



The birth of the giants: gas & dust in QSO host galaxies at $z \sim 6$

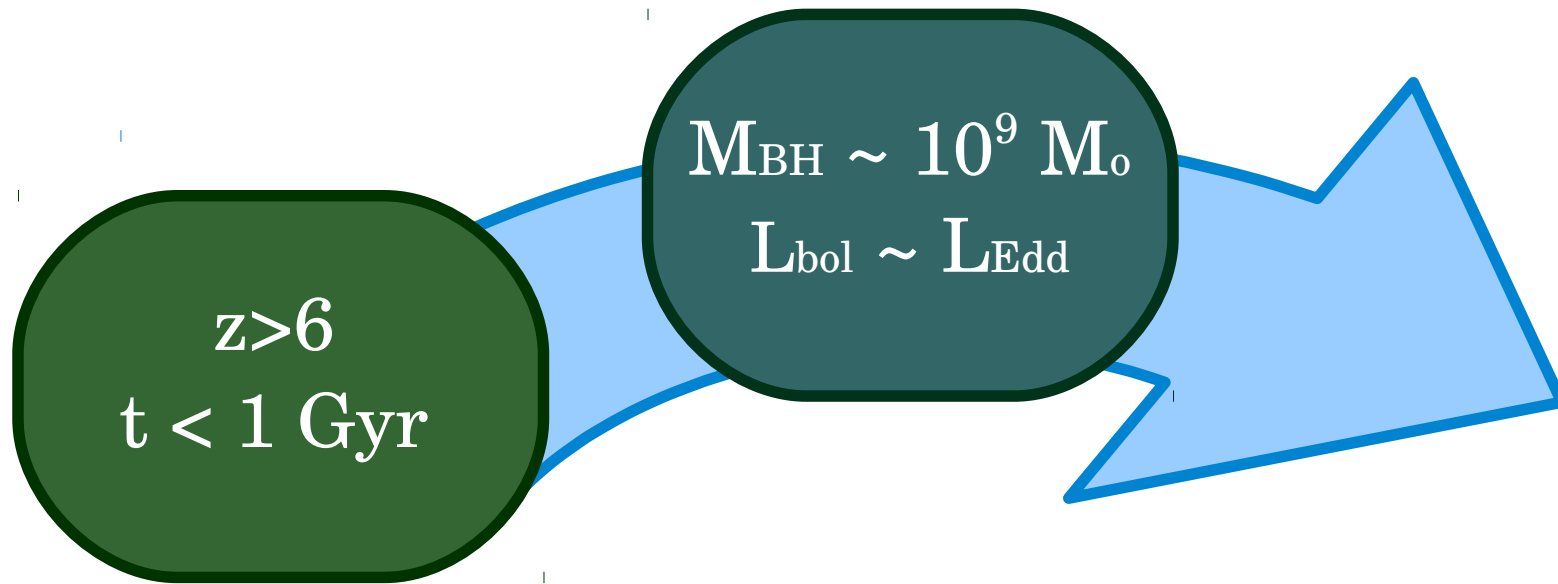
Roberto Decarli
MPIA, Heidelberg

Fabian Walter, Bram Venemans, Emanuele Paolo Farina, Eduardo Banados,
Chiara Mazzucchelli, Xiaohui Fan, Frank Bertoldi, Chris Carilli, Dominik Riechers,
Hans-Walter Rix, Michael Strauss, Ran Wang, Yujin Yang

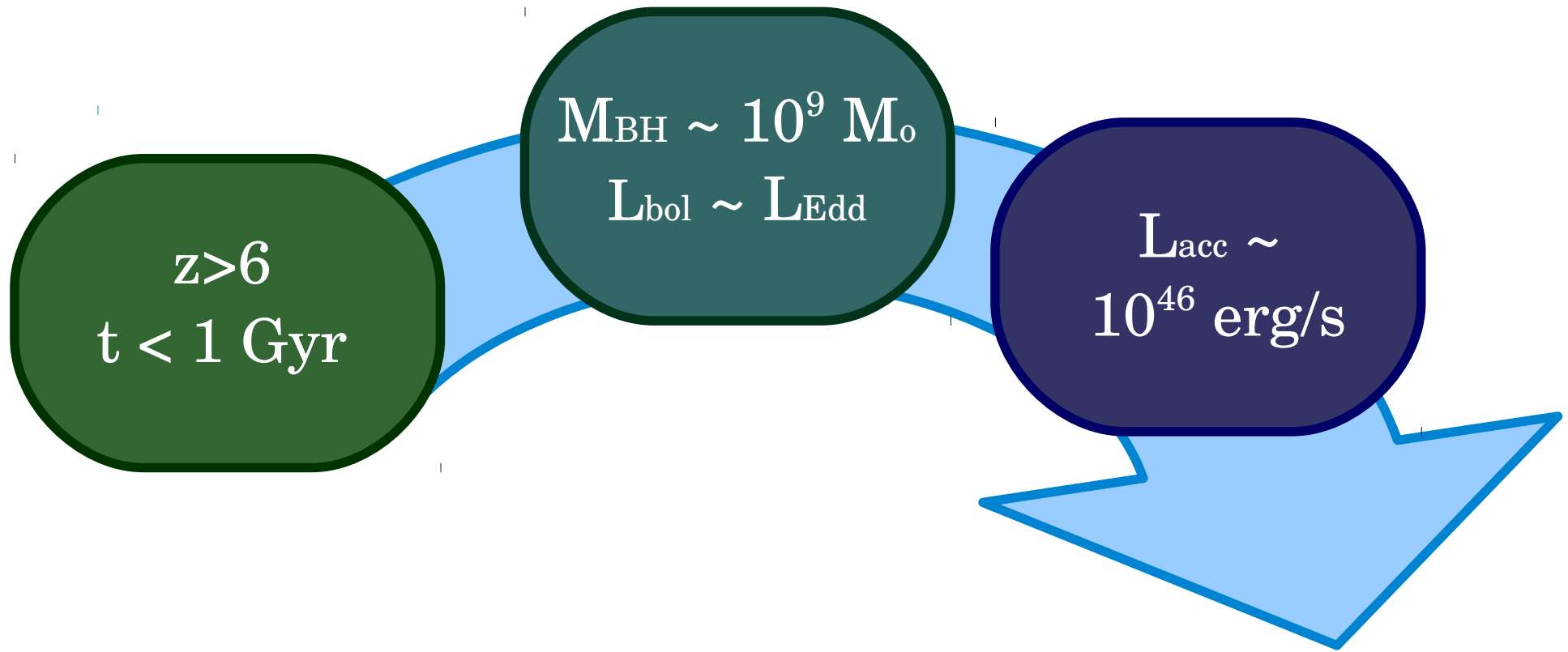
Yet again... $z > 6$ QSOs!

$z > 6$
 $t < 1 \text{ Gyr}$

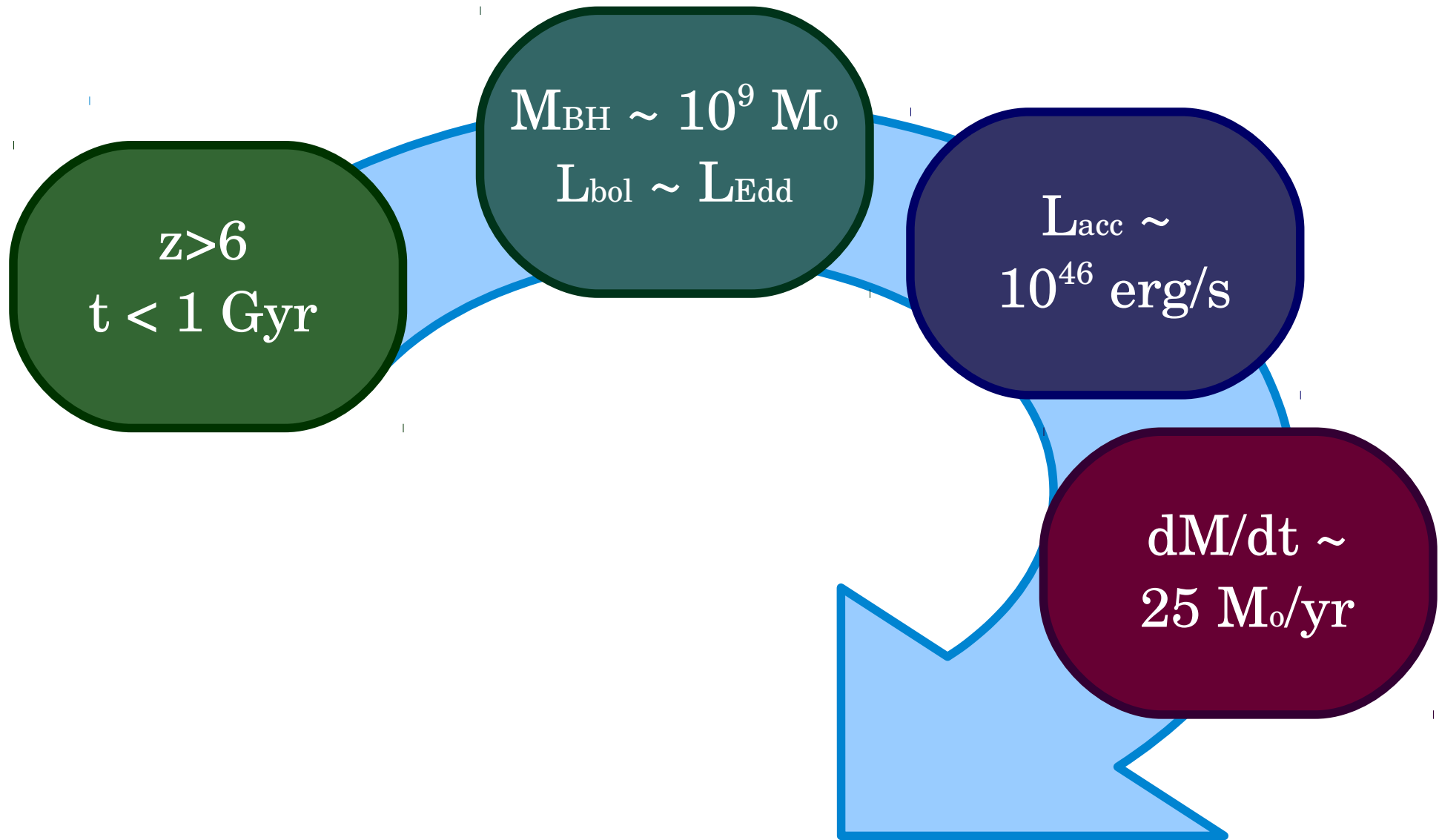
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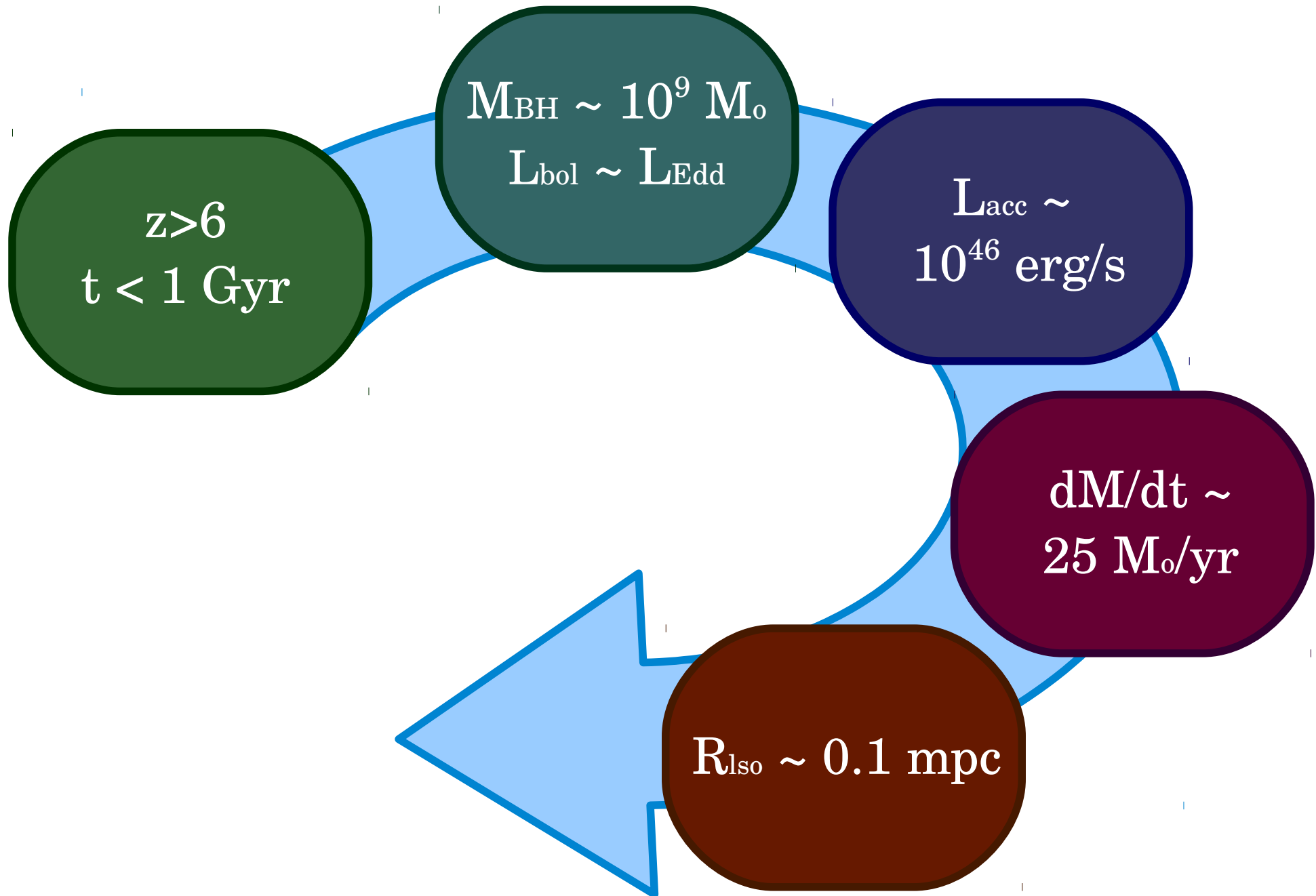
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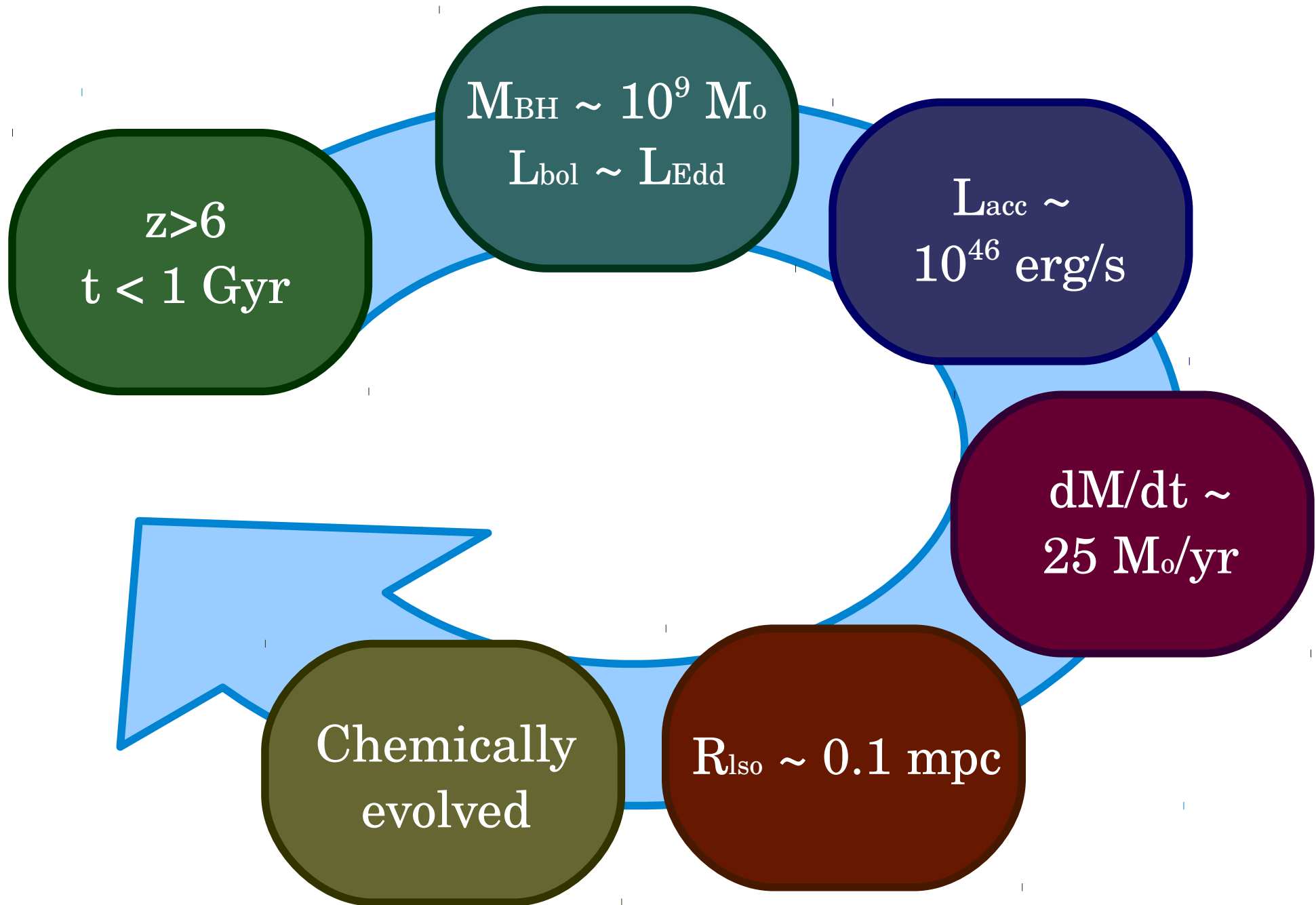
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The gas cycle

Broad Lines
(talks Mon/Tue)



0.1-1 pc



Distance

The gas cycle

Absorption

(talks on Mon/Tue)

IGM

Broad Lines

(talks Mon/Tue)

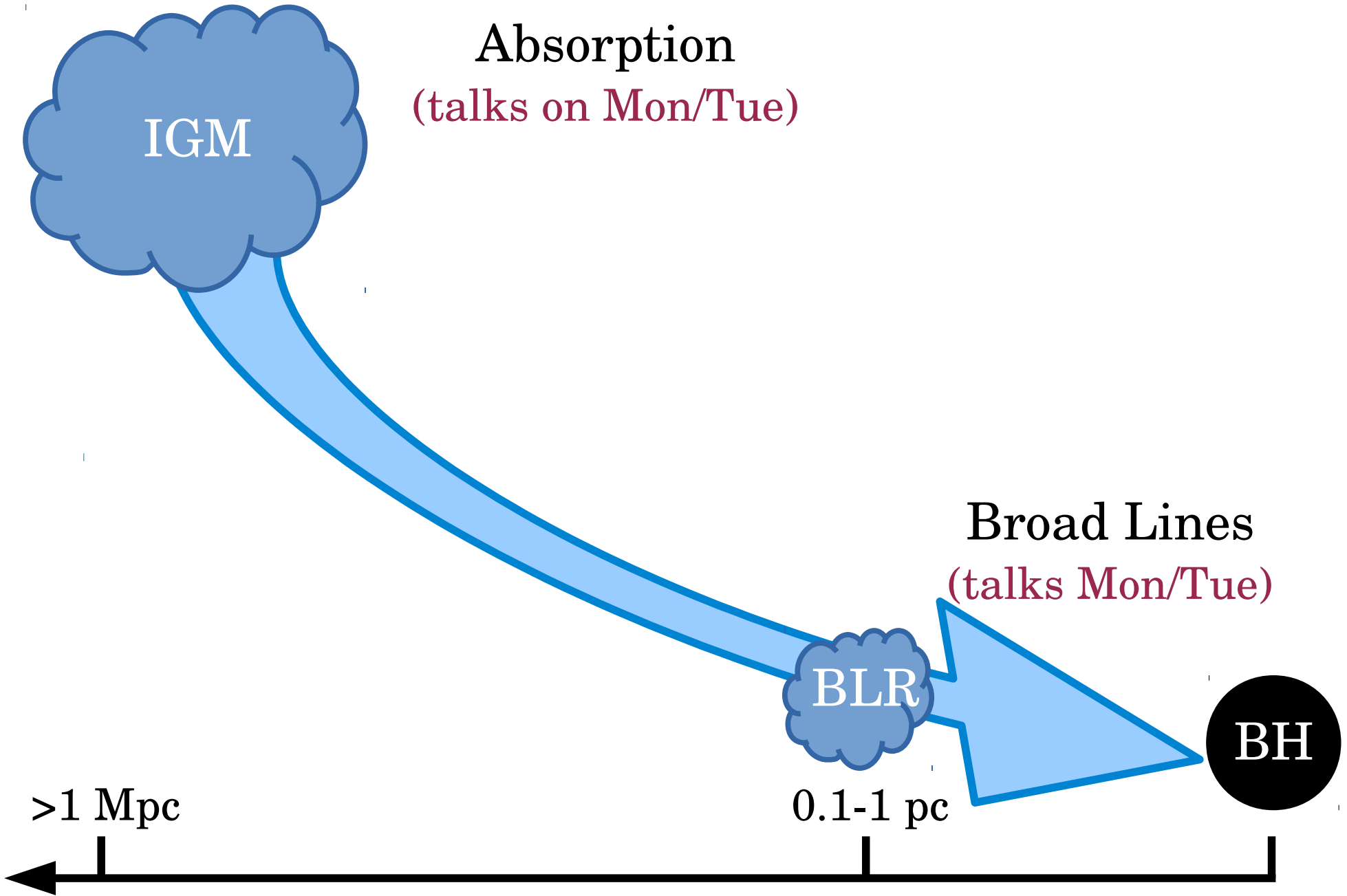
BLR

BH

>1 Mpc

0.1-1 pc

Distance

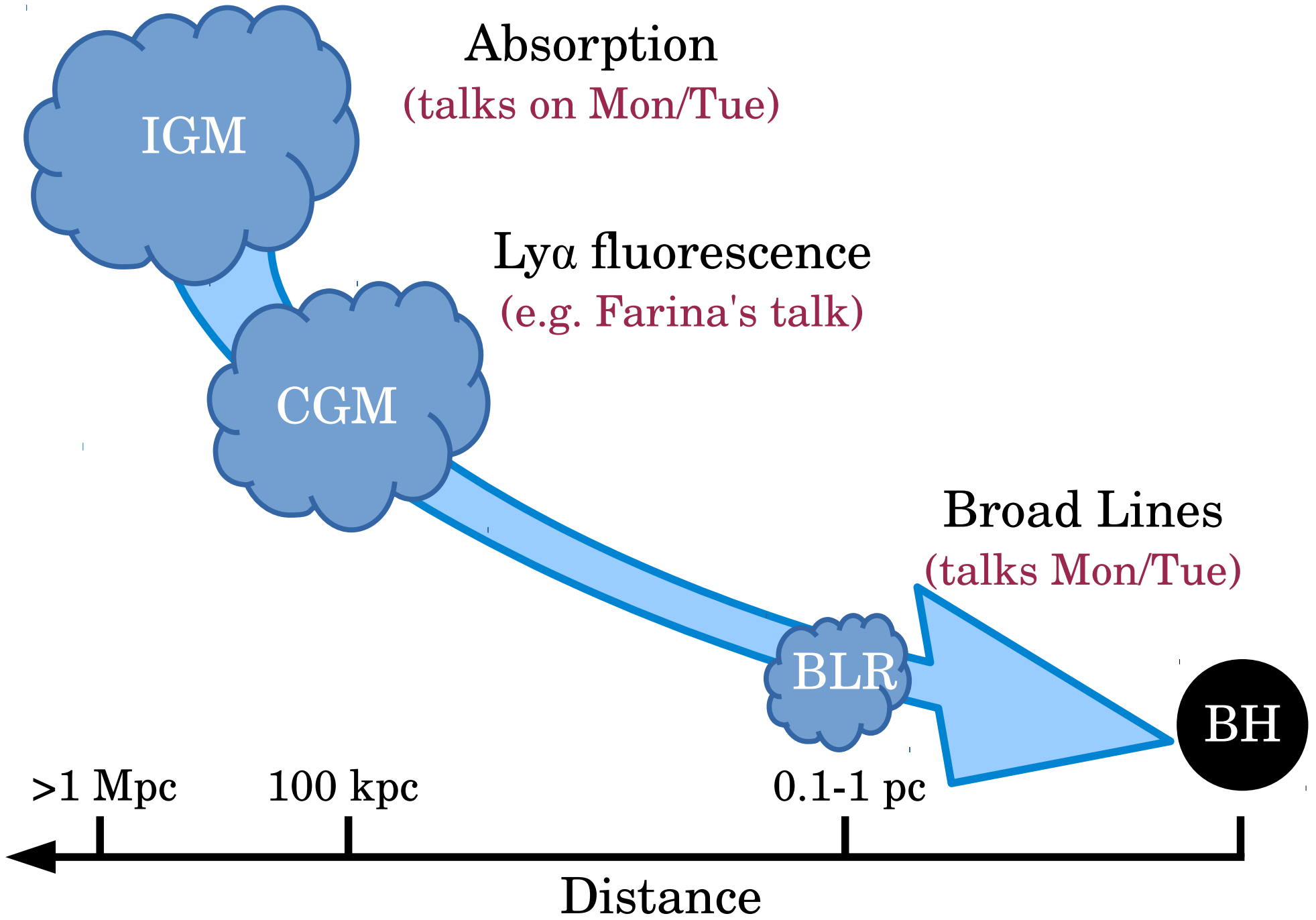


The gas cycle

Absorption
(talks on Mon/Tue)

$\text{Ly}\alpha$ fluorescence
(e.g. Farina's talk)

Broad Lines
(talks Mon/Tue)



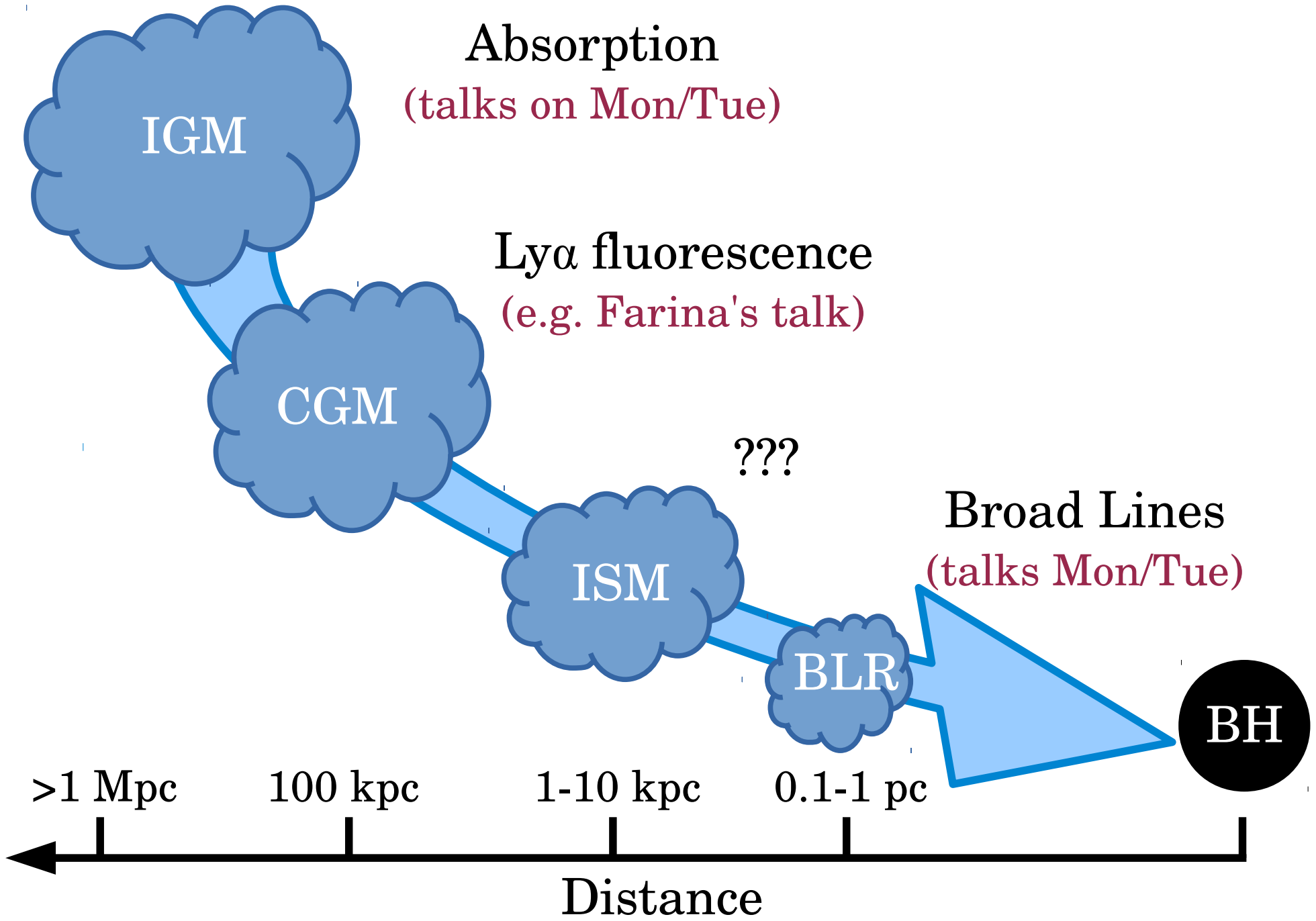
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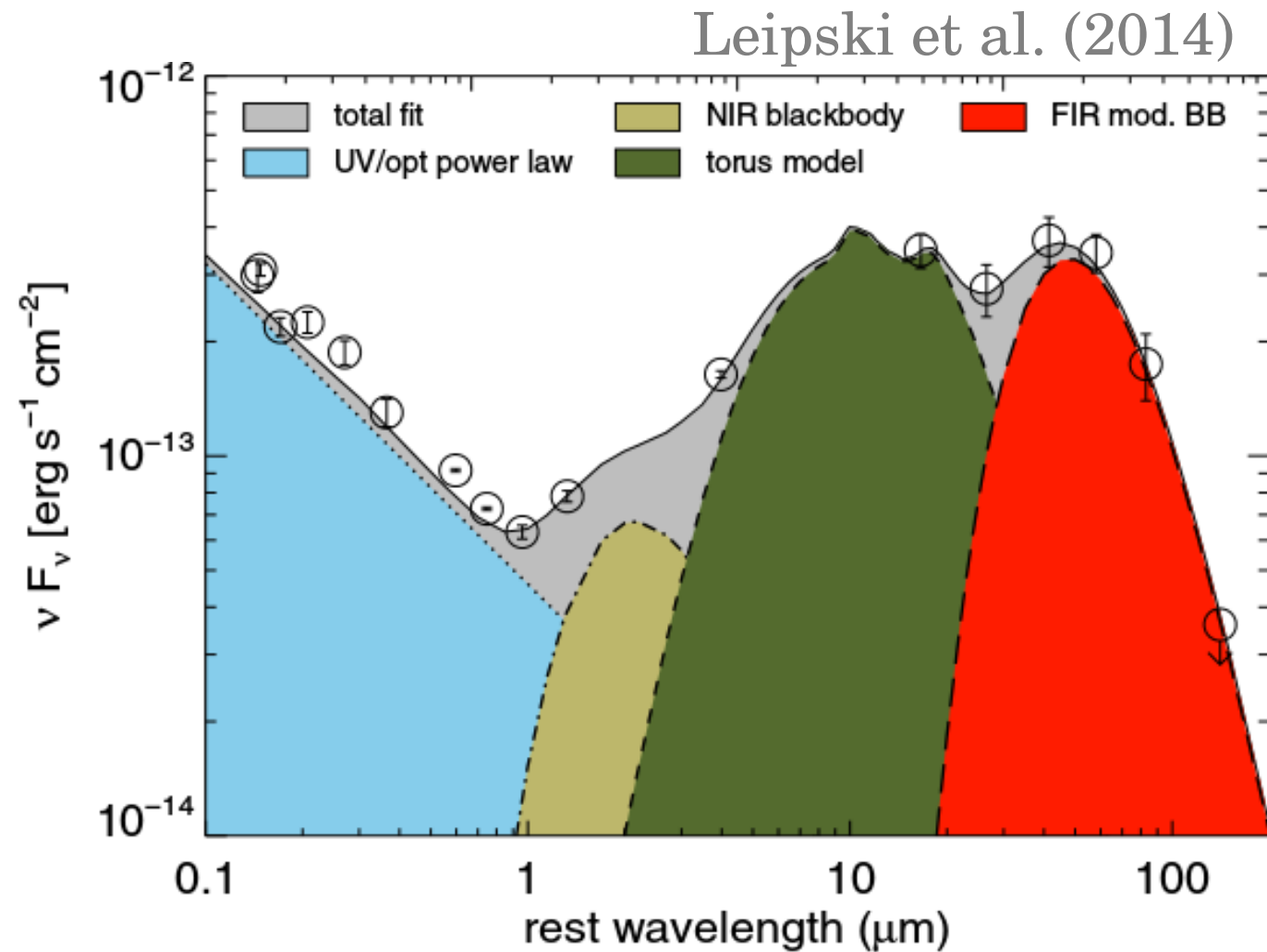
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(e.g. Farina's talk)

???

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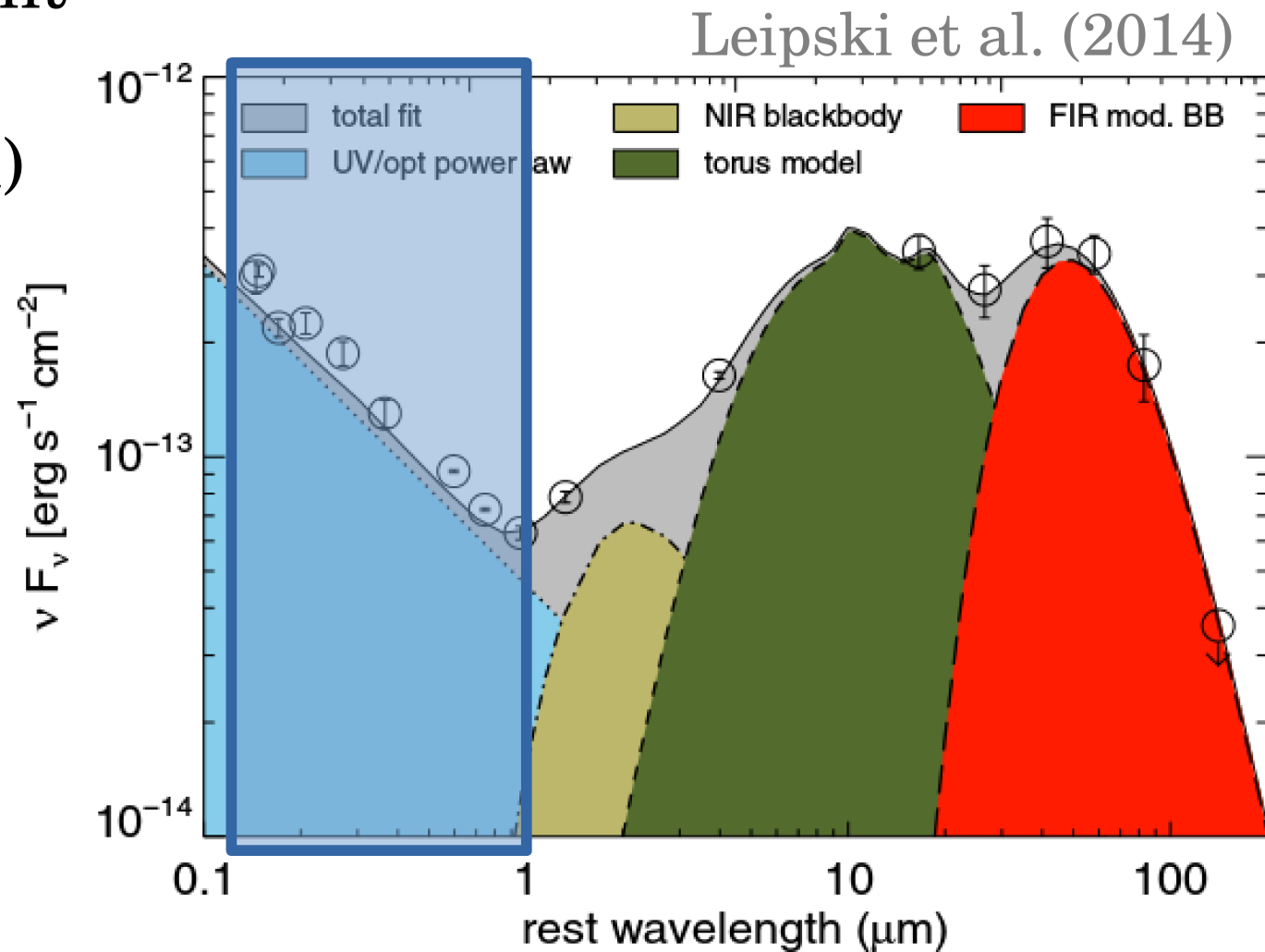


Exposing the build-up of the host galaxies



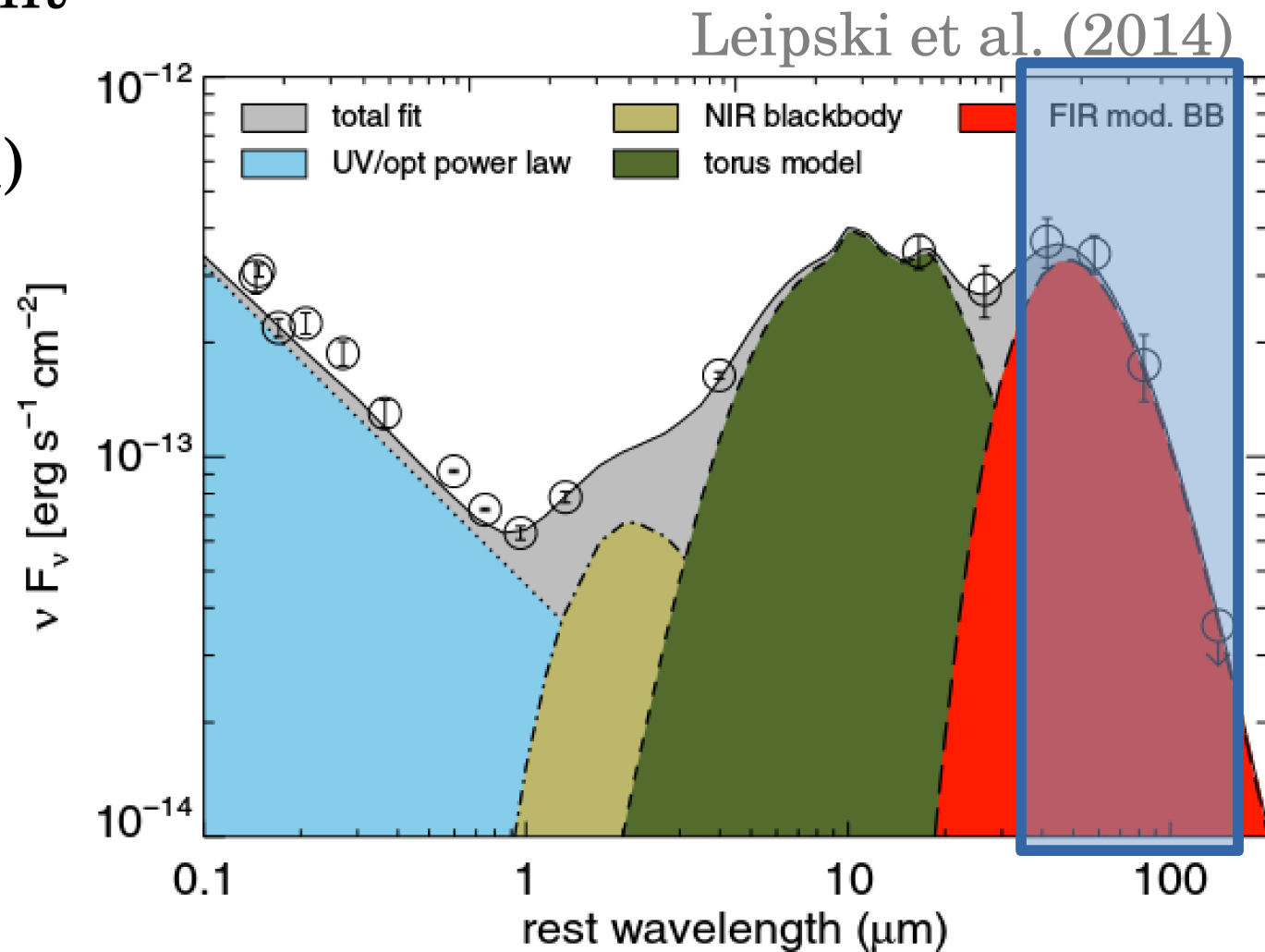
Exposing the build-up of the host galaxies

The nucleus
outshines starlight
from the host
galaxies (>3 dex)



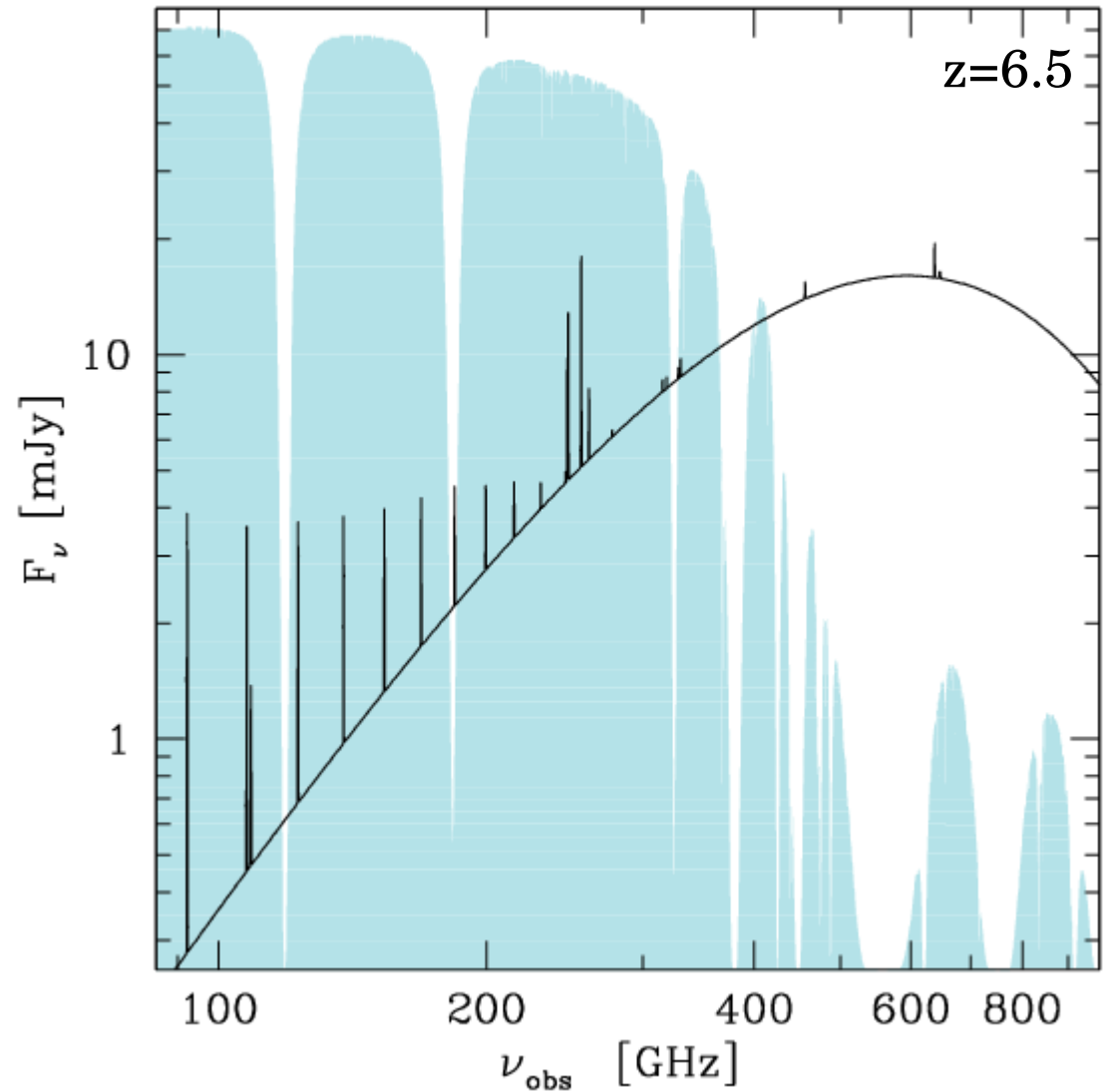
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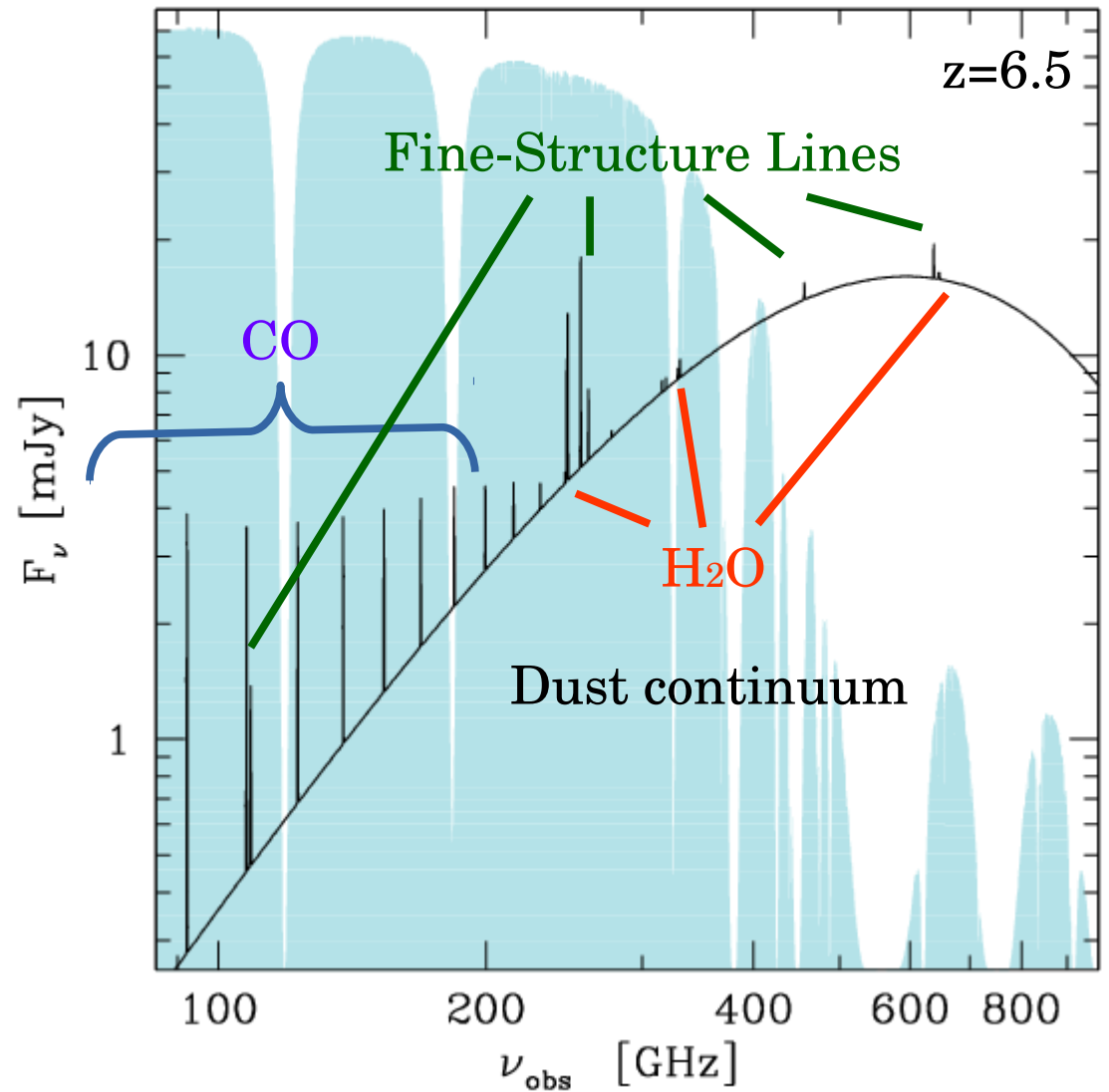
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FIR (= obs [sub-]mm)
is the way to go



Exposing the build-up of the host galaxies

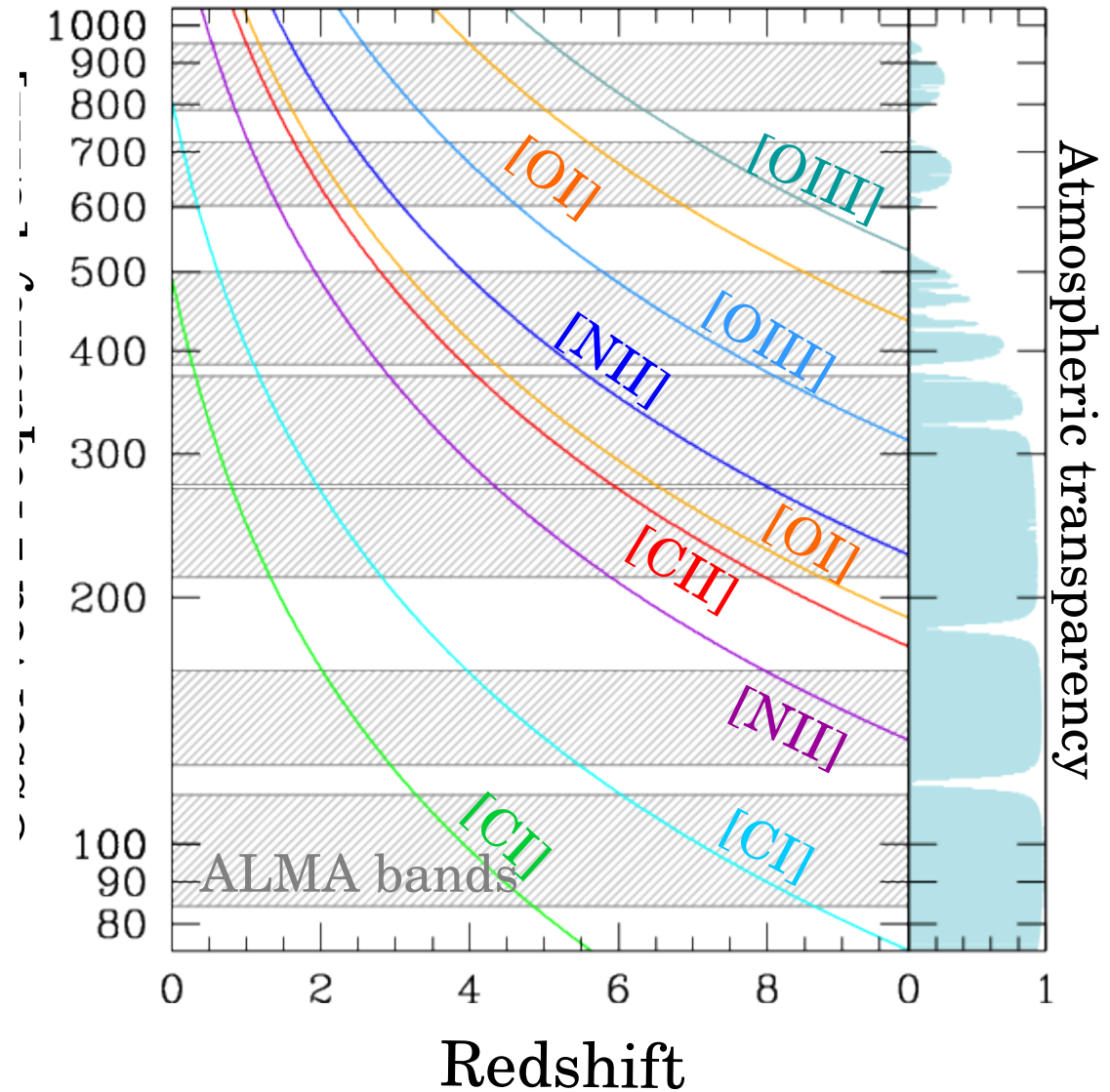
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Exposing the build-up of the host galaxies

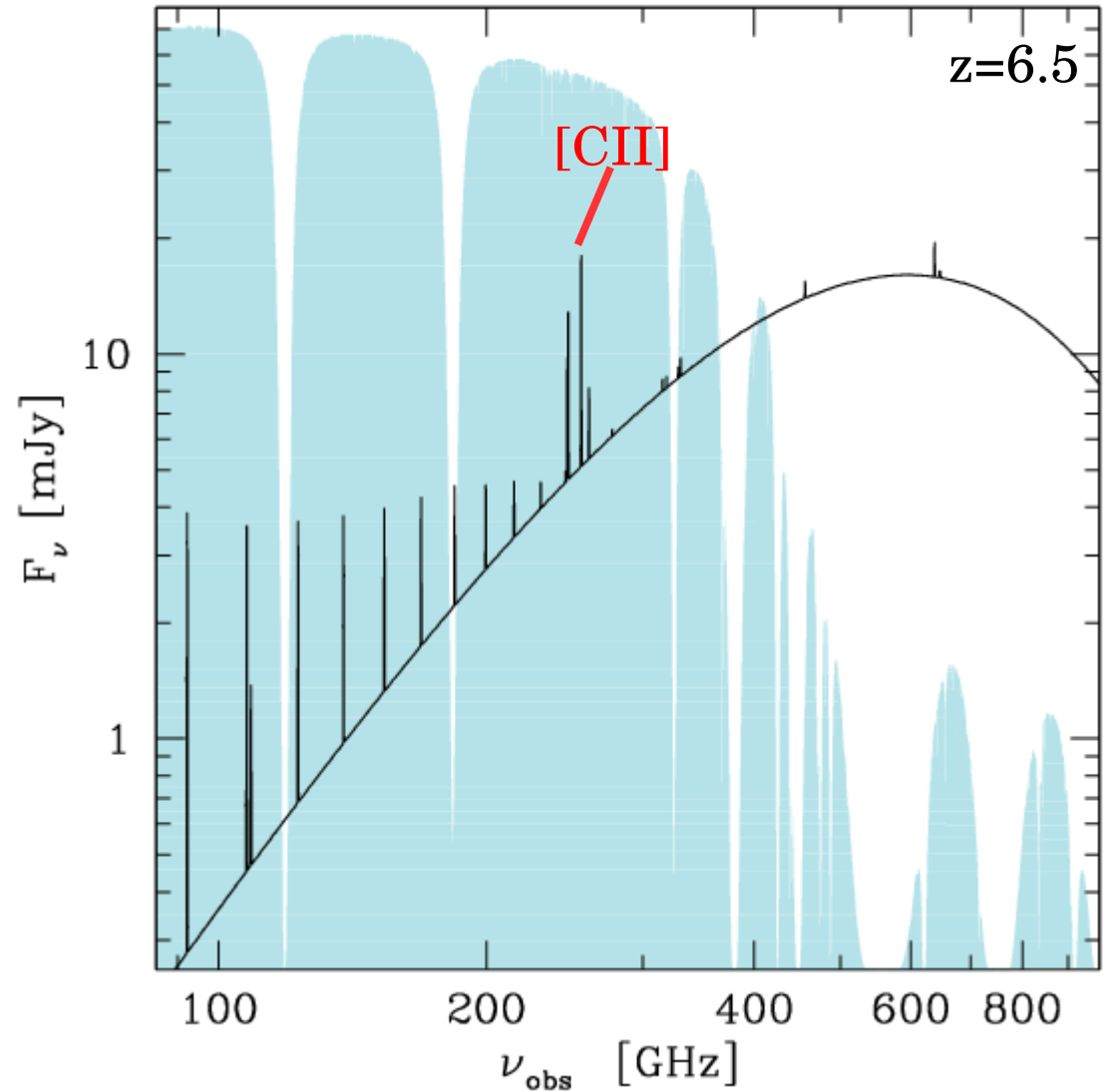
Fine-Structure Lines:

- main coolant of the cold ISM
- accessible at [sub-]mm wavelengths



Exposing the build-up of the host galaxies

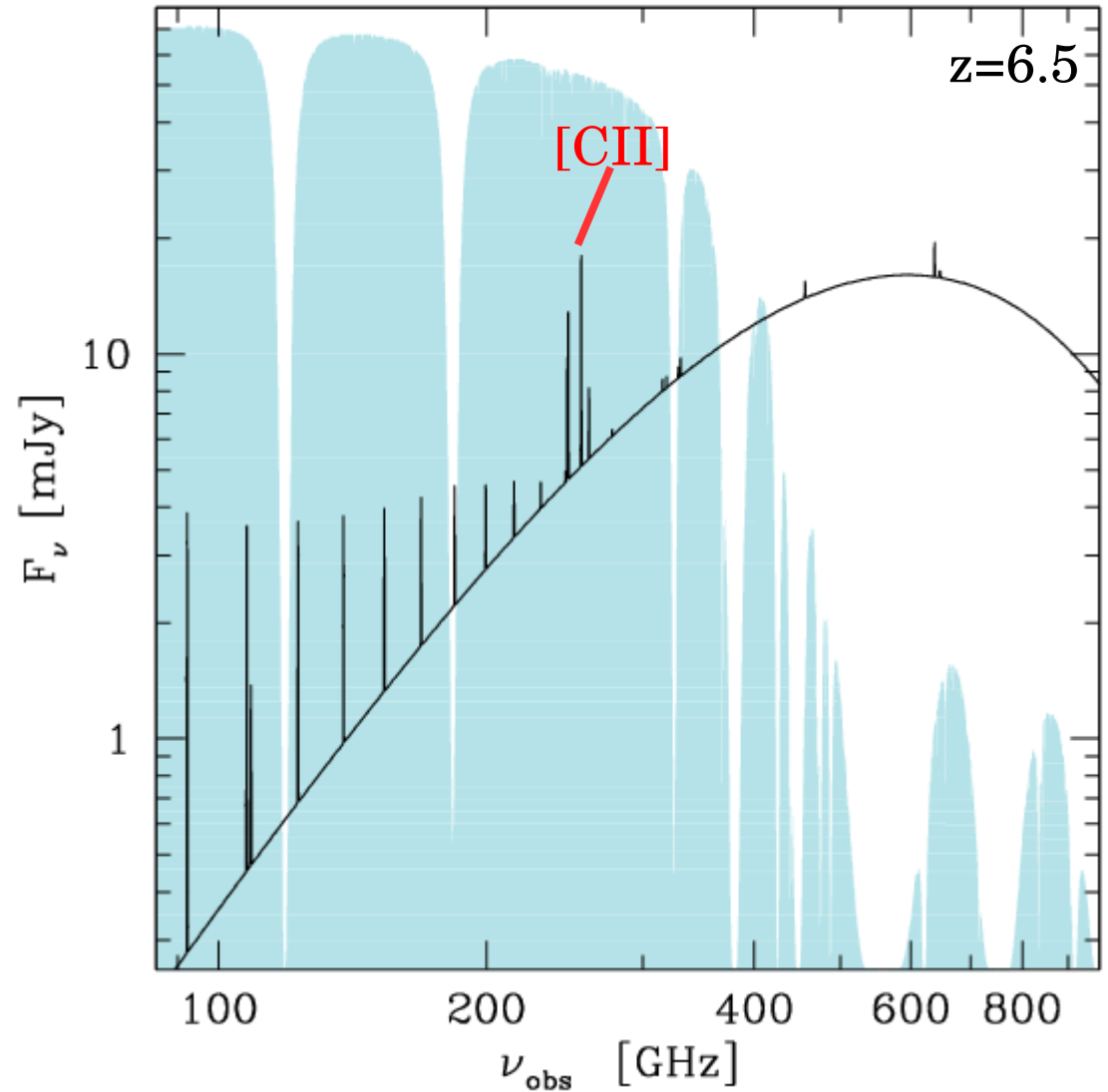
[CII] 158 μm :
carries $\sim 0.3\%$ of the
ENTIRE luminosity
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Exposing the build-up of the host galaxies

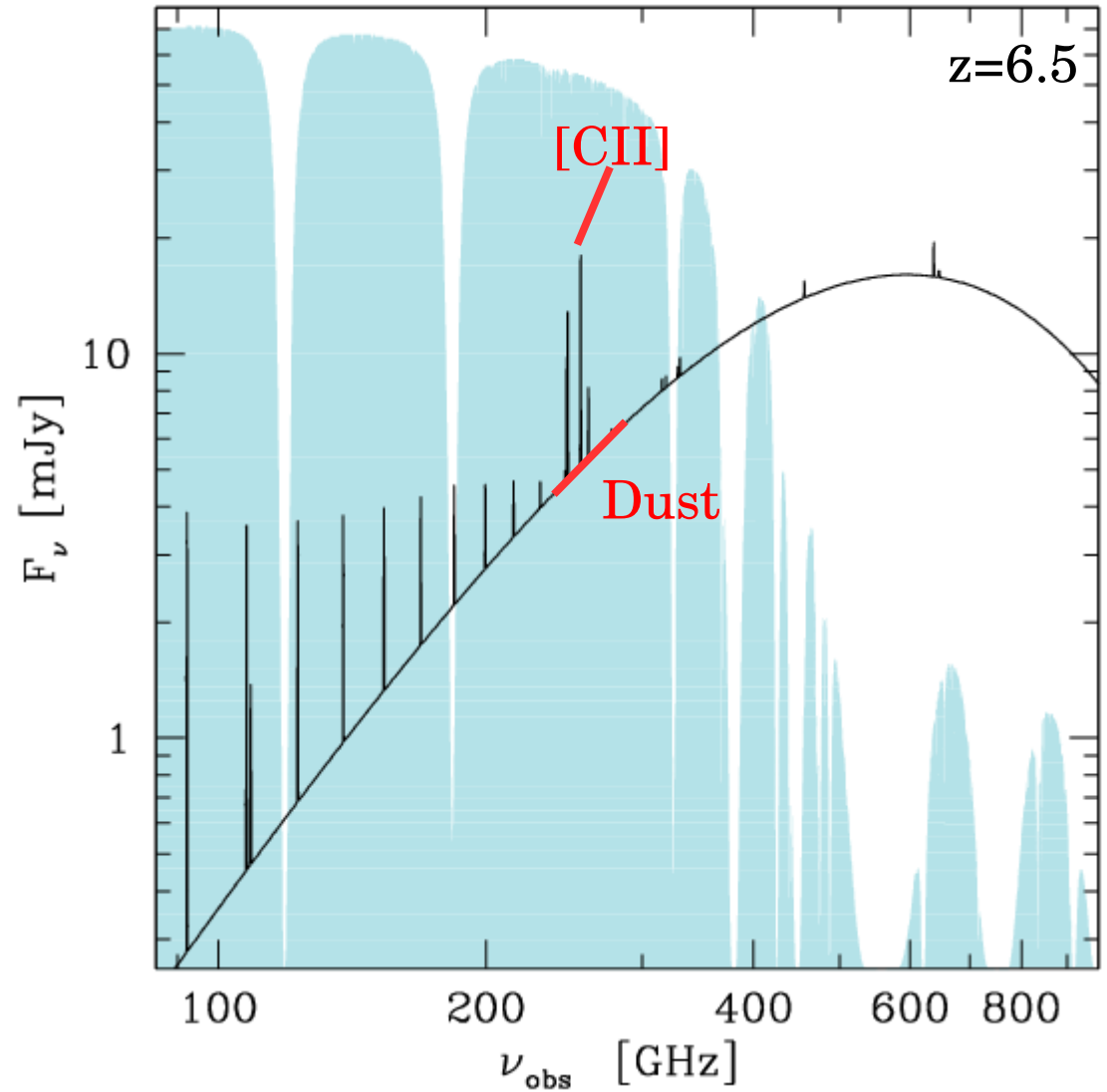
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Ideal for redshift /
dynamical
measurements at
high-z



Exposing the build-up of the host galaxies

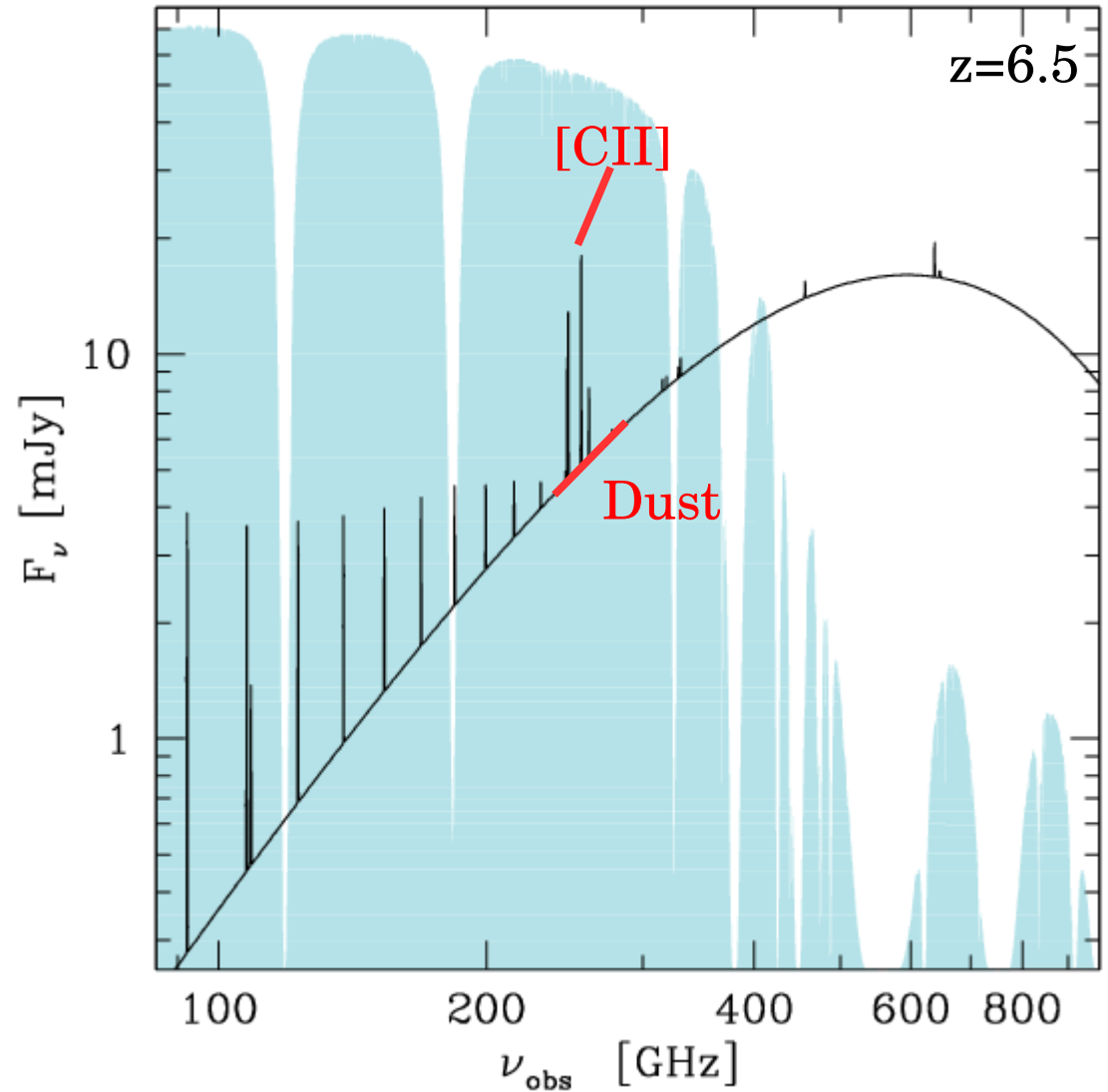
[CII] obs give dust RJ
continuum for free!



Exposing the build-up of the host galaxies

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$$RJ \sim M_{\text{dust}}$$



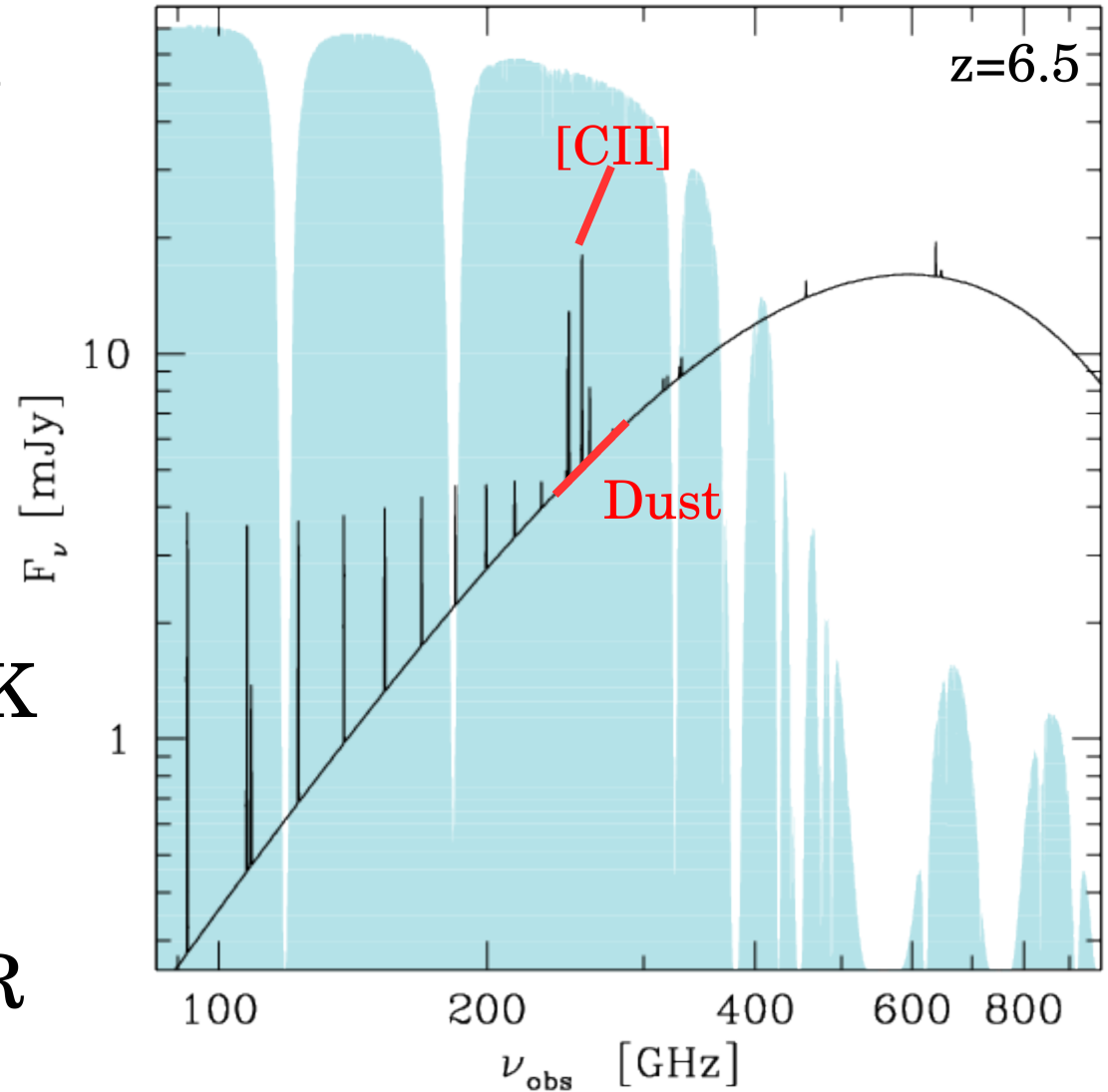
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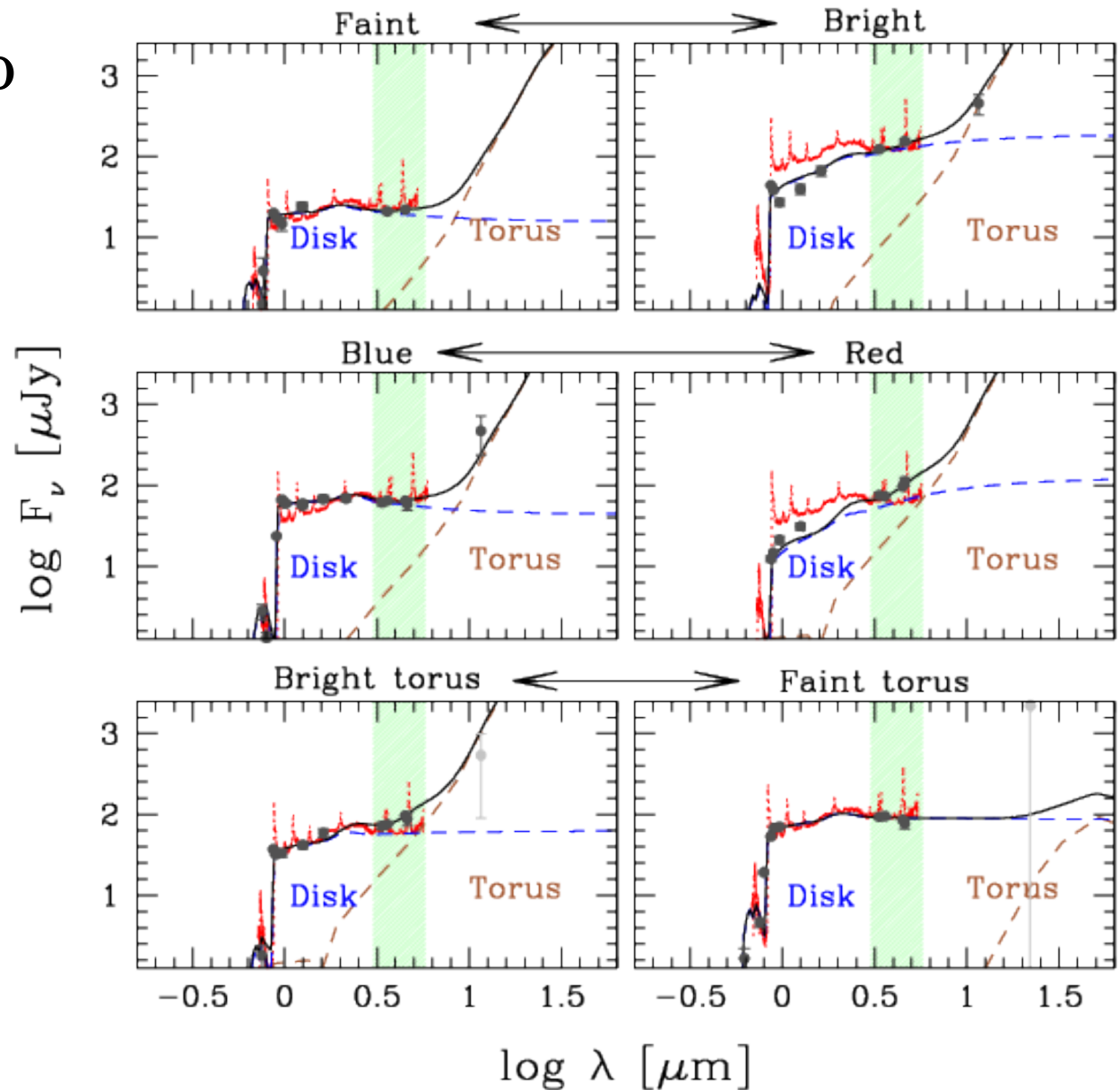
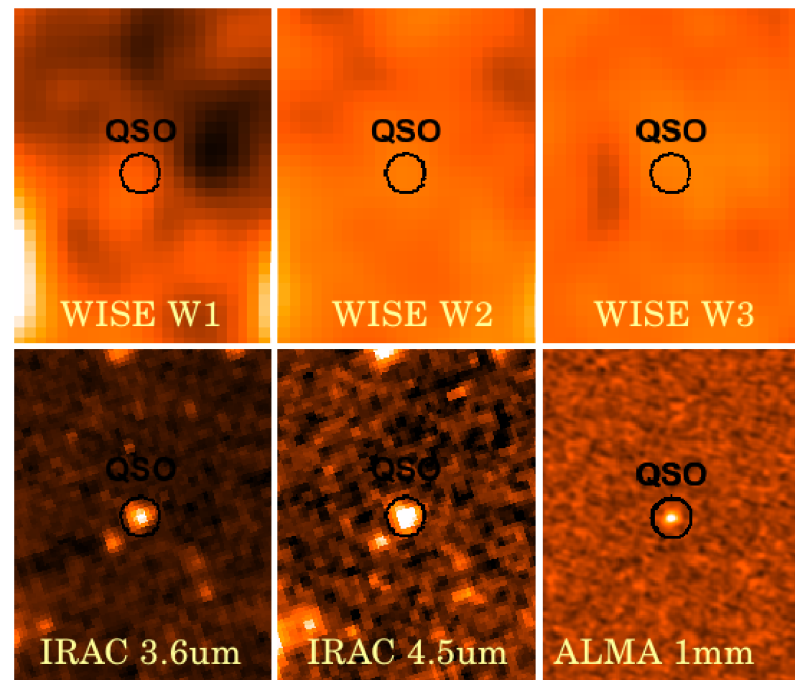
If we assume $T_{\text{dust}} \sim 47$ K
(Beelen+06, Leipski+14):

$$F_{\nu}(\text{dust}) \Rightarrow L_{\text{IR}} \Rightarrow \text{SFR}$$



Exposing the build-up of the host galaxies

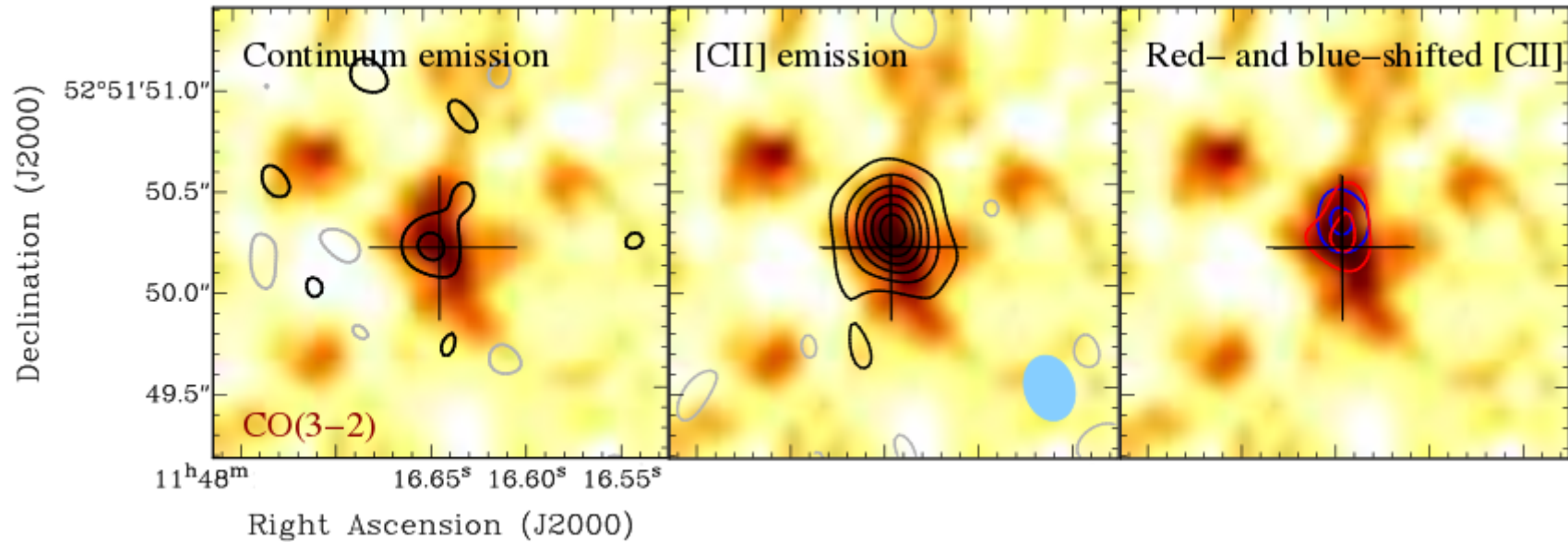
Dust constraints also
from Opt/NIR/MIR
SED fits



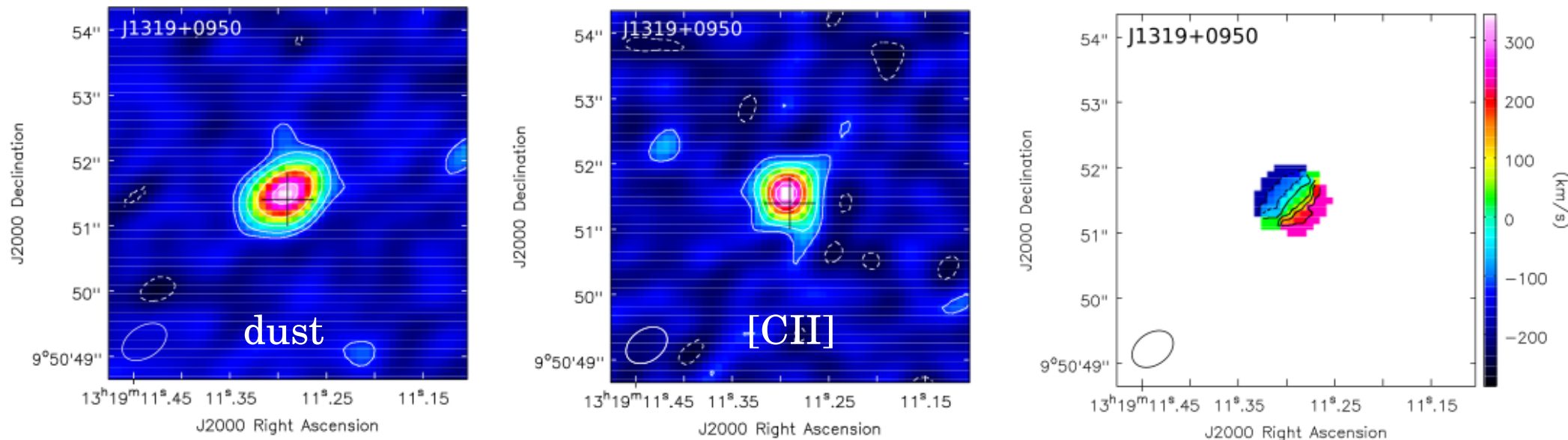
Decarli et al. (in prep)

[CII] in high-z QSOs

Walter et al. (2009)



Wang et al. (2013)



[CII] in high-z QSOs

Maiolino+05, Walter+09 (1)

Venemans+12 (1)

Willott+13 (2)

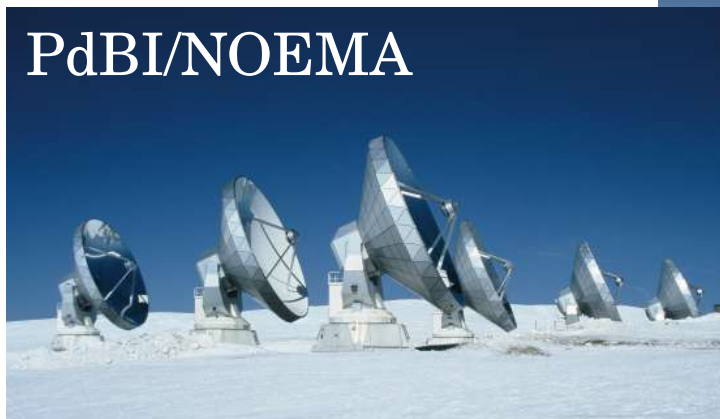
Banados+15 (1)

Decarli+ in prep (2)

IRAM 30m



PdBI/NOEMA



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ALMA



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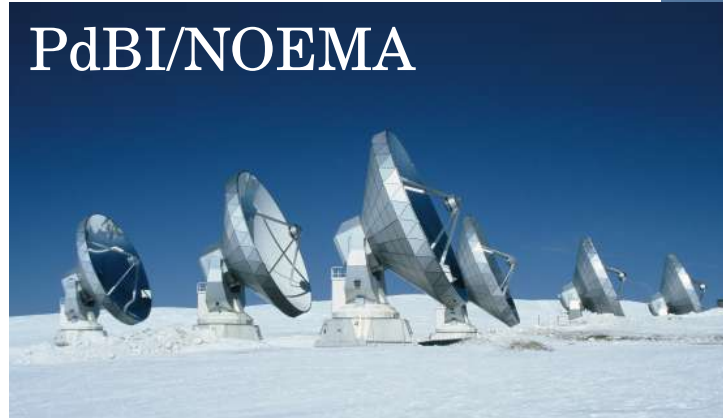
Willott+15 (2)

[CII] survey in $z > 6$ QSOs (35)

IRAM 30m



PdBI/NOEMA



ALMA



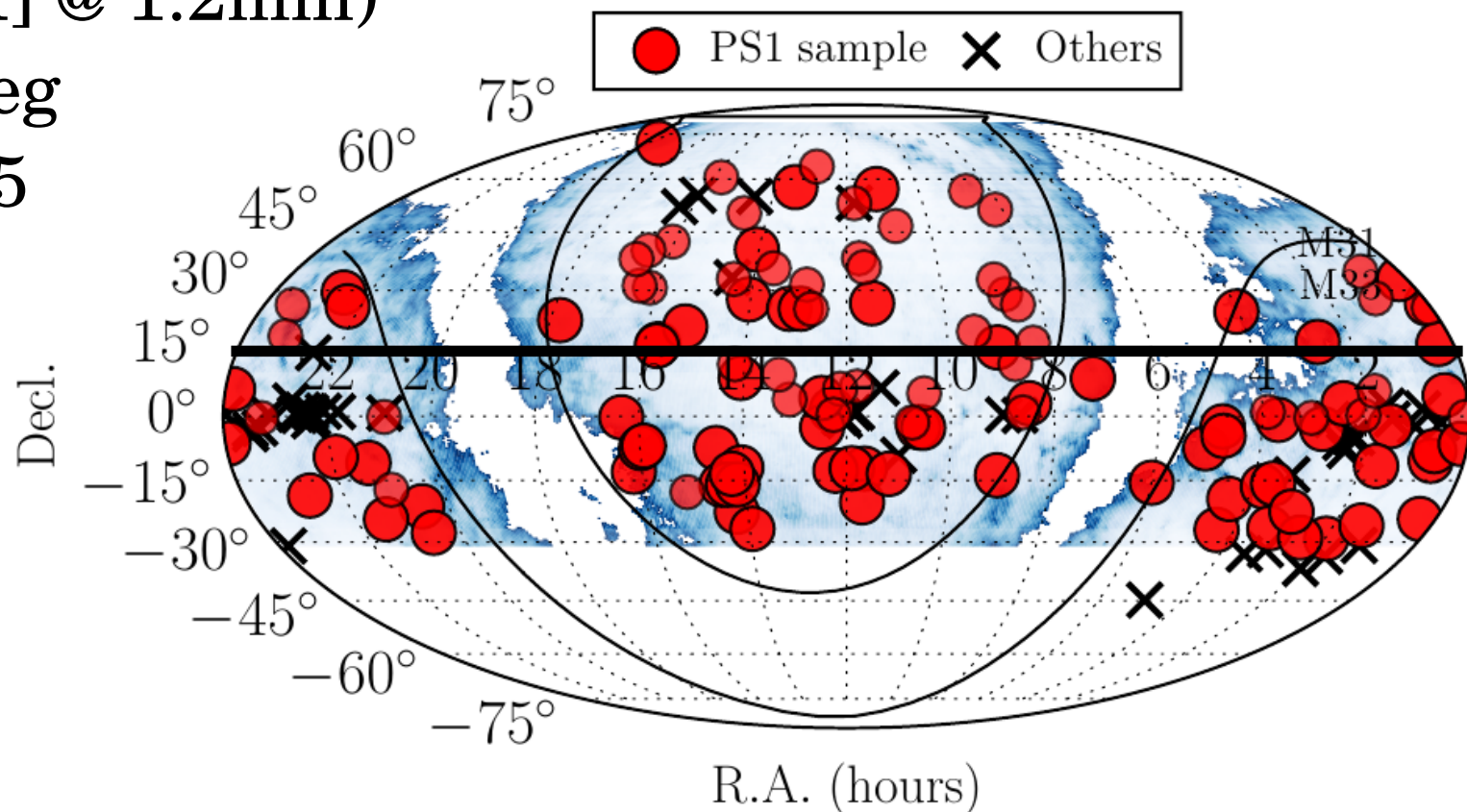
Survey design

$z > 5.94$ ([CII] @ 1.2mm)

Dec $< +15$ deg

$M_{1450} < -25.25$

PI: Walter



Banados et al. (subm)

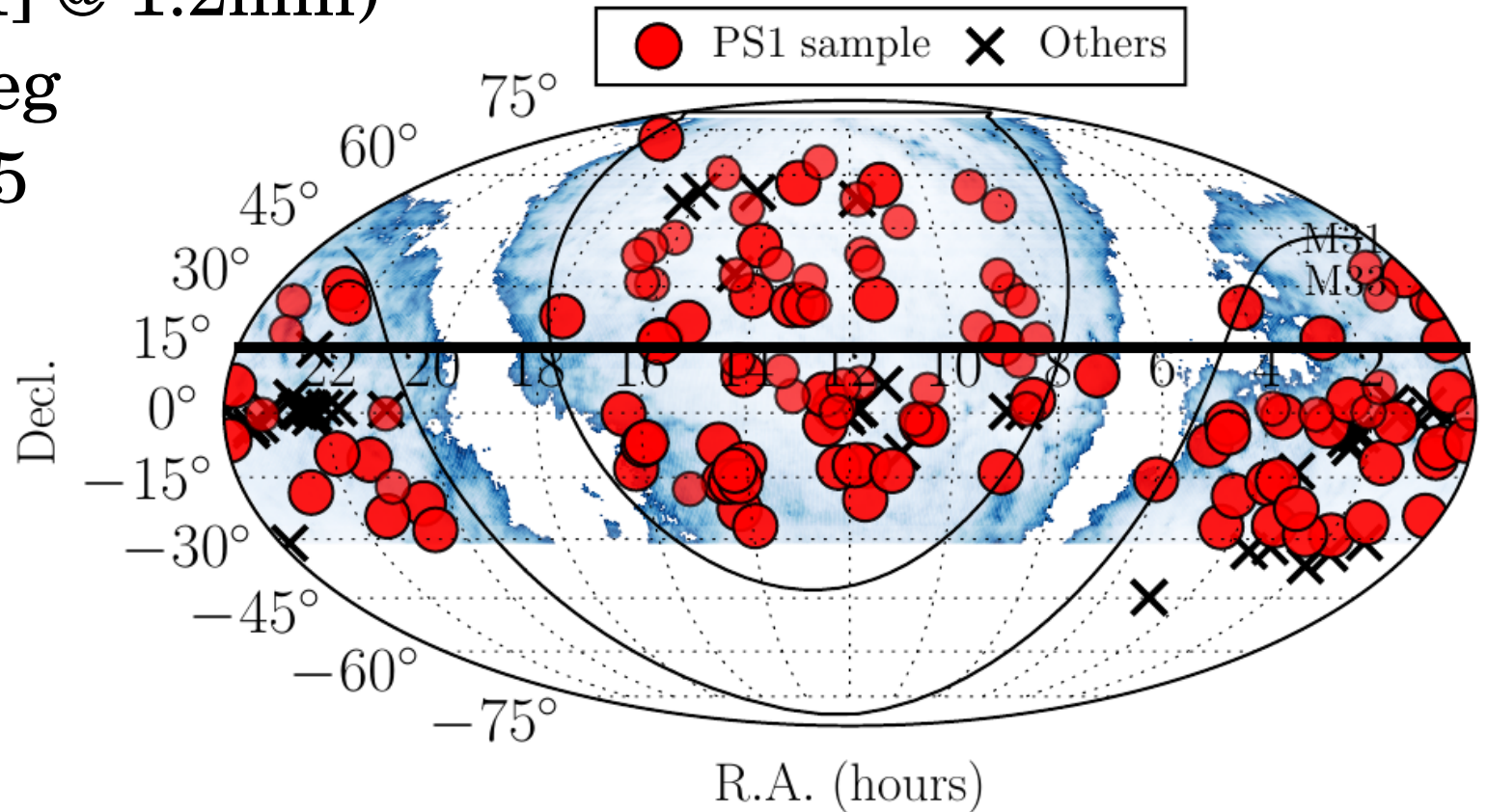
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8 min on source

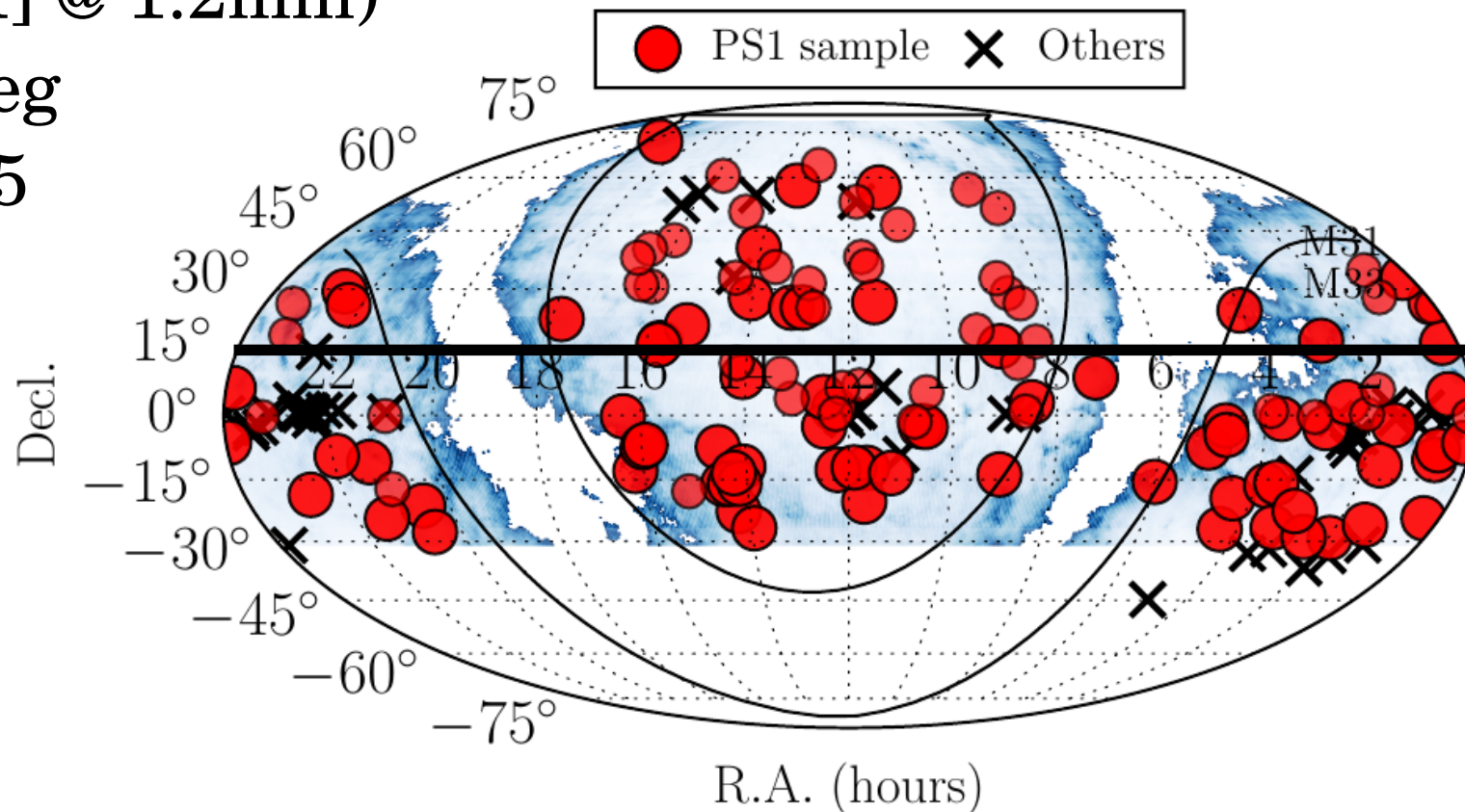
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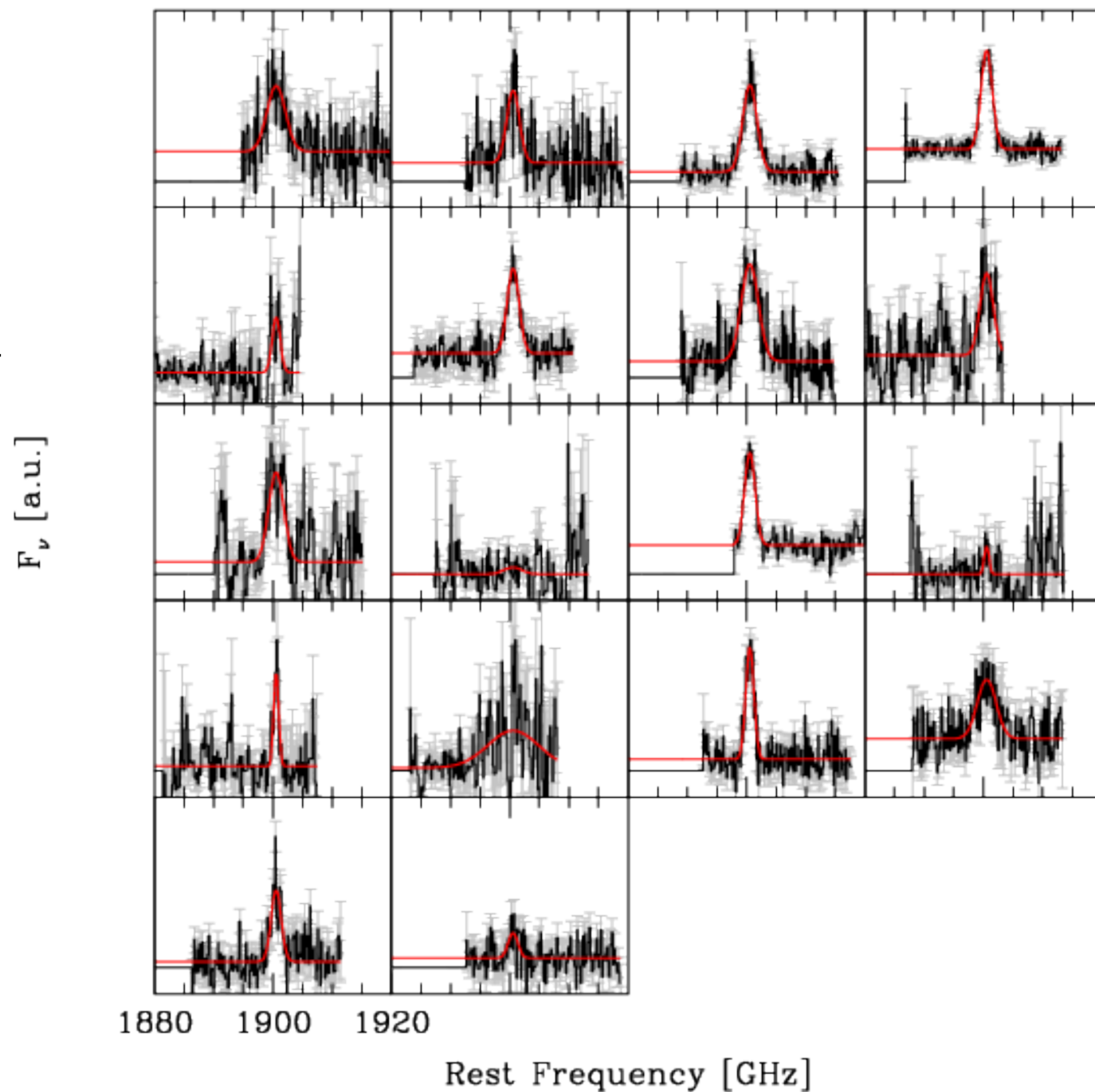
8 min on source

18 (+2) delivered so far

[CII] survey in $z > 6$ QSOs

15 clear [CII]
detections (83%)

16 (89%) detected
in the continuum
(at various
significance)

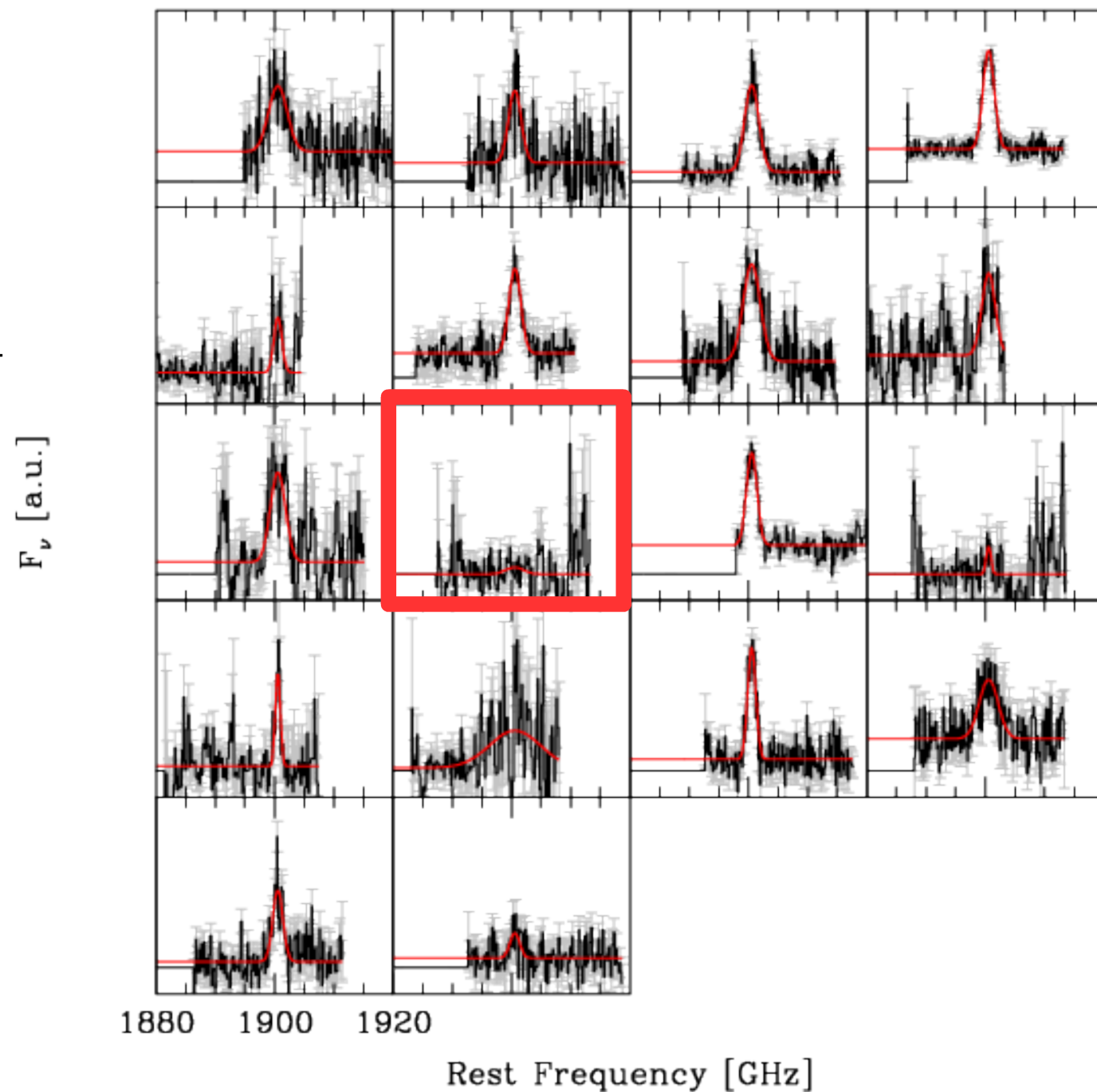


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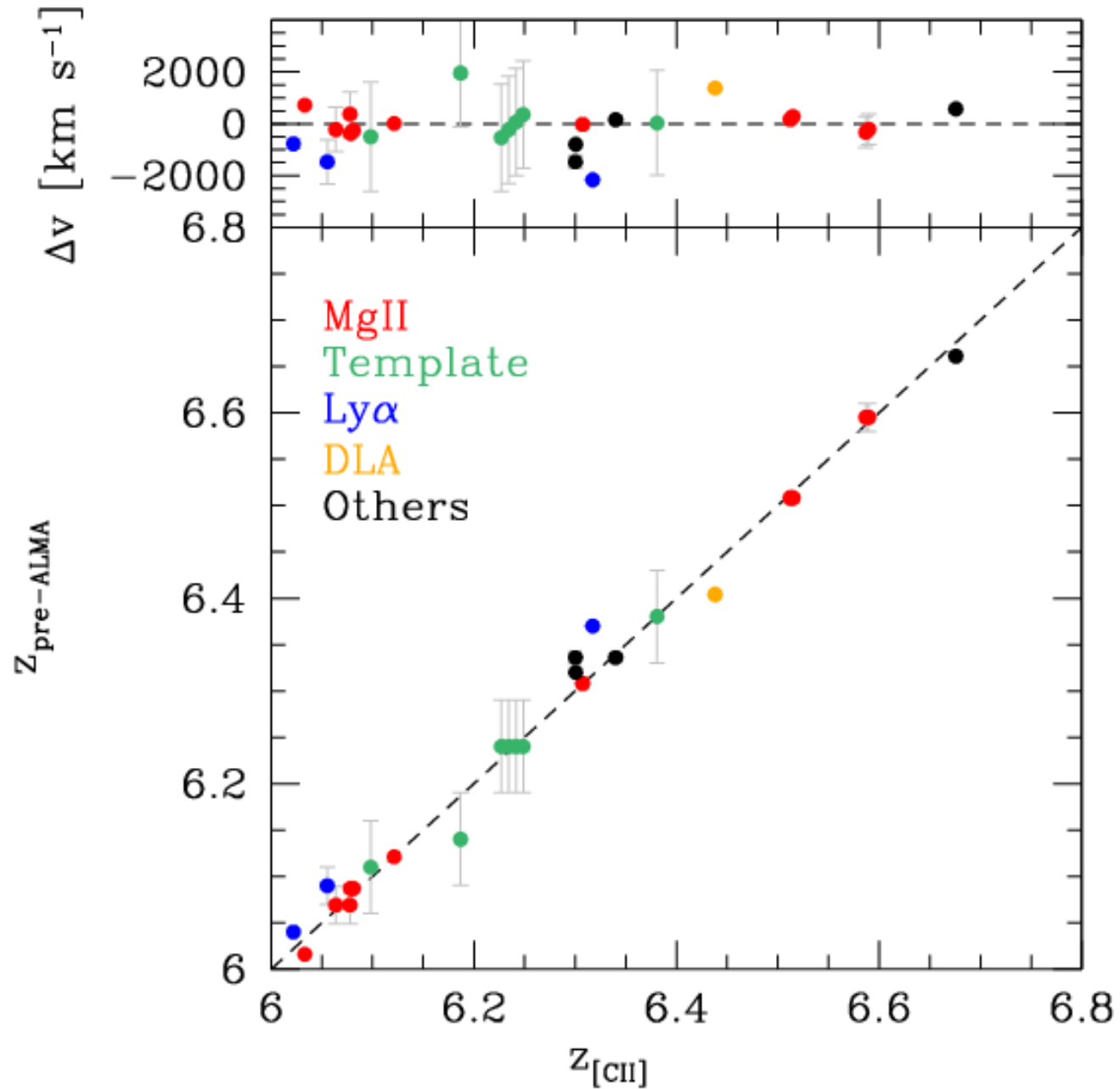
Laura's favorite
J1030+0524 is
NOT detected!



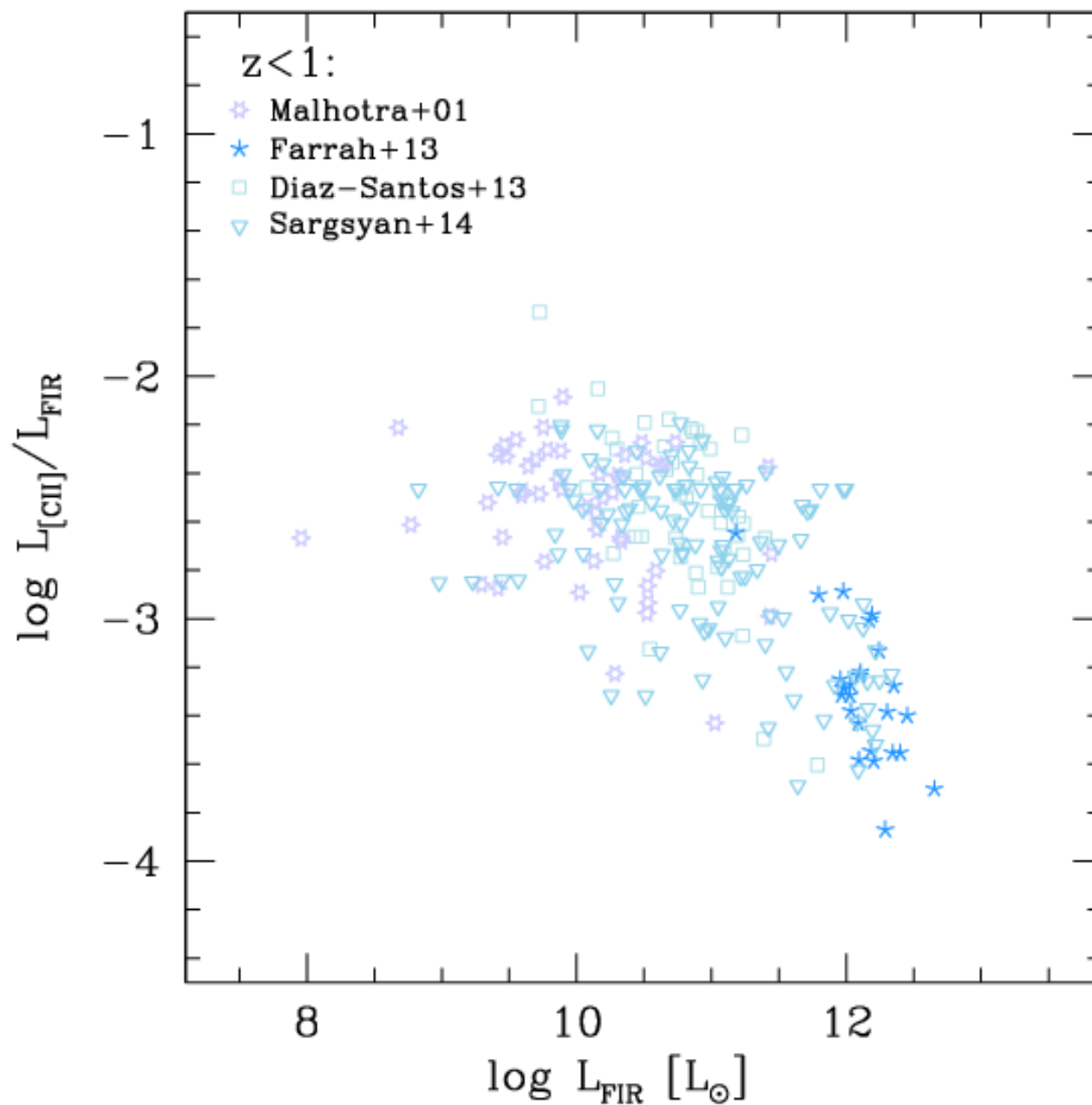
[CII] survey in $z > 6$ QSOs

MgII-based
redshifts:
 $|\Delta v| < 500$ km/s

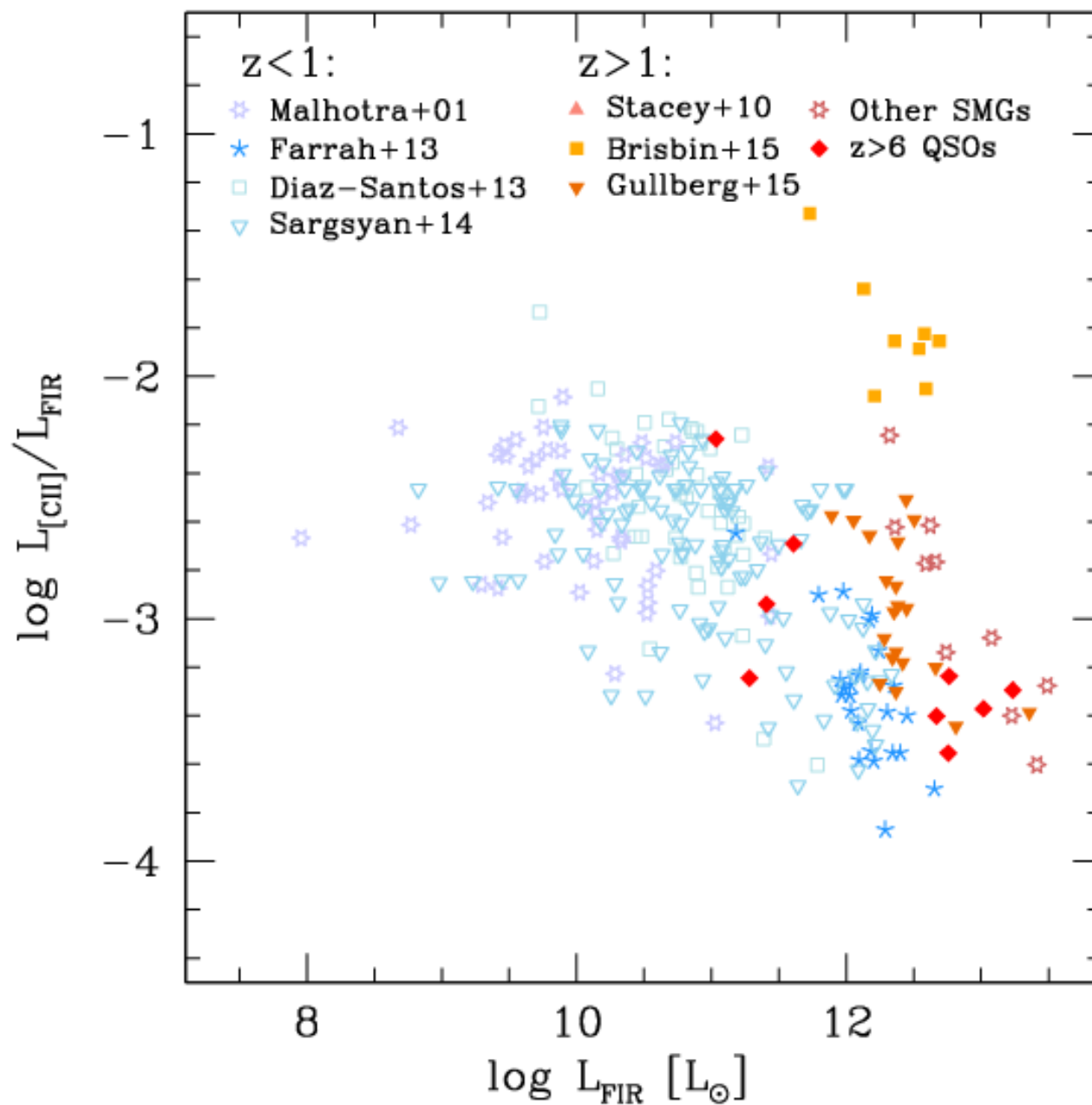
Ly α -based redshifts:
off by ~ 1000 km/s



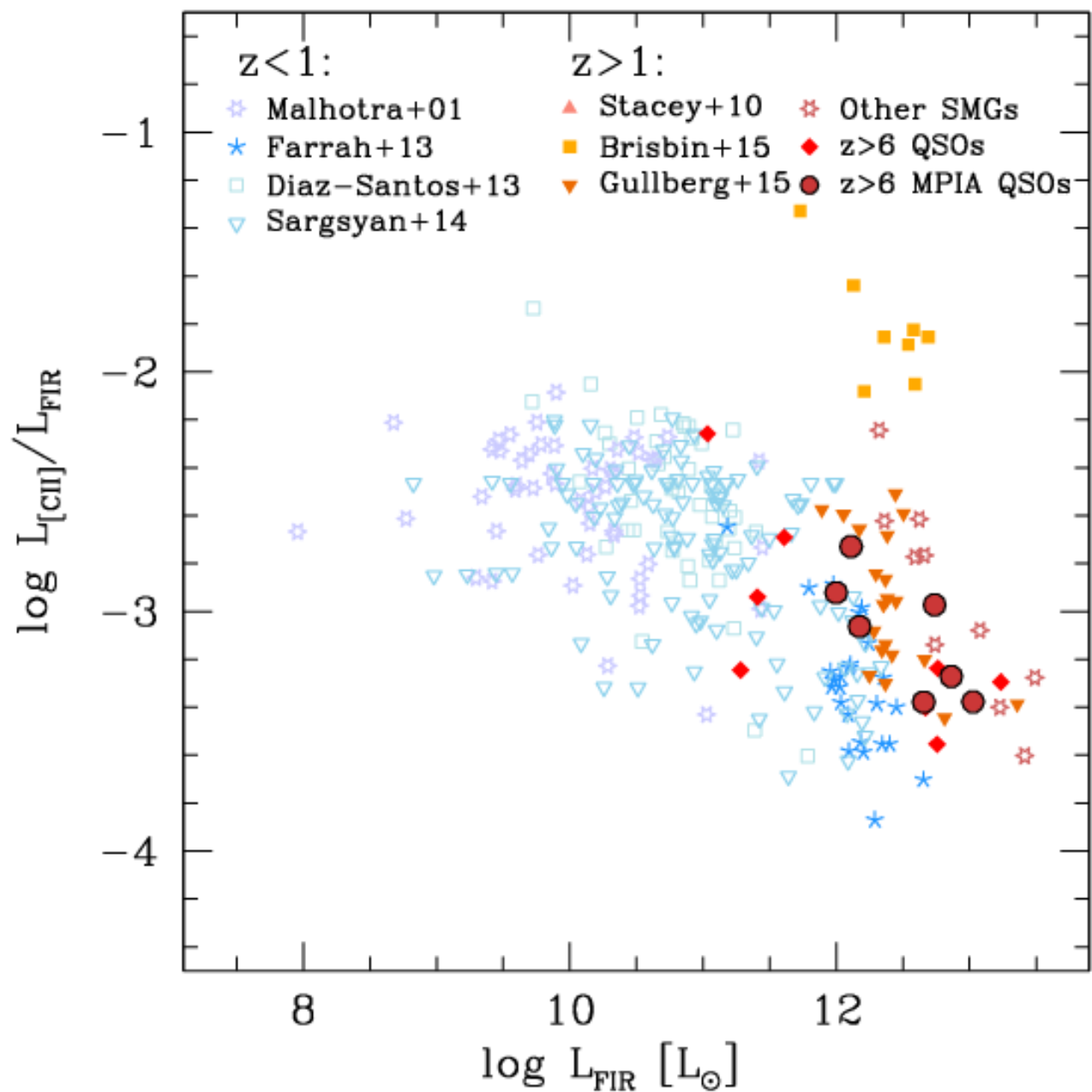
[CII] contribution to ISM cooling



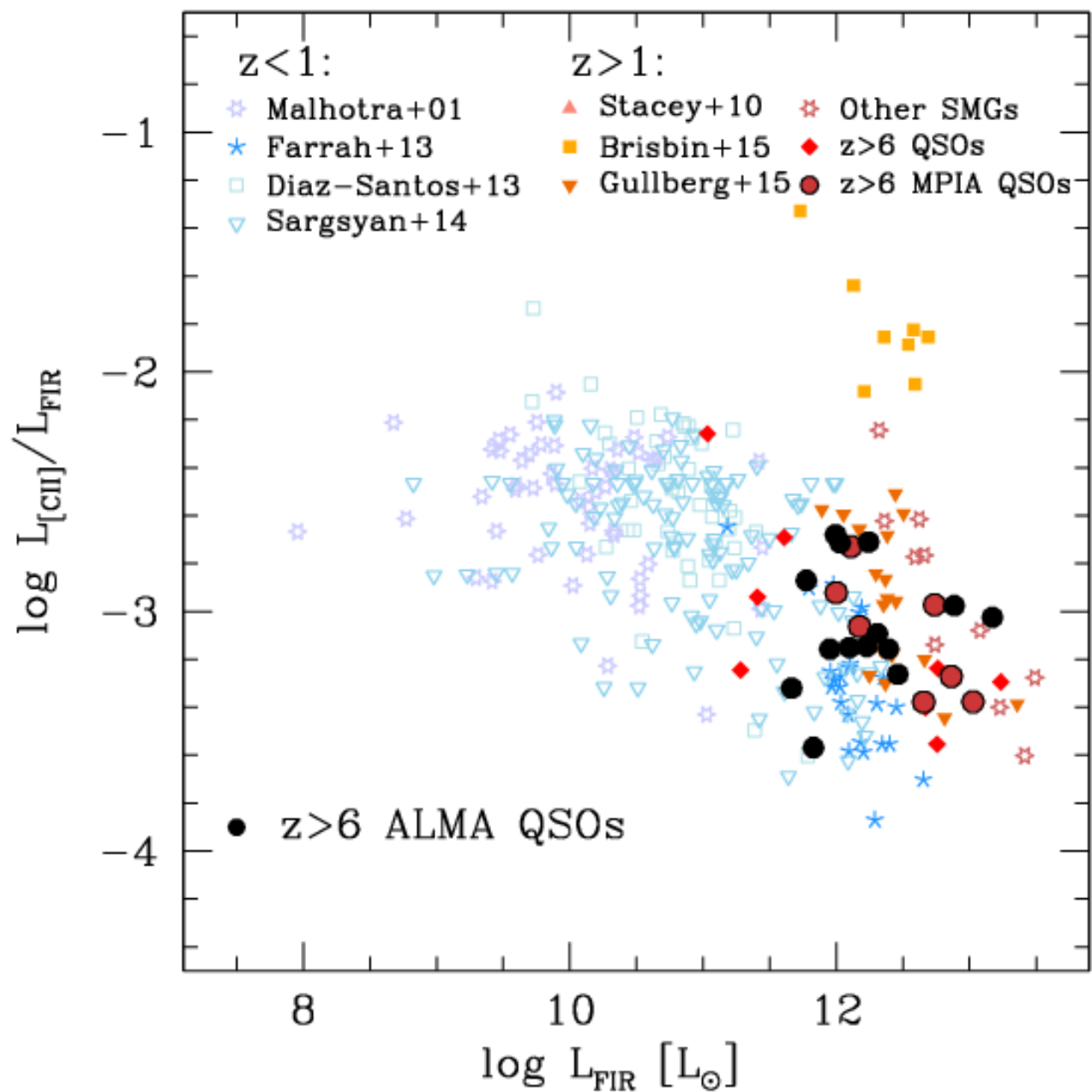
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Unveiling ISM physics with [CII]

In ionized gas:

$$[\text{CII}] \sim Z$$

$$[\text{NII}] \sim Z^2$$

$$[\text{CII}]/[\text{NII}] \sim 1/Z$$

(0th order, it's actually
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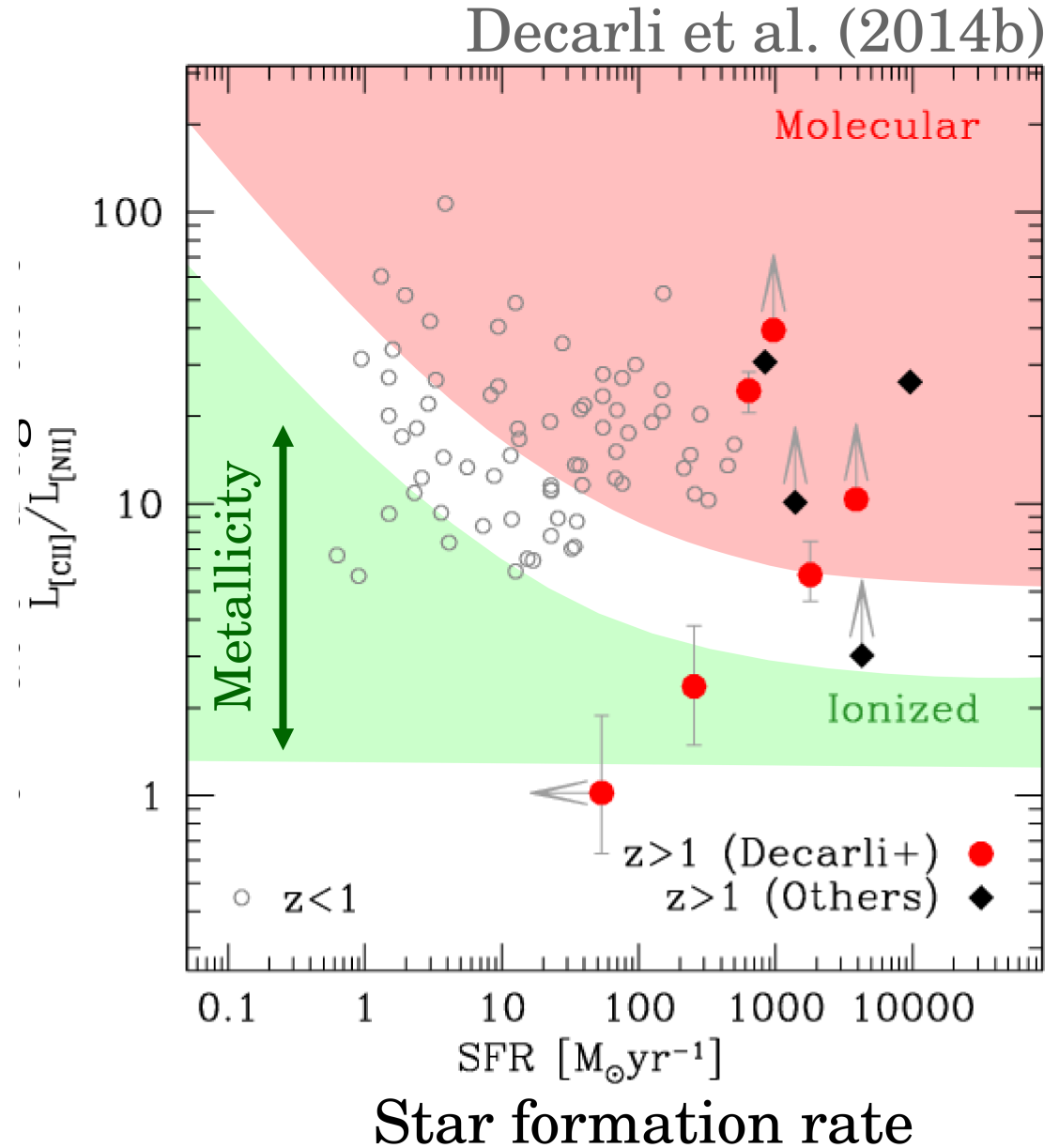
$$[\text{CII}]/[\text{NII}] \sim 1/Z$$

(0th order, it's actually more complicated...)

Also,

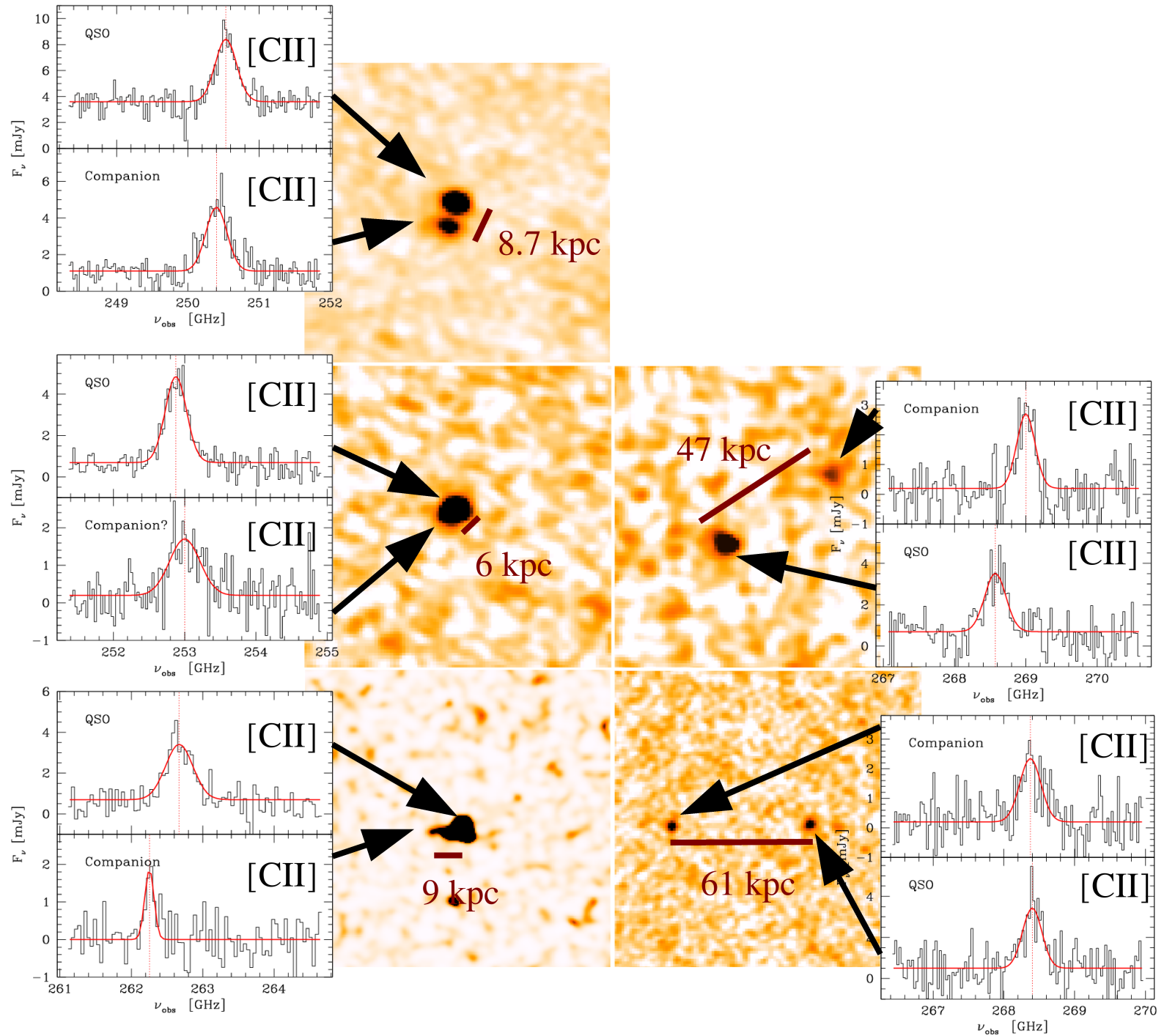
[NII] = only ionized gas

[CII] = ionized+neutral phase

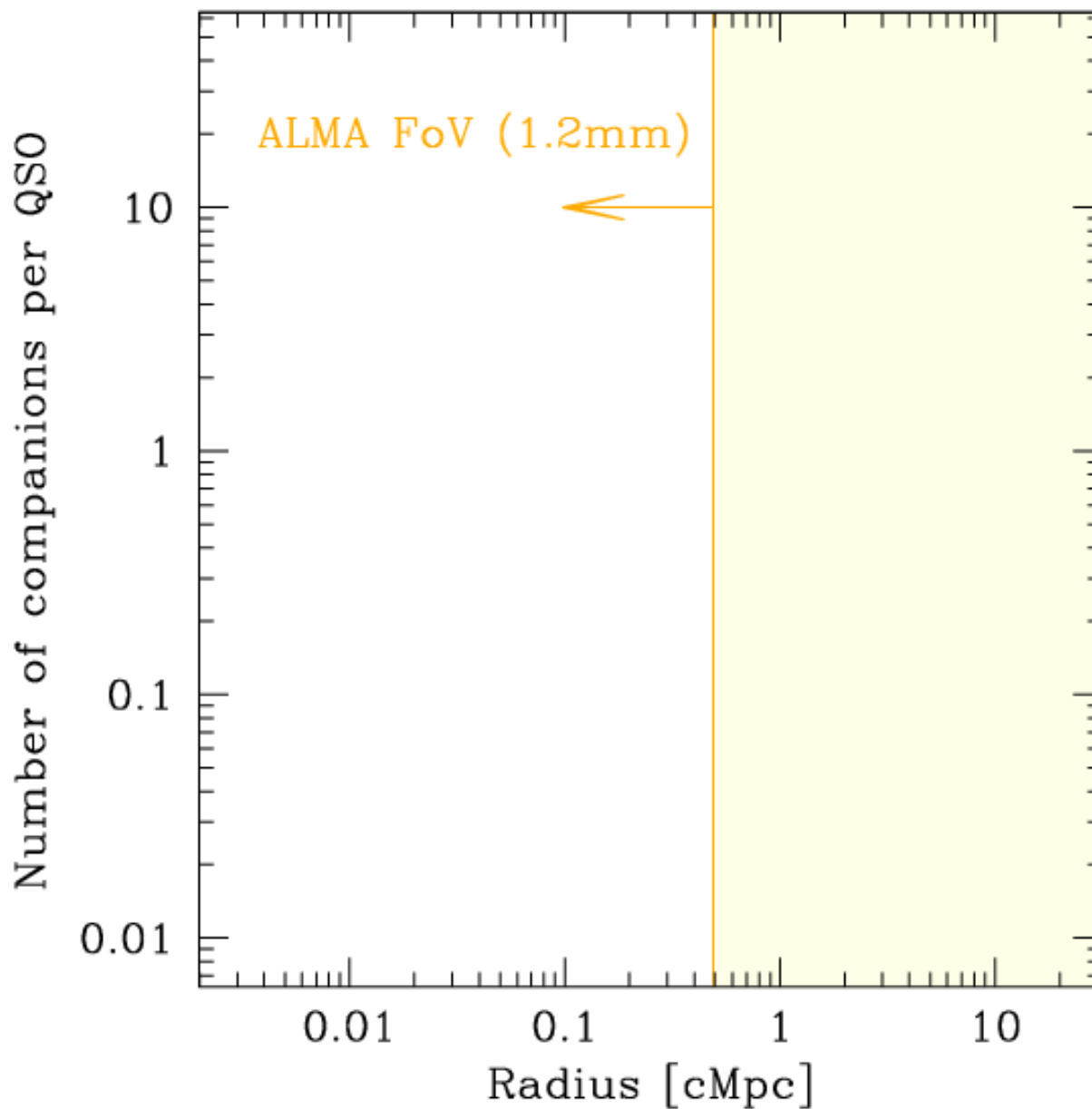


[CII]-emitting companions!

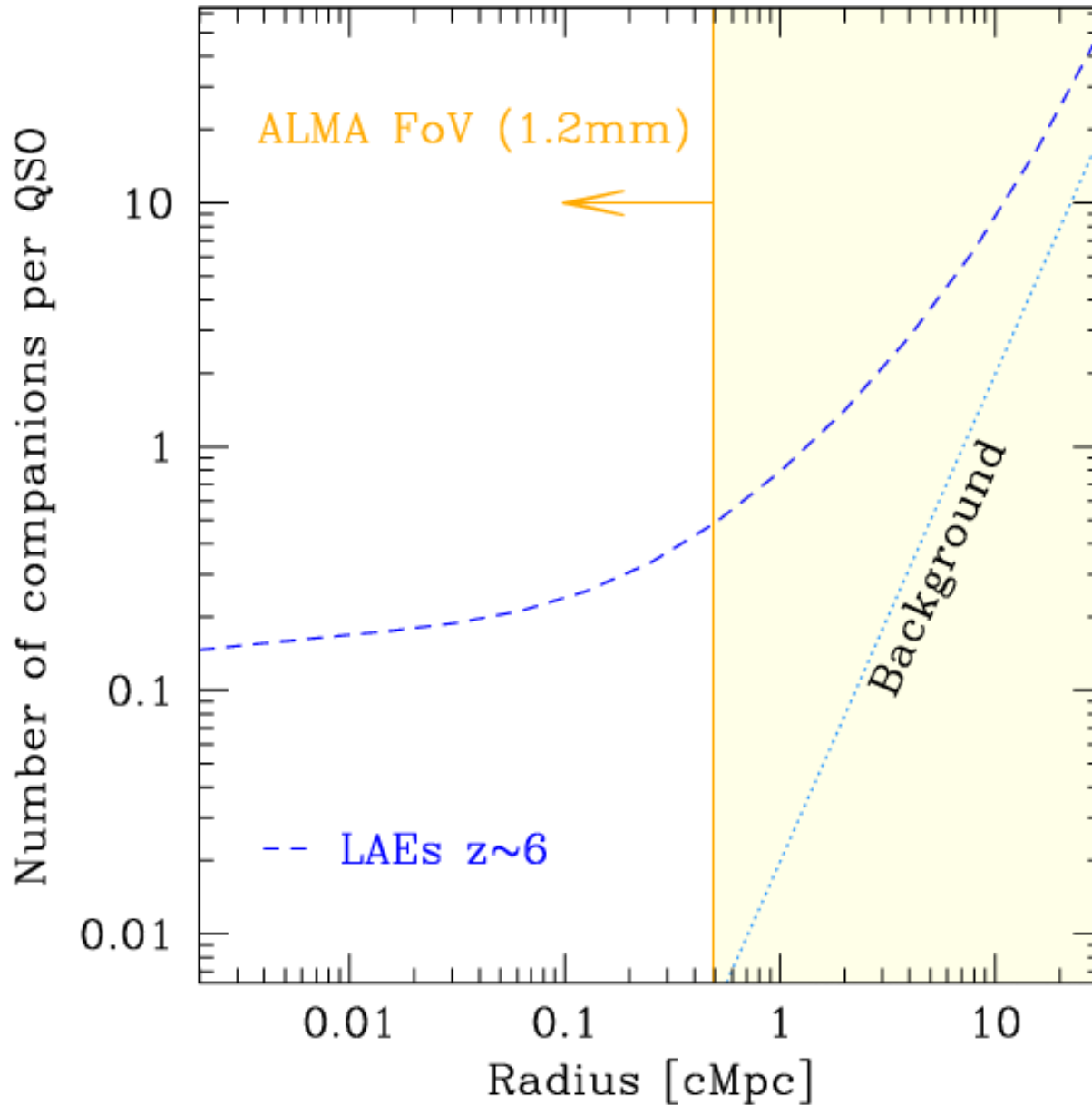
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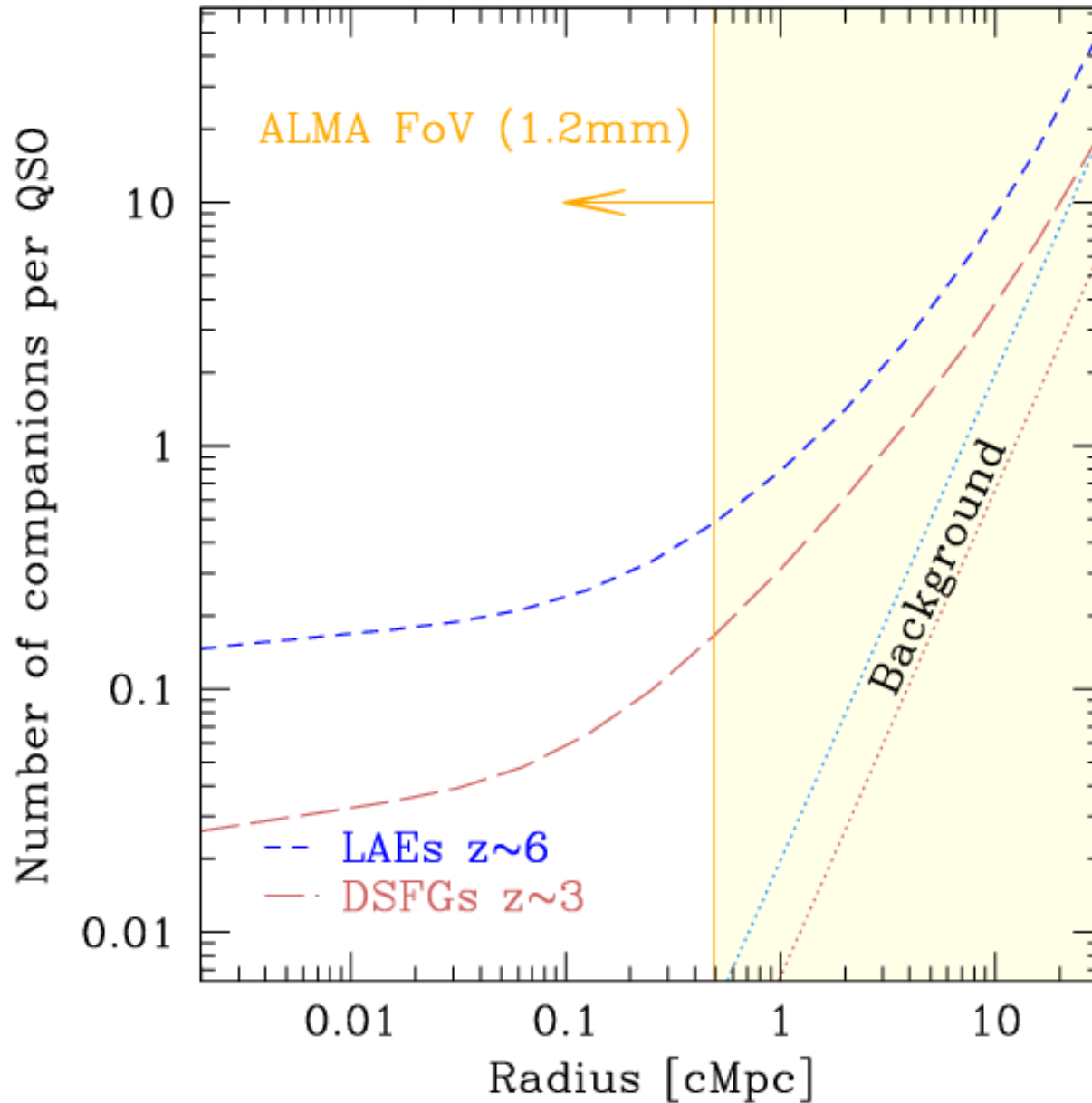
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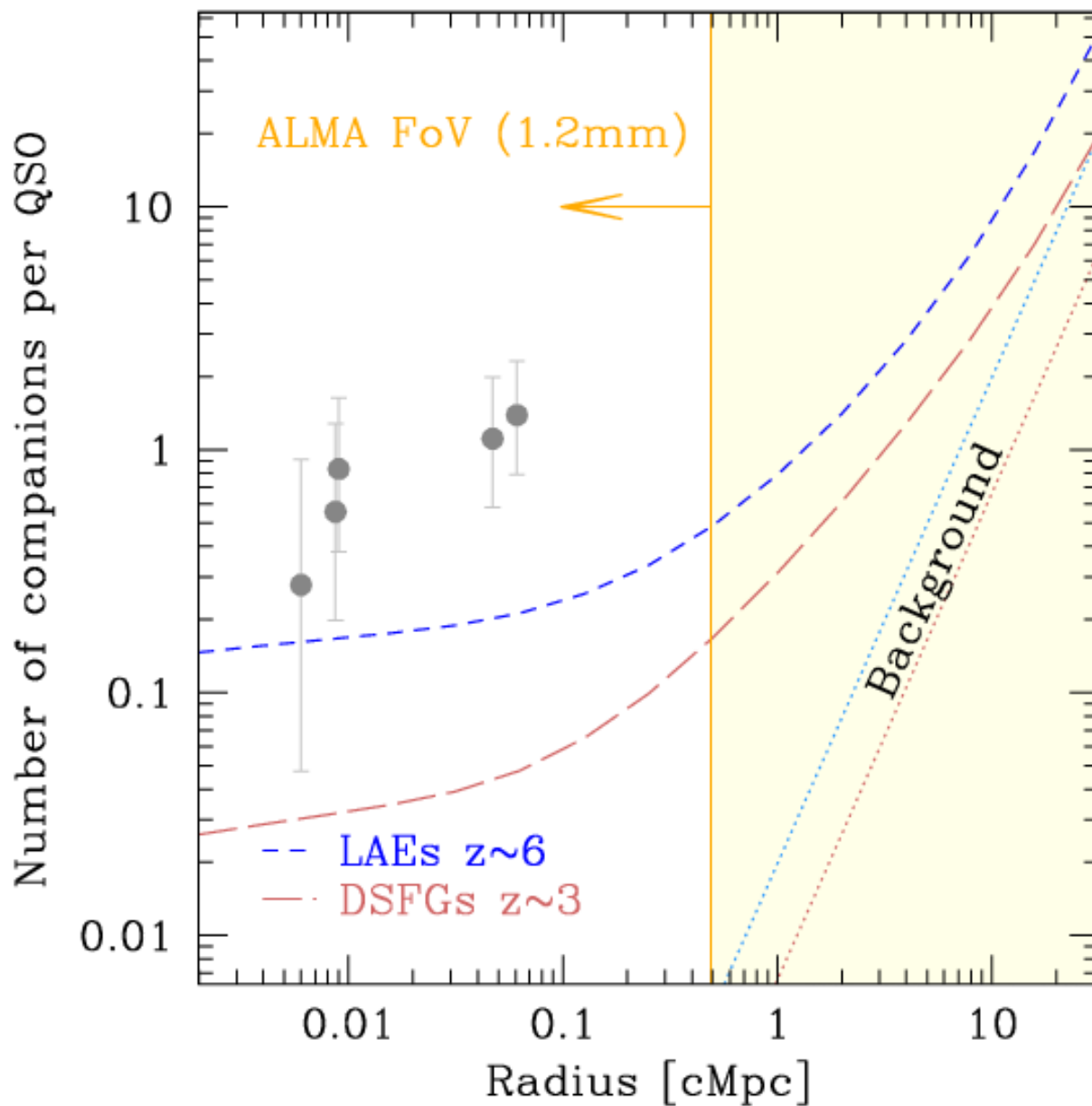
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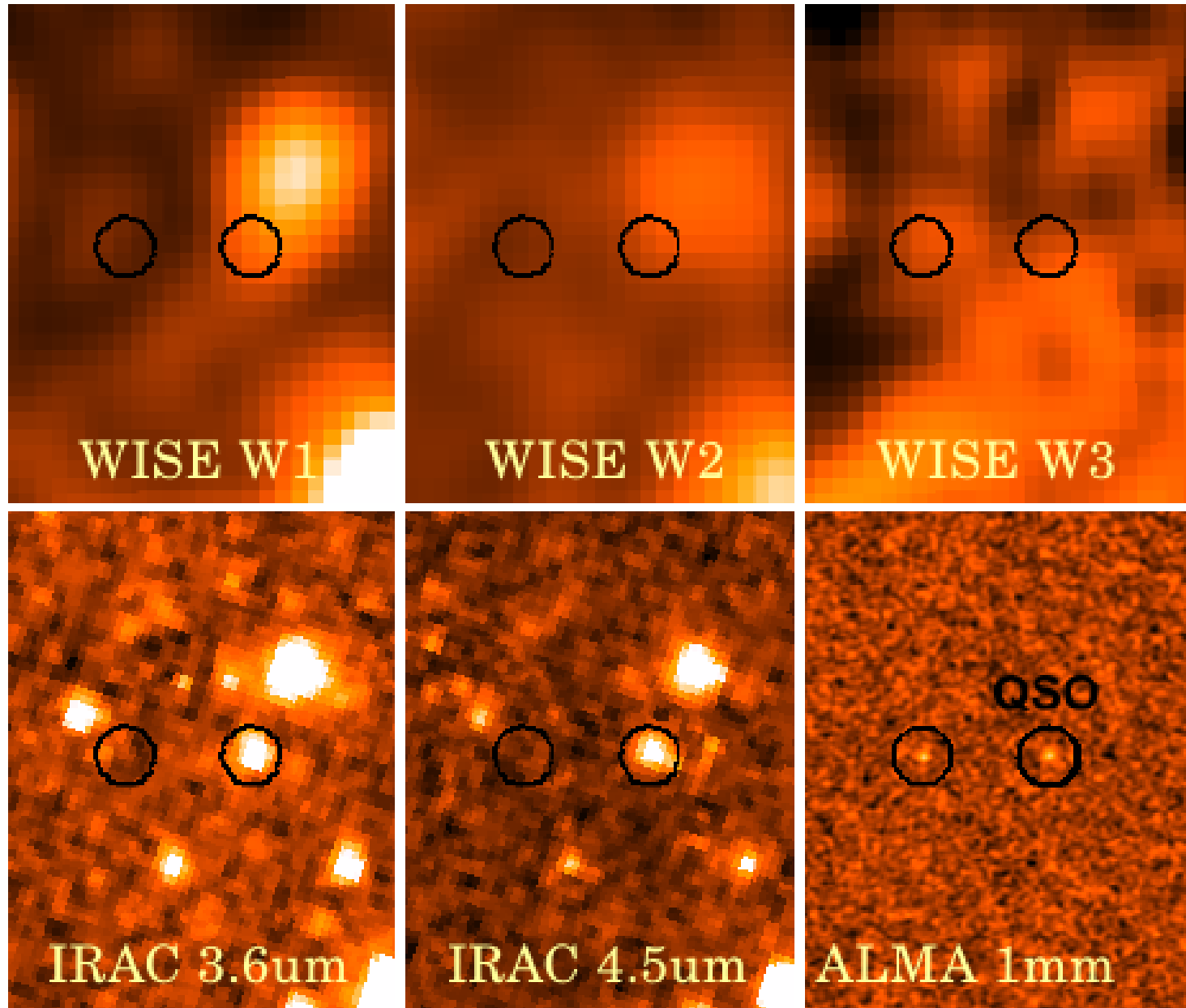
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What kind of galaxies are they?



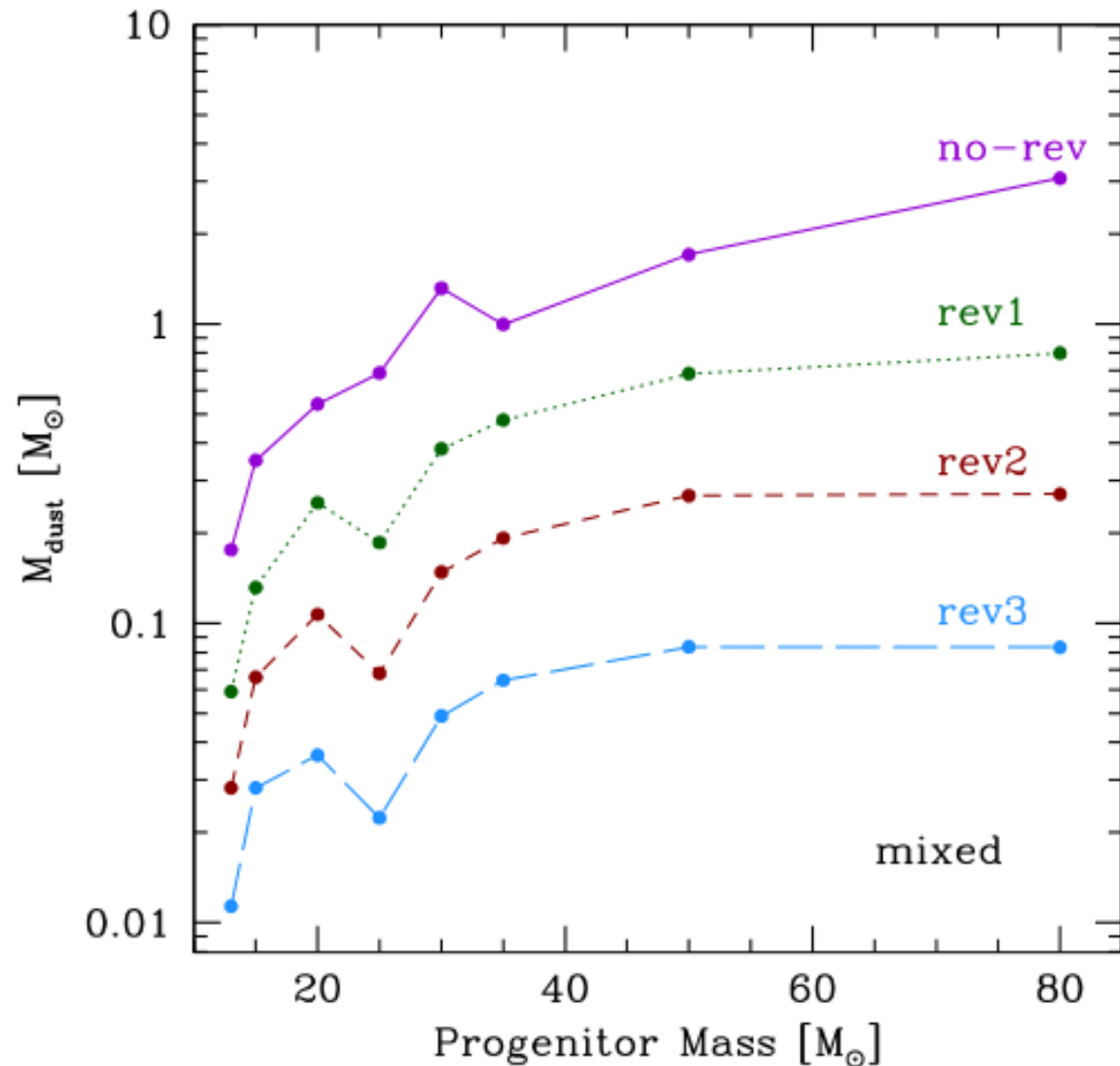
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Spitzer/IRAC non-detections: $M_{\text{star}} < (1-7)e10 M_{\text{sun}}$

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Dust RJ detections:

$$M_{\text{dust}} \sim (1-8)e7 M_{\text{sun}}$$

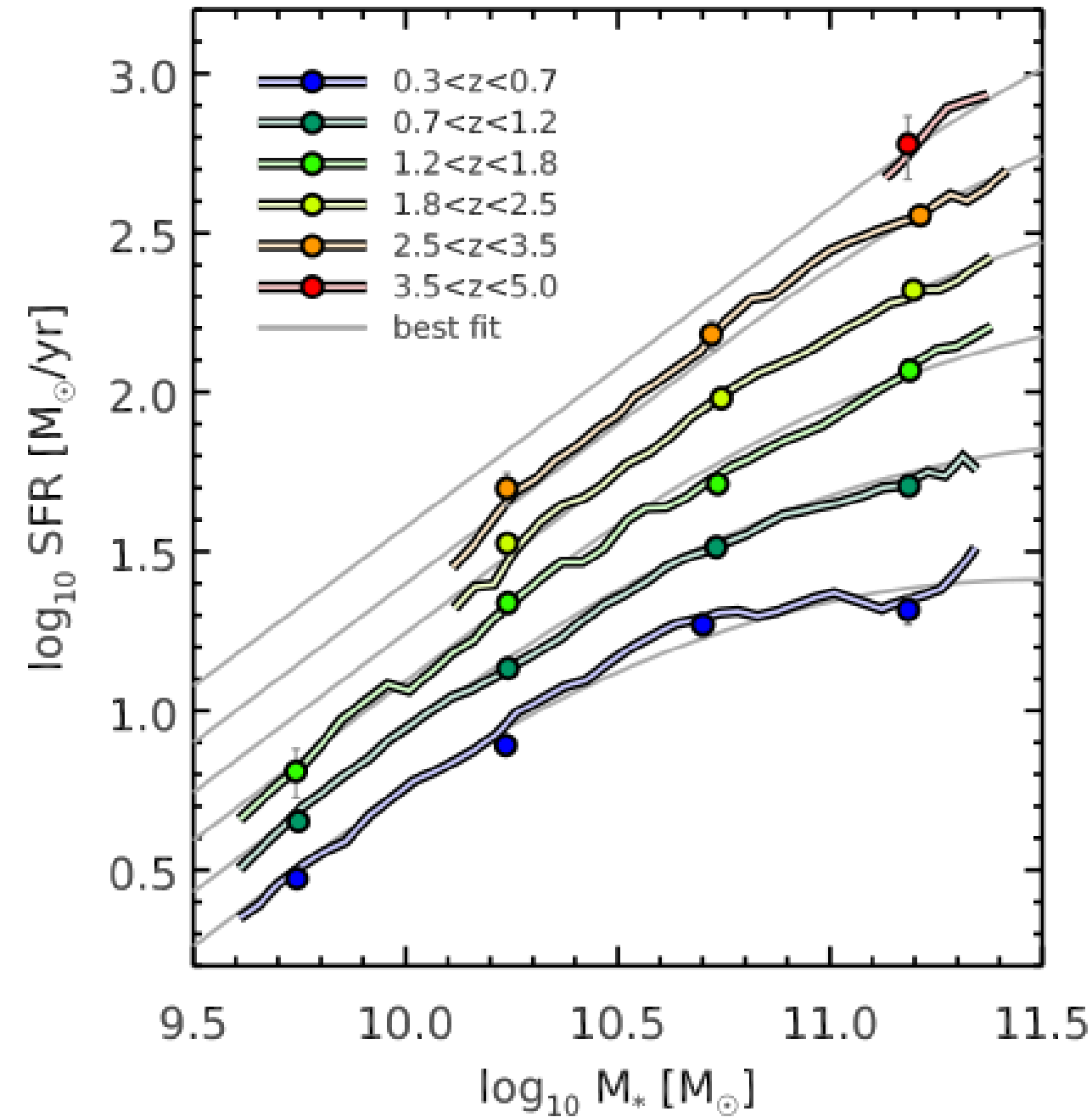
But:

dust yield $y < 1/20$
of progenitor mass

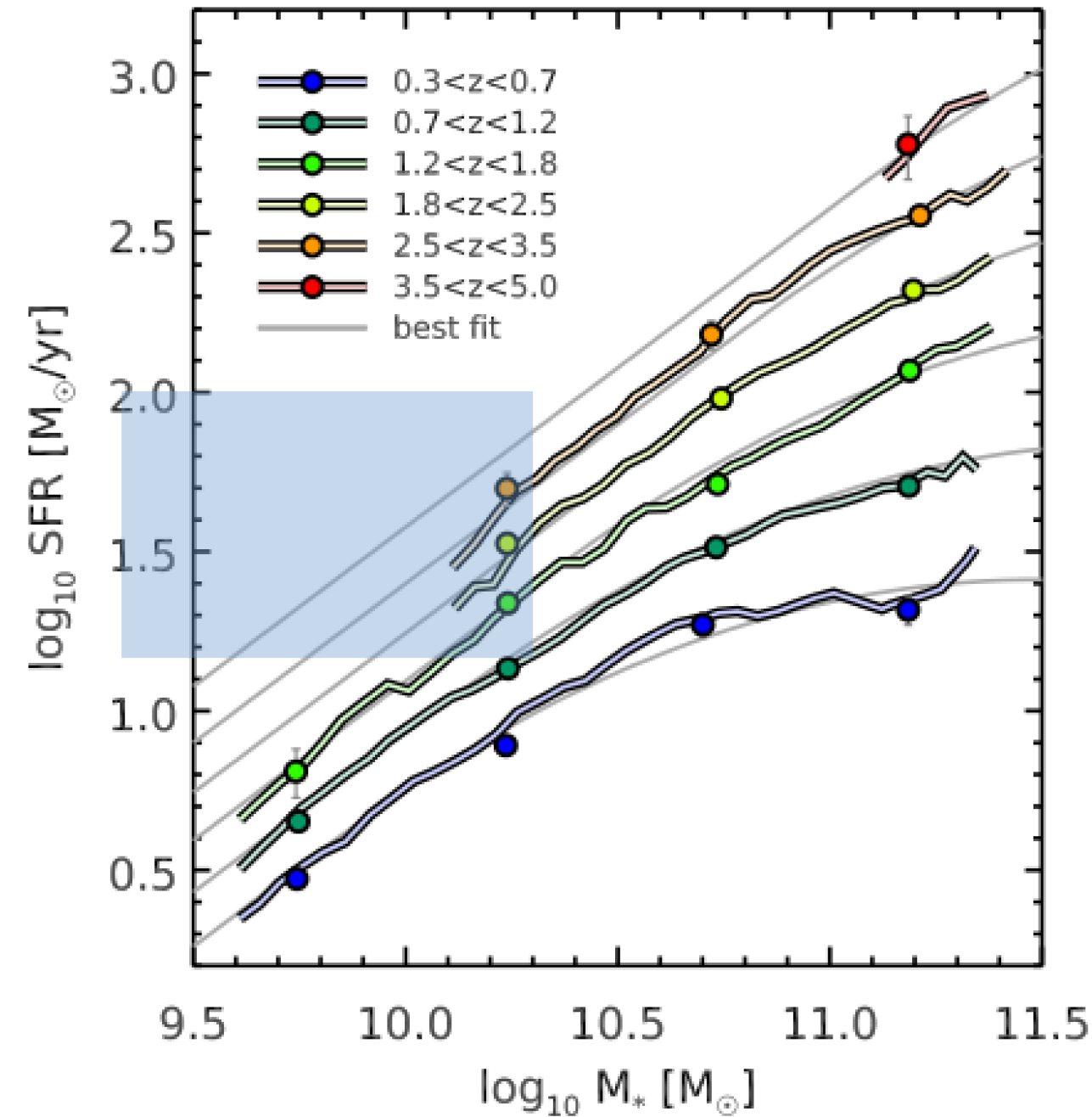
$$\Rightarrow M_{\text{star}} > (2-16)e9 M_{\text{sun}}$$

Marassi et al. (2015)

What kind of galaxies are they?

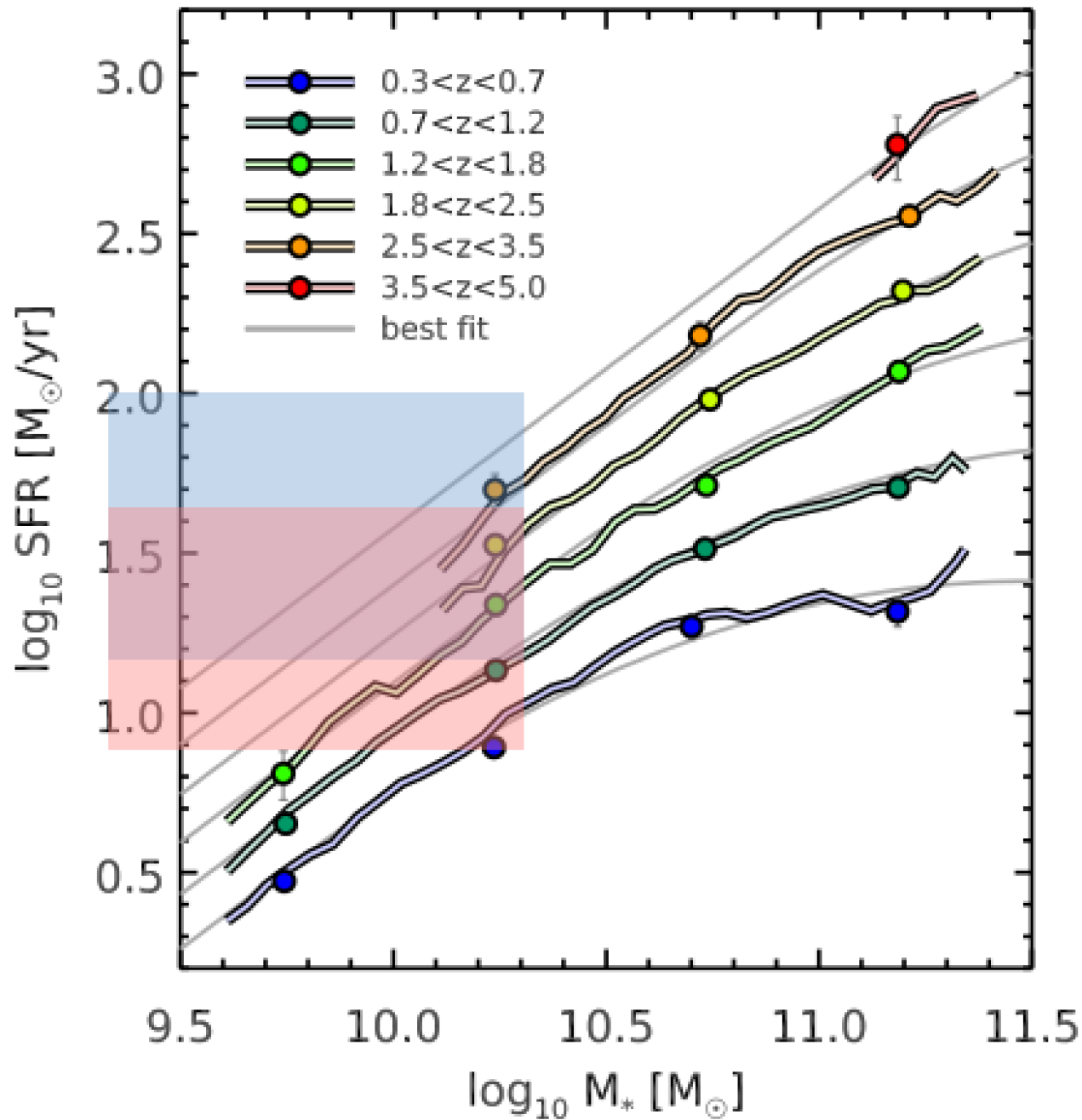


What kind of galaxies are they?



If $T_{\text{dust}} \sim 47 \text{ K}$,
SFR = 15-120 M_{sun}/yr

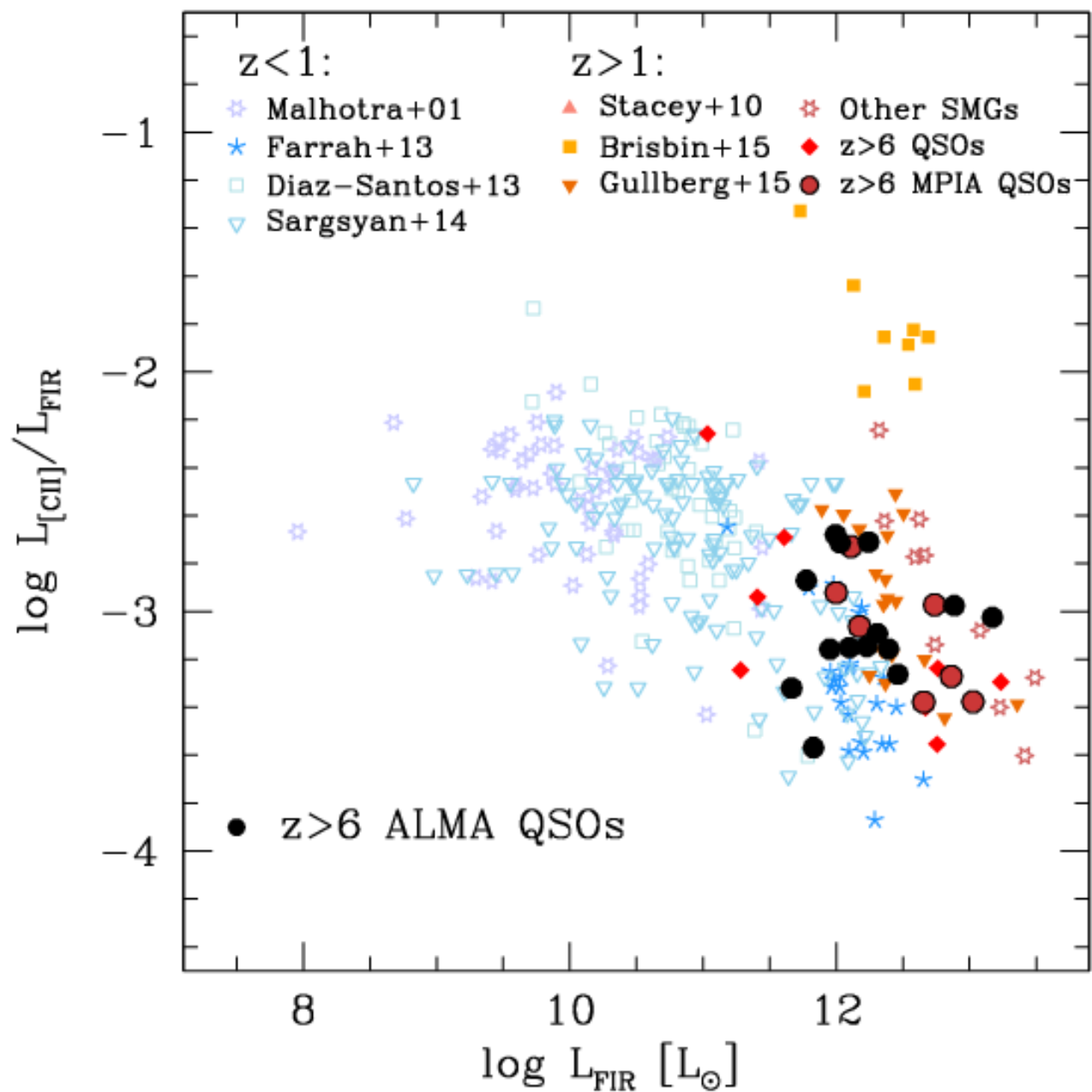
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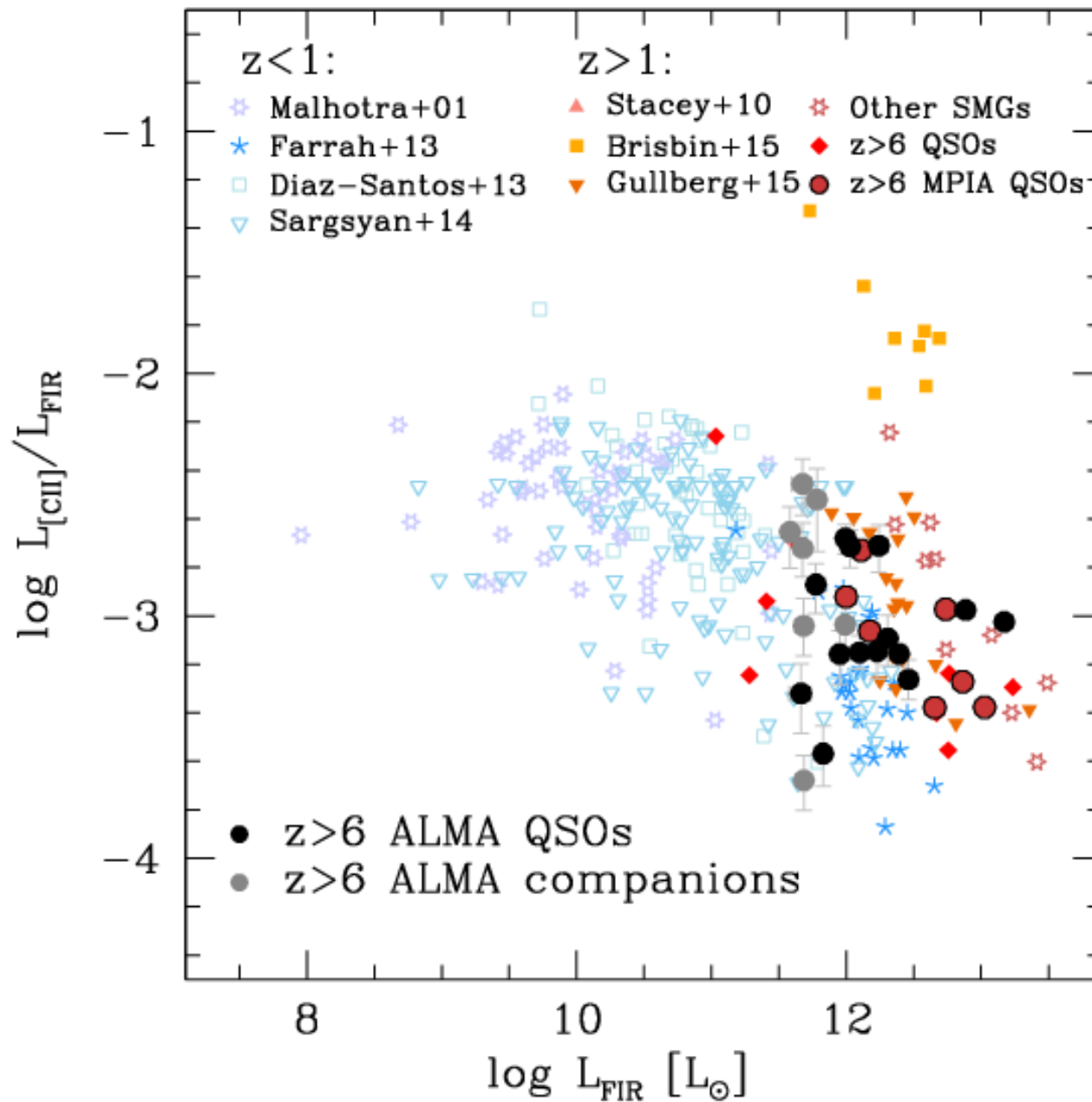
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If $T_{\text{dust}} \sim 25 \text{ K}$,
SFR = 5-40 M_{sun}/yr

[CII] contribution to ISM cooling



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Summary

First systematic survey of [CII] & dust in a statistical sample of $z > 6$ QSOs

>80% [CII] detection rate,
~90% continuum detection rate

~1/3 of the QSOs show a dusty companion galaxy!

