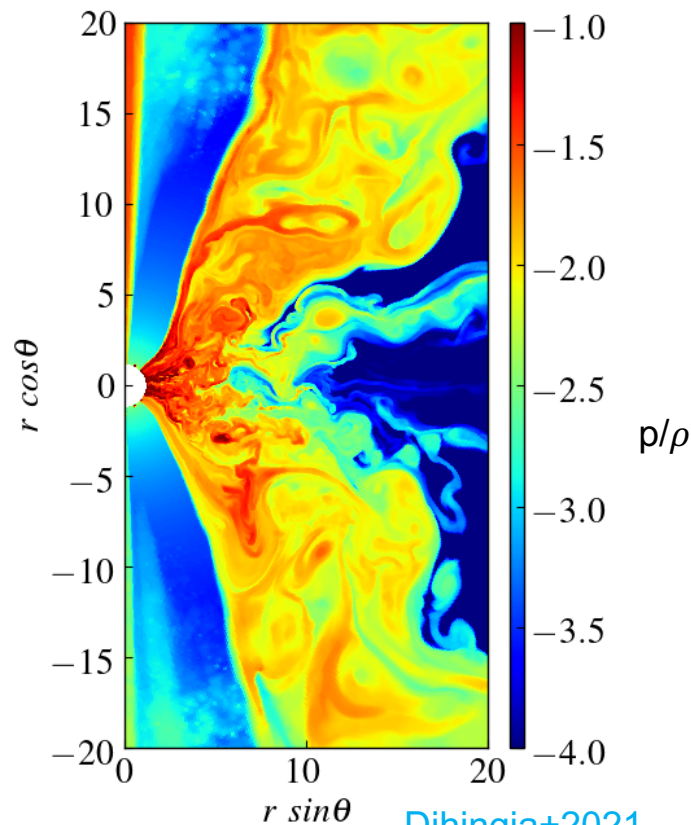
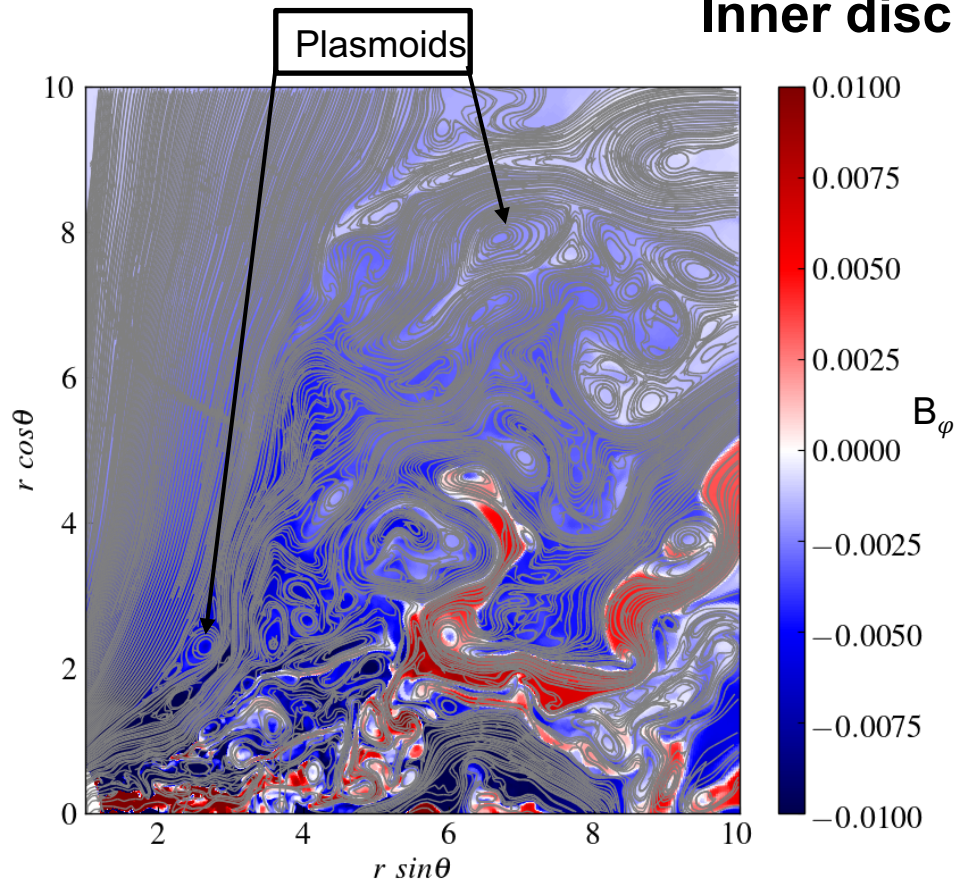
## Density evolution



# Wind and jet:



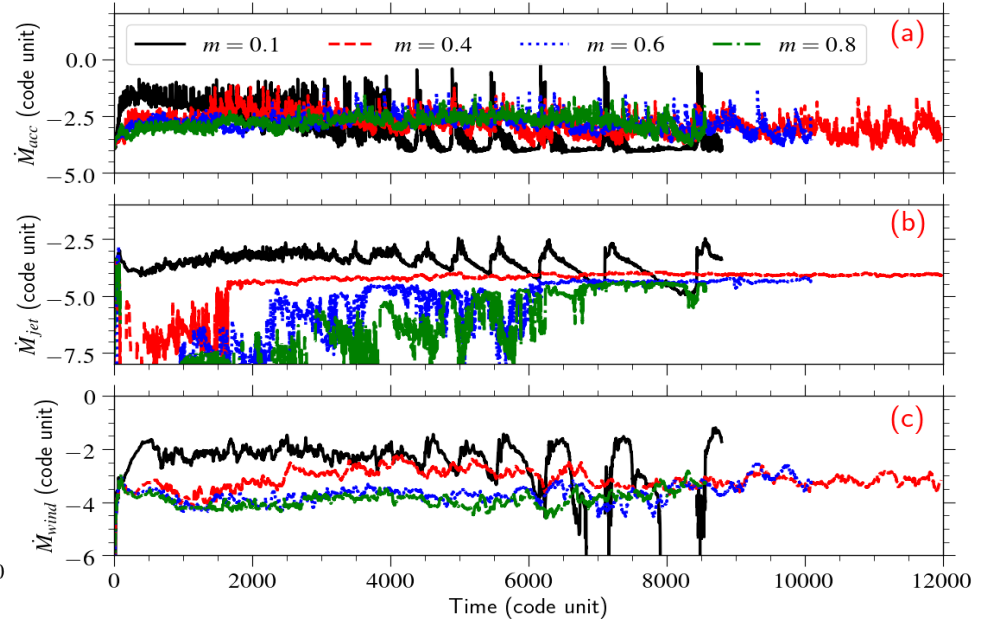
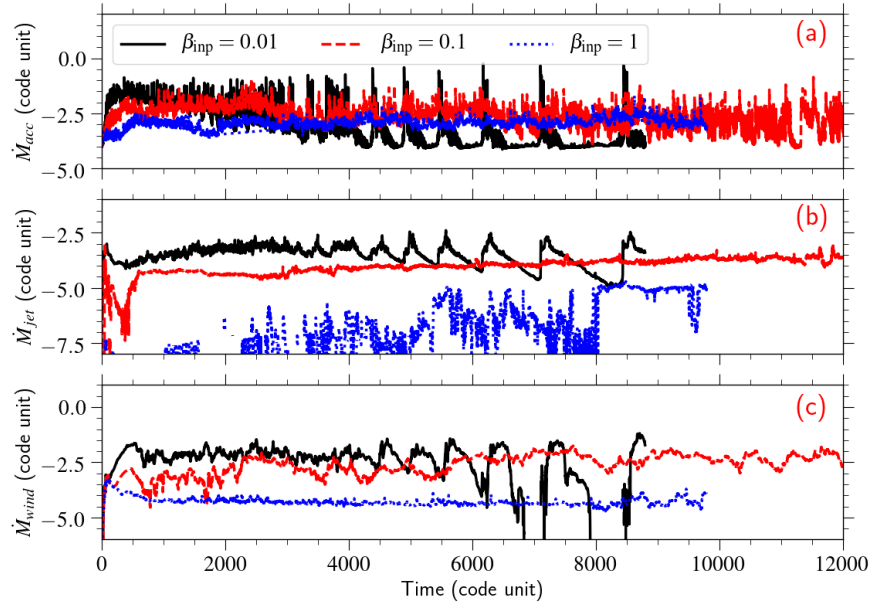
## Inner disc structure:



Dihingia+2021

Source of turbulence: **MRI** ( $\beta > 1$ ) and **MRTI** ( $\beta < 1$ ).

# Mass flux rates:



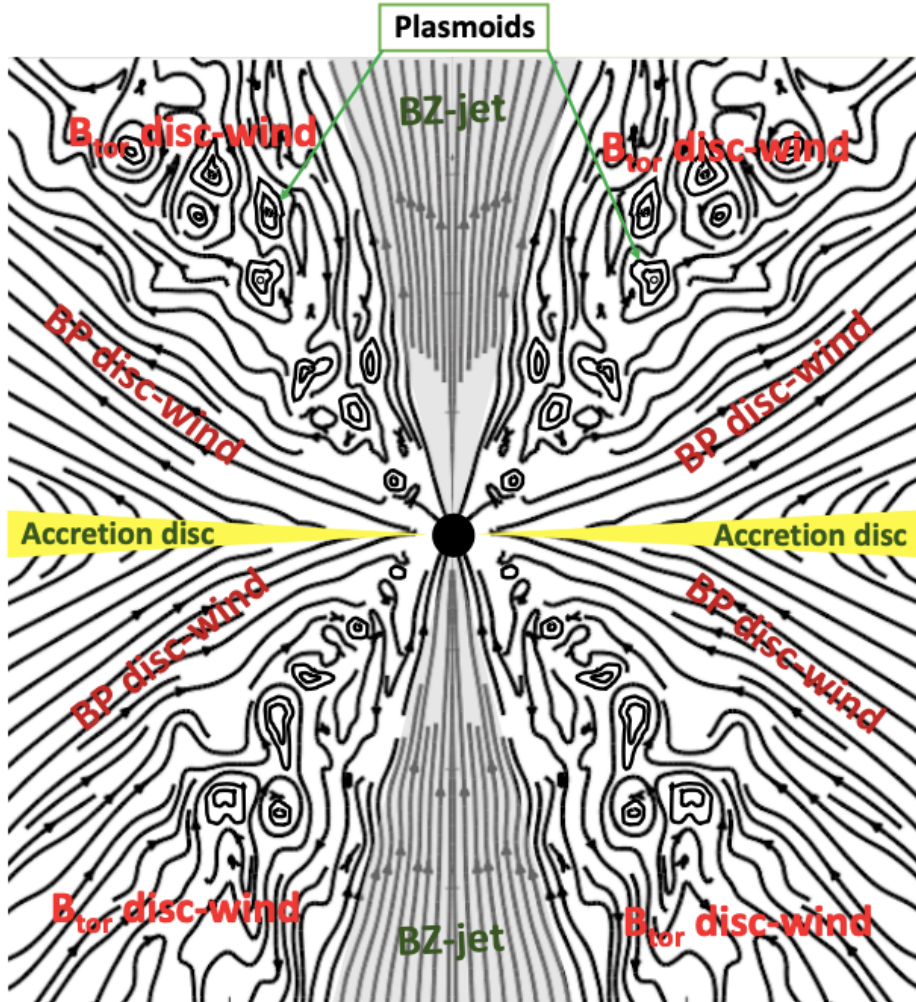
$$\dot{M} = 2\pi \int \sqrt{-g\rho u^r} d\theta.$$

$$\sigma < 1 \text{ or } \eta < 2, \text{ and } -hu_t > 1. \quad \dot{M}_{wind}$$

$$\sigma > 1 \text{ or } \eta > 2, \text{ and } -hu_t > 1 \quad \dot{M}_{jet}$$

Dihingia+2021

## Toy model:



## Summary:

1. We Observe **BZ jet**, turbulent  **$B_{\text{tor}}$  disc-wind**, and **BP disc-wind** in our simulation.
2. Lower value of  $m$  and plasma beta are prone to **BP** driven wind.
3. **Plasmoids** are formed in the  **$B_{\text{tor}}$  disc-wind**.
4. The oscillations in the inner part of the accretion disc can be relate to flaring activities in the jet.