

# Analytical Model of Magnetically Dominated Jet:

— jet launching, acceleration and collimation

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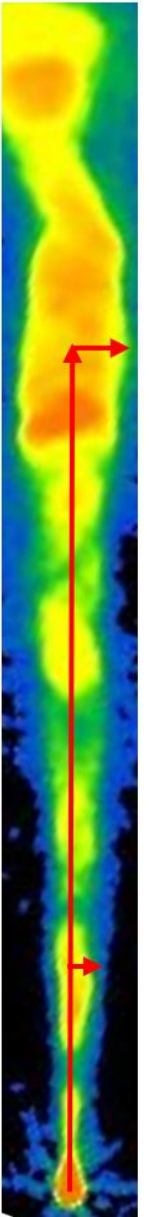
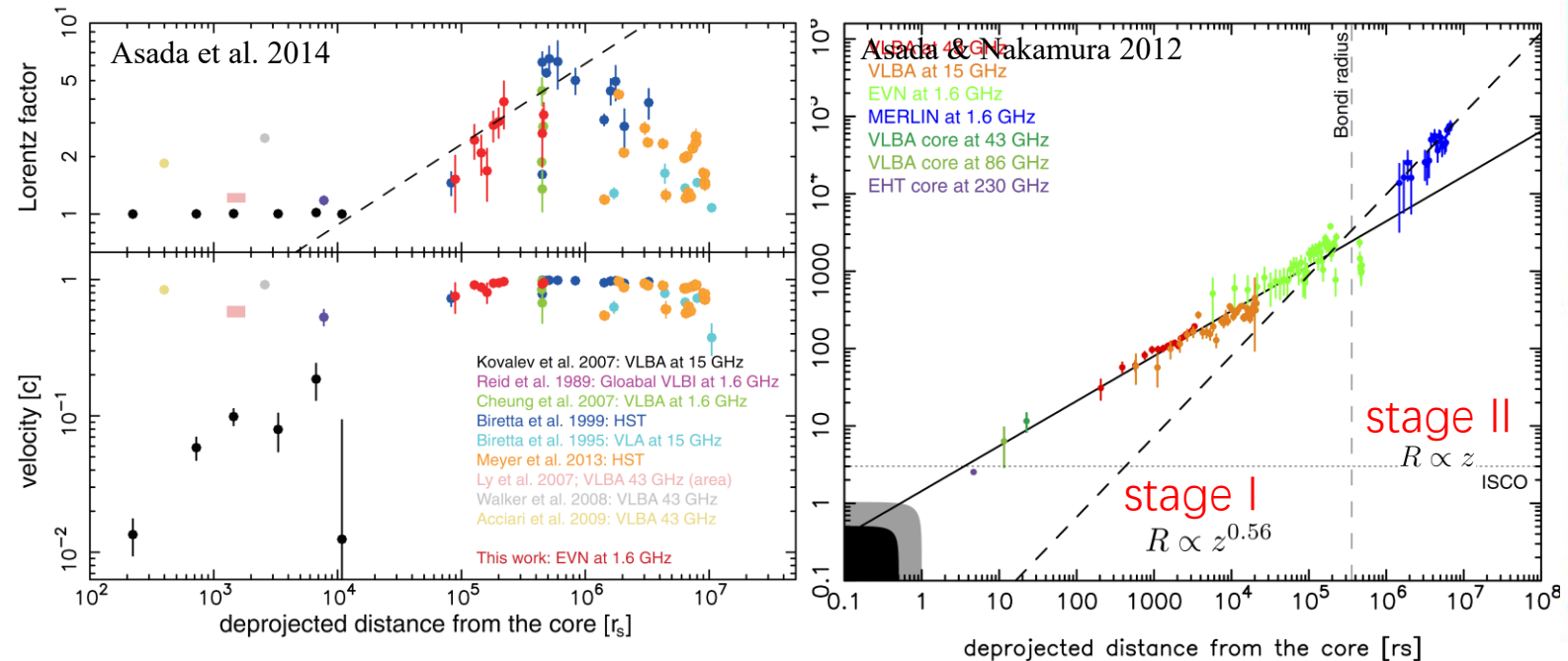
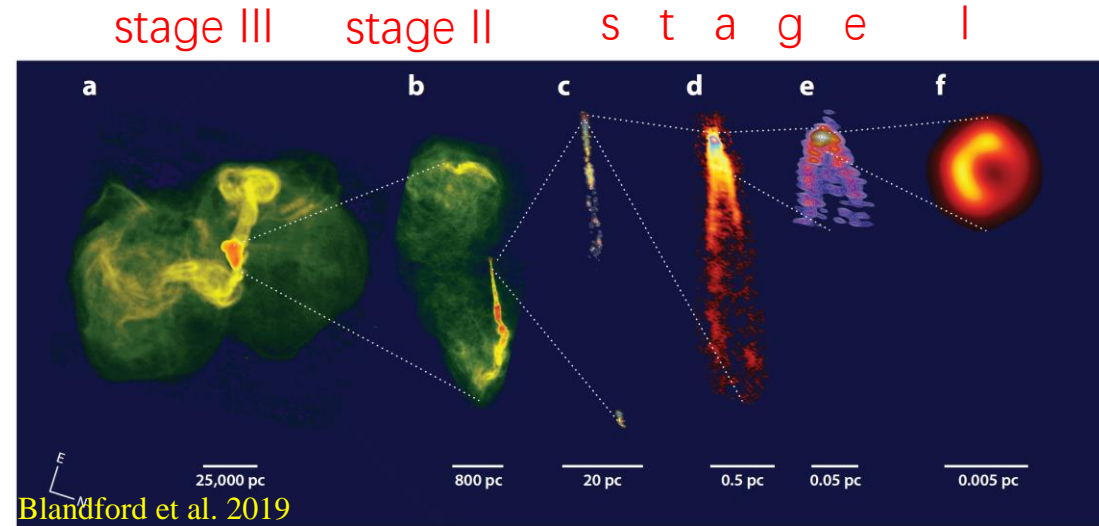
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# Jets: evolution

- stage I:
- collimating
- accelerating
- “parabolic”
- stage II:
- “collimating”
- “conical”
- “stage III”
- terminal, lobe



Jets: stage I: magnetically dominated  $\Rightarrow$  “force-free”

$$\mathbf{B} = \frac{1}{r^2 \sin \theta} \frac{\partial \Psi}{\partial \theta} \hat{\mathbf{r}} - \frac{1}{r \sin \theta} \frac{\partial \Psi}{\partial r} \hat{\theta} + \frac{\Phi}{r \sin \theta} \hat{\phi}$$

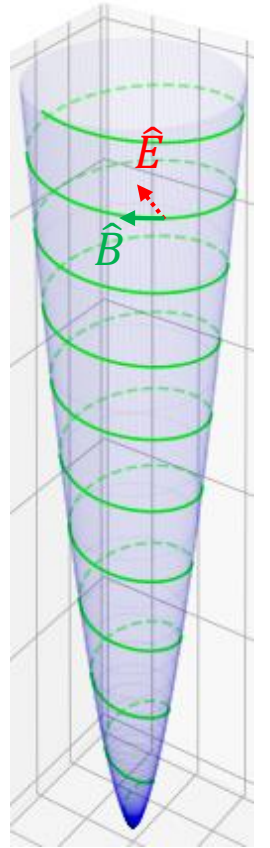
$$\boxed{\Psi} = r \sin \theta A_\phi \quad \boxed{\Phi} = r \sin \theta B_\phi$$

$$\mathbf{B} = \nabla \times \mathbf{A}$$

$$\mathbf{E} = -\boxed{\Omega} \nabla \Psi = -\Omega r \sin \theta \hat{\phi} \times \mathbf{B}$$

$$\frac{\mathbf{j}}{\sigma_c} = \mathbf{E} + \mathbf{v} \times \mathbf{B} \stackrel{\text{ideal MHD}}{=} 0$$

$$\rho (\mathbf{u} \cdot \nabla) \mathbf{u} = \rho_e \mathbf{E} + \mathbf{j} \times \mathbf{B} \stackrel{\text{force free}}{=} 0.$$



$$+\Phi' \Phi - \left\{ \frac{\Omega'}{\Omega} \left[ \left( \frac{\partial \Psi}{\partial r} \right)^2 + \left( \frac{1}{r} \frac{\partial \Psi}{\partial \theta} \right)^2 \right] + \frac{\partial^2 \Psi}{\partial r^2} + \frac{2}{r} \frac{\partial \Psi}{\partial r} + \frac{1}{r^2} \frac{\partial^2 \Psi}{\partial \theta^2} + \frac{\cot \theta}{r^2} \frac{\partial \Psi}{\partial \theta} \right\} (\Omega r \sin \theta)^2 = 0$$

the “pulsar” equation (established 1960s)

# Jets: solve equation

- simulation: BH rotating **slow or fast** produce **similar jet configuration**  
(e.g., Tchekhovskoy, McKinney & Narayan 2008)
- **math** expect: two terms (equations): **non-rotating** and **rotating**

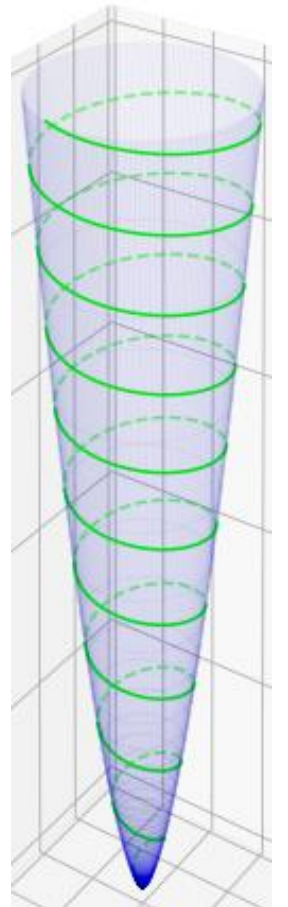
$$\frac{\partial^2 \Psi}{\partial r^2} + \frac{1}{r^2} \frac{\partial^2 \Psi}{\partial \theta^2} - \frac{\cot \theta}{r^2} \frac{\partial \Psi}{\partial \theta} + \Phi' \Phi - \left\{ \frac{\Omega'}{\Omega} \left[ \left( \frac{\partial \Psi}{\partial r} \right)^2 + \left( \frac{1}{r} \frac{\partial \Psi}{\partial \theta} \right)^2 \right] + \frac{\partial^2 \Psi}{\partial r^2} + \frac{2}{r} \frac{\partial \Psi}{\partial r} + \frac{1}{r^2} \frac{\partial^2 \Psi}{\partial \theta^2} + \frac{\cot \theta}{r^2} \frac{\partial \Psi}{\partial \theta} \right\} (\Omega r \sin \theta)^2 = 0$$

if real

non-rotating  $\frac{\partial^2 \Psi}{\partial r^2} + \frac{1}{r^2} \frac{\partial^2 \Psi}{\partial \theta^2} - \frac{\cot \theta}{r^2} \frac{\partial \Psi}{\partial \theta} = 0$

rotating  $\frac{\Phi' \Phi}{\Omega^2 r^2 \sin^2 \theta} - \left\{ \frac{\Omega'}{\Omega} \left[ \left( \frac{\partial \Psi}{\partial r} \right)^2 + \left( \frac{1}{r} \frac{\partial \Psi}{\partial \theta} \right)^2 \right] + \frac{\partial^2 \Psi}{\partial r^2} + \frac{2}{r} \frac{\partial \Psi}{\partial r} + \frac{1}{r^2} \frac{\partial^2 \Psi}{\partial \theta^2} + \frac{\cot \theta}{r^2} \frac{\partial \Psi}{\partial \theta} \right\} = 0$

- The two solutions match each other!



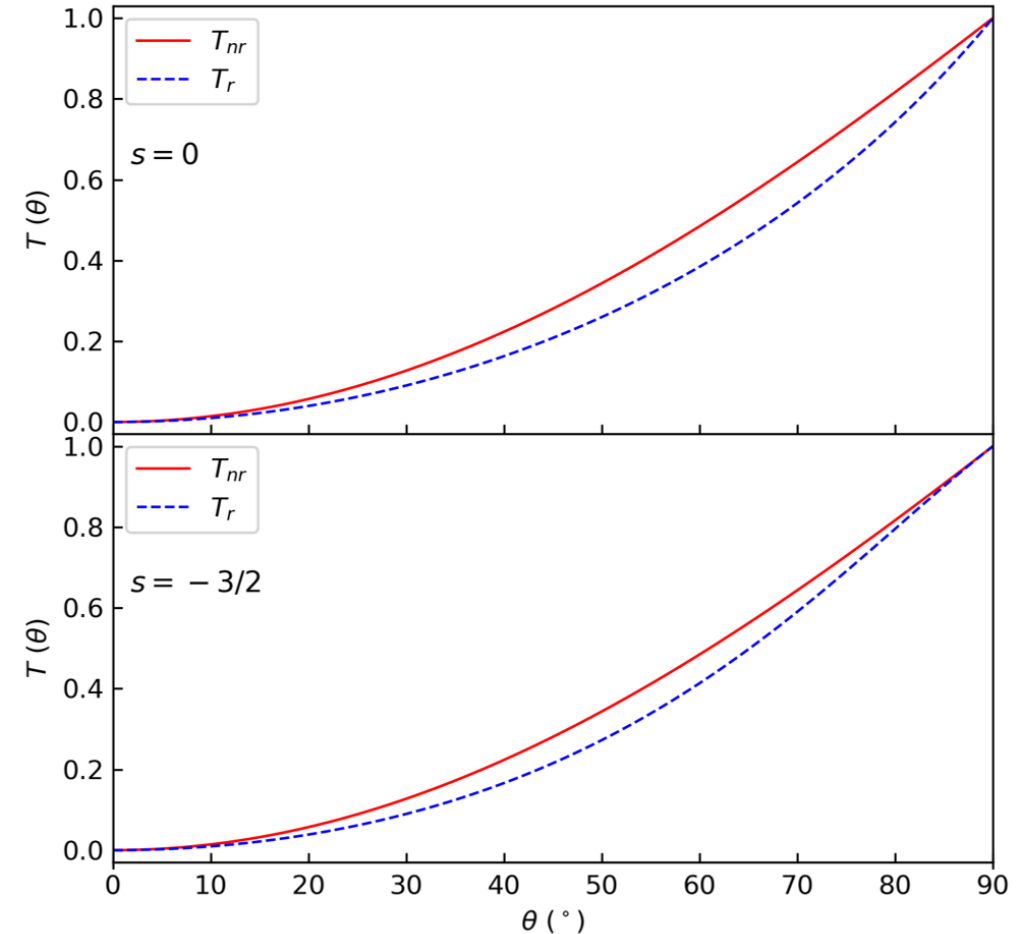
# Jets: solve equation

$$\Psi = r^\nu T_{\text{nr}}(\theta) \quad 0 \leq \nu \leq 2$$
$$T_{\text{nr}}(\theta) = C_2 y {}_2F_1 \left( 1 - \frac{\nu}{2}, \frac{1}{2} + \frac{\nu}{2}, 2, y \right)$$
$$T_r(y) = A_2 e^{\frac{\nu}{s+\nu}} \int_1^y \frac{G_1(t) + A_1 G_2(t)}{A_1 G_3(t) + G_4(t)} dt$$

non-relativistic to relativistic regimes

apply:  $\theta \ll 1$  or  $\theta \rightarrow \pi/2$

$\Omega r \sin \theta \gg 1$  or  $\Omega r \sin \theta \ll 1$



# Jets: magnetic field and velocity

- magnetic field

$$B_p = \frac{2\Psi}{R^2}$$

$$B_\phi = -\frac{2\Omega\Psi}{R}$$

- drift velocity  $\leftrightarrow$  cold plasma velocity

$$D_{\text{fd}} \equiv \frac{(v\Gamma)^2 - (v_d\Gamma_d)^2}{(v_d\Gamma_d)^2} \ll 1 \quad (\Omega R \gg 1 \text{ or } \Omega R \ll 1)$$

- velocity

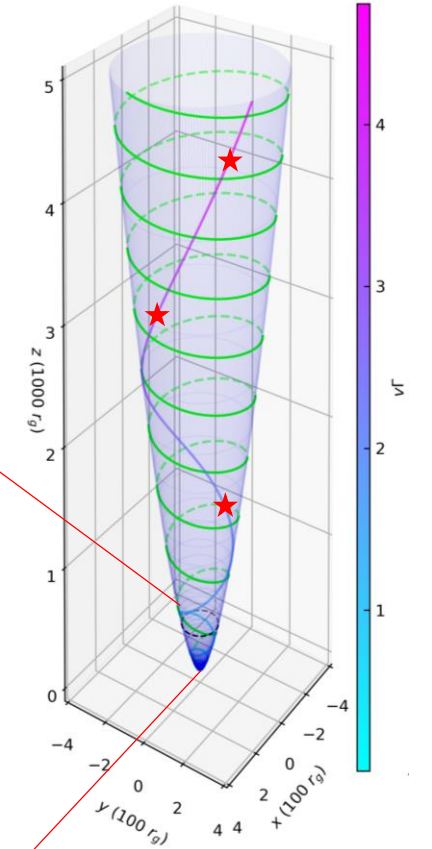
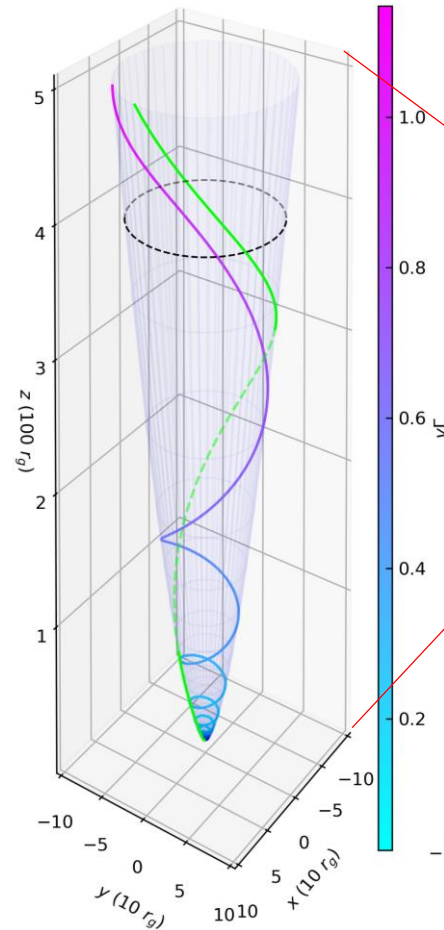
$$v_\phi = \Omega r \sin \theta \frac{B_p^2}{B^2} \approx \frac{\Omega R}{1 + (\Omega R)^2},$$

$$v_p = -\Omega r \sin \theta \frac{B_\phi B_p}{B^2} \approx \frac{(\Omega R)^2}{1 + (\Omega R)^2},$$

$$v = \Omega r \sin \theta \frac{B_p}{B} \approx \frac{\Omega R}{\sqrt{1 + (\Omega R)^2}},$$

$$v\Gamma \approx \Omega R.$$

From non-relativistic to relativistic



Helical jet

Chen & Zhang 2021

Consist with previous asymptotic results at ultra-relativistic regime (Blandford, Narayan, Tchekhovskoy, Beskin, Komissarov, Lyubarsky, ...)

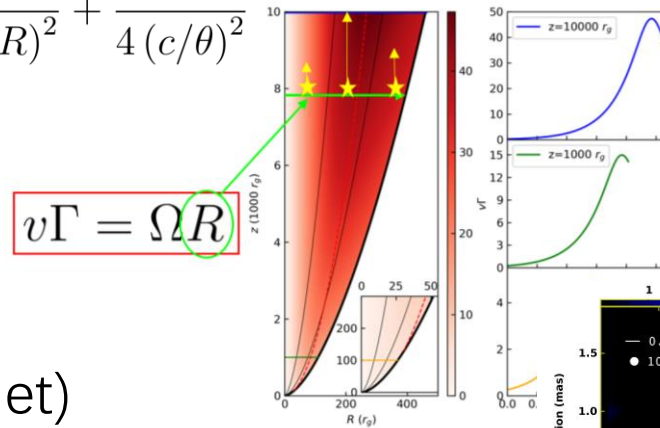
# Jets: acceleration and collimation

- configuration : (quasi-parabolic at  $\theta \ll 1$ )

$$\Psi = Cr^\nu \sin^2 \theta {}_2F_1\left(1 - \frac{\nu}{2}, \frac{1}{2} + \frac{\nu}{2}, 2, \sin^2 \theta\right) \Rightarrow R = C_2^{-1/2} \Psi^{1/2} z^{1-\nu/2}$$

- acceleration: stages I, II, III (non-relativistic to relativistic)

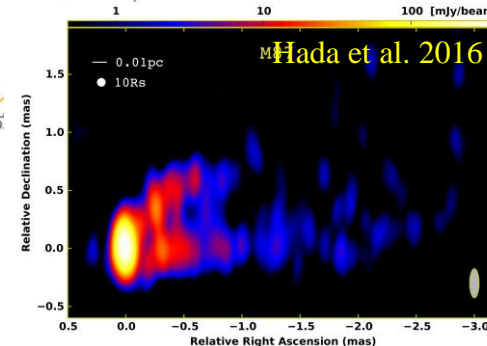
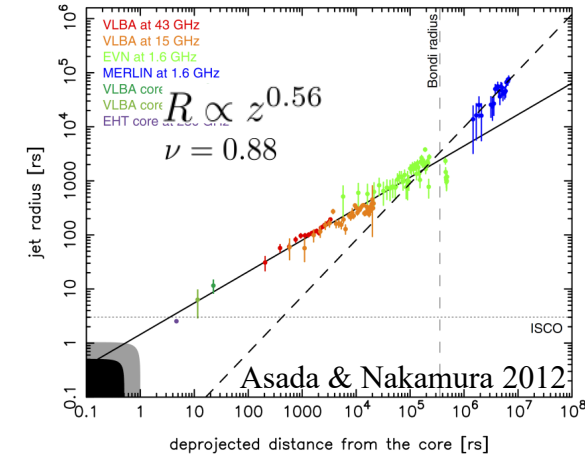
$$\frac{1}{(v\Gamma)^2} \simeq \frac{1}{(\Omega R)^2} + \frac{2-\nu}{4(c/\theta)^2}$$



- spine/layer jet

- hollow jet (for a BZ jet)

$$S_z = \frac{B_\phi^2}{4\pi} = \frac{C_2}{\pi} \Omega^2 \Psi r^{\nu-2} = \frac{C_2^{2/\nu}}{\pi} \Omega^2 \Psi^{2-2/\nu} \theta^{-2+4/\nu} = \frac{\alpha^2}{\pi} C_2^{2\lambda+2} z^{2\lambda\nu+2\nu-4\lambda-4} R^{4\lambda+2}$$



Chen & Zhang 2021

Consist with previous asymptotic results at ultra-relativistic regime (Blandford, Narayan, Tchekhovskoy, Beskin, Komissarov, Lyubarsky, ...)

# Jets: current and charge

- electric current

$$J = \sqrt{cP_{\text{jet}}} \approx 5.8 \times 10^{17} \sqrt{P_{44}} \text{ A}$$

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$\sim 3.9 \times 10^{18} \text{ A}$  Kronberg et al. 2011

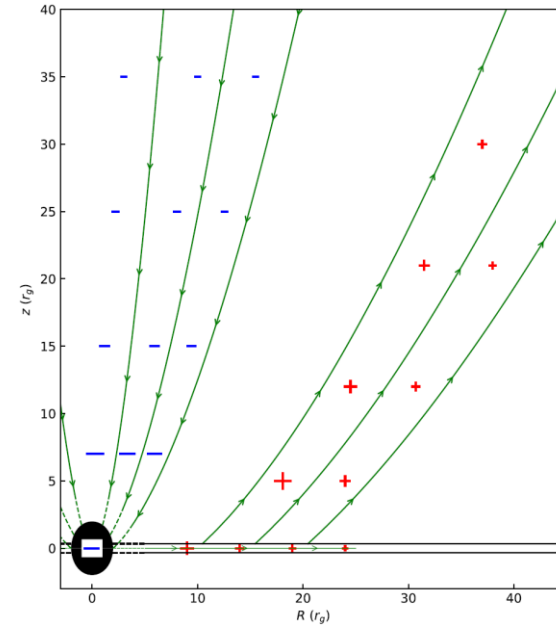
$\sim 1.0 \times 10^{46} \text{ erg s}^{-1}$  Zhang et al. 2018

- electric potential difference  
("gap" near BH horizon)

$$\Delta V = \sqrt{P_{\text{jet}}/c} \approx 1.7 \times 10^{19} \sqrt{P_{44}} \text{ Volts}$$

- black hole charge

$$r_Q = \sqrt{GQ/M} \approx \sqrt{GP_{\text{jet}}/c^5} \approx 1.7 \times 10^{-8} \sqrt{P_{44}}$$



Chen & Zhang 2021

Consist with previous asymptotic results at ultra-relativistic regime (Blandford, Narayan, Tchekhovskoy, Beskin, Komissarov, Lyubarsky, ...)



Thanks!