

Early QSOs and their SMBHs from KiDS+VIKING

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Bram Venemans, team QSO at MPIA, Kuijken, Mwebaze, Valentijn

Survey Science:

KiDS/VIKING production team, OmegaCEN, Target
McFarland, Helmich, de Jong, Irisarri, Williams, ++
Astro-WISE information system

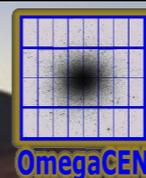


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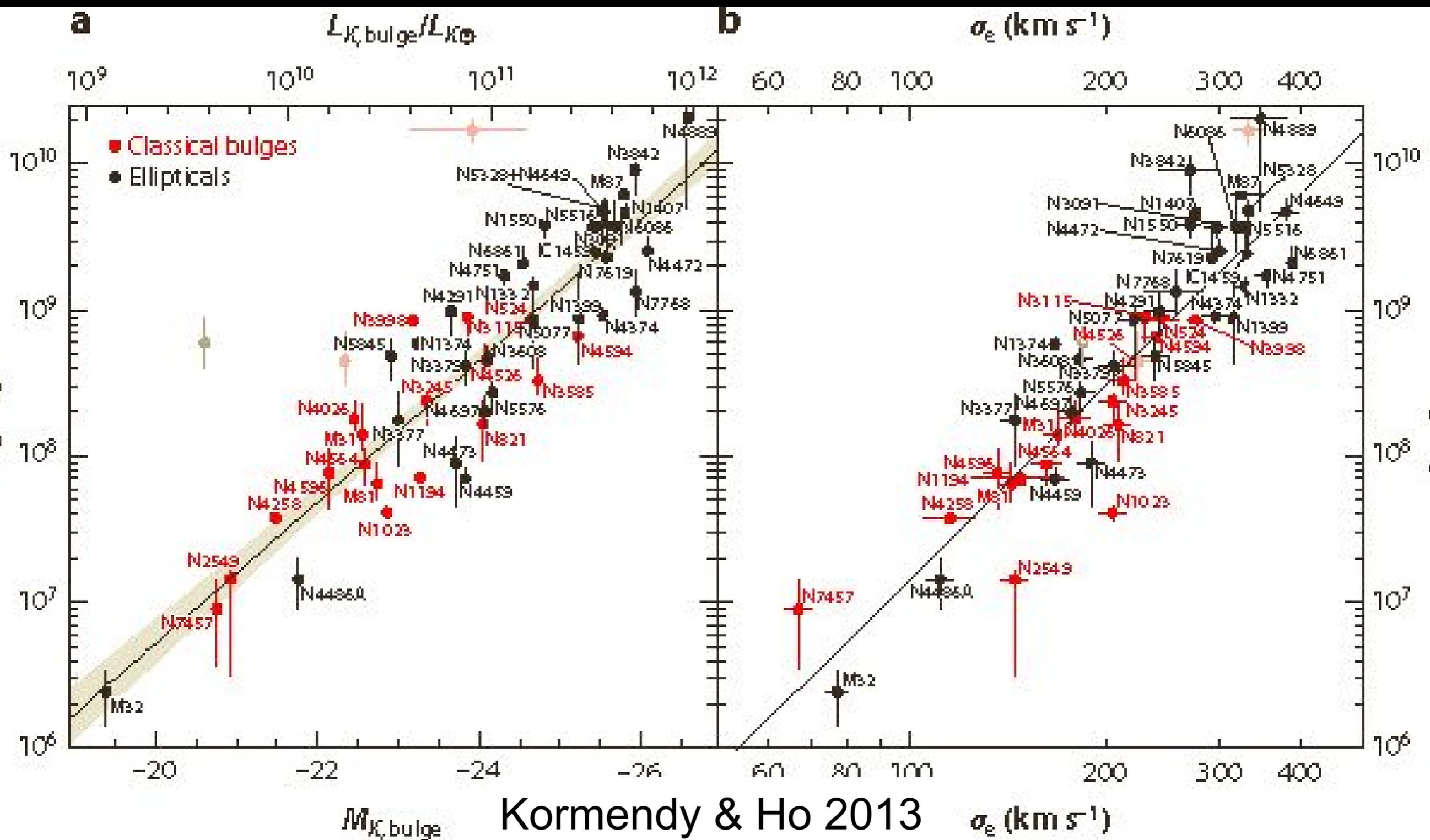
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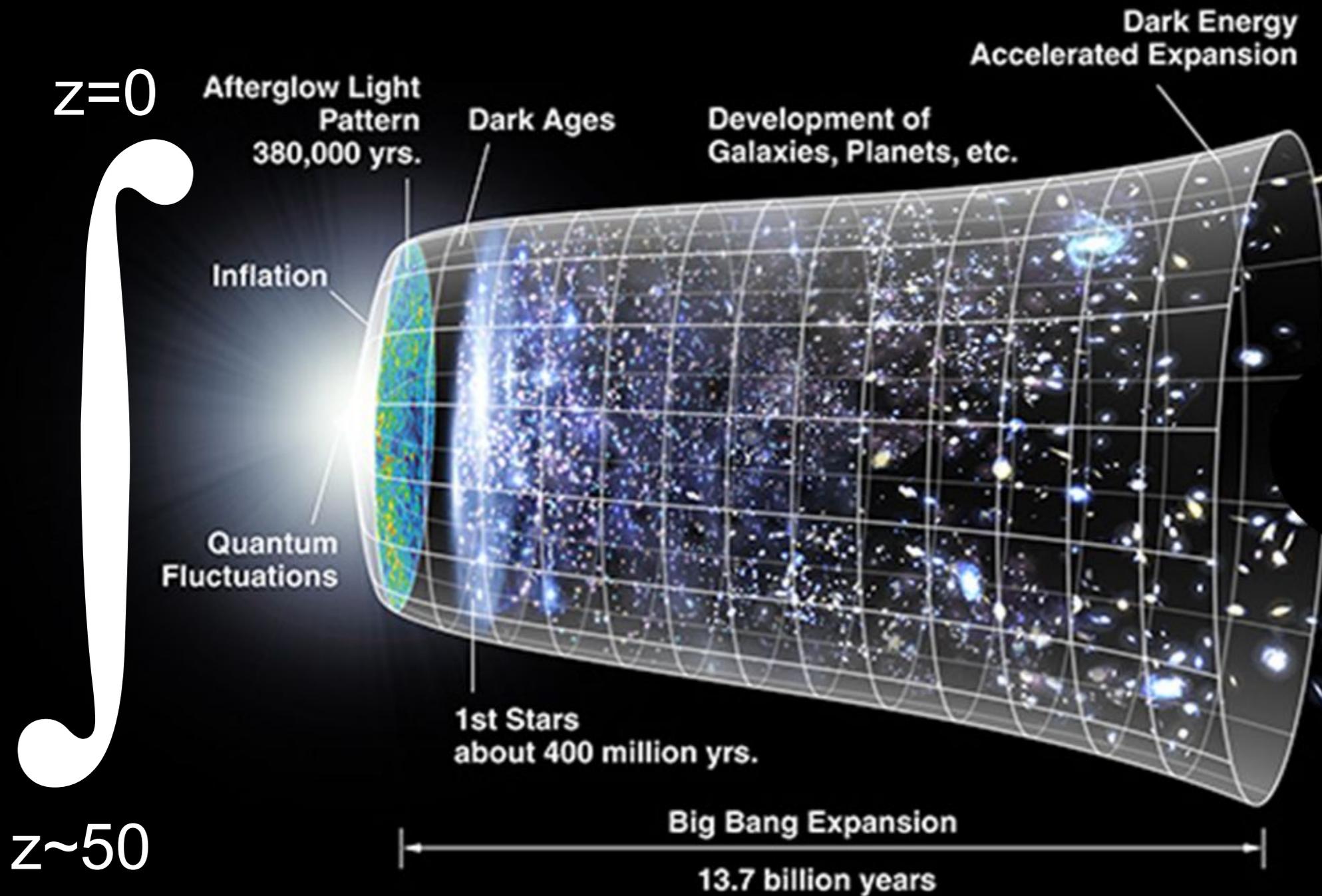
KiDS



We need to end up with this....

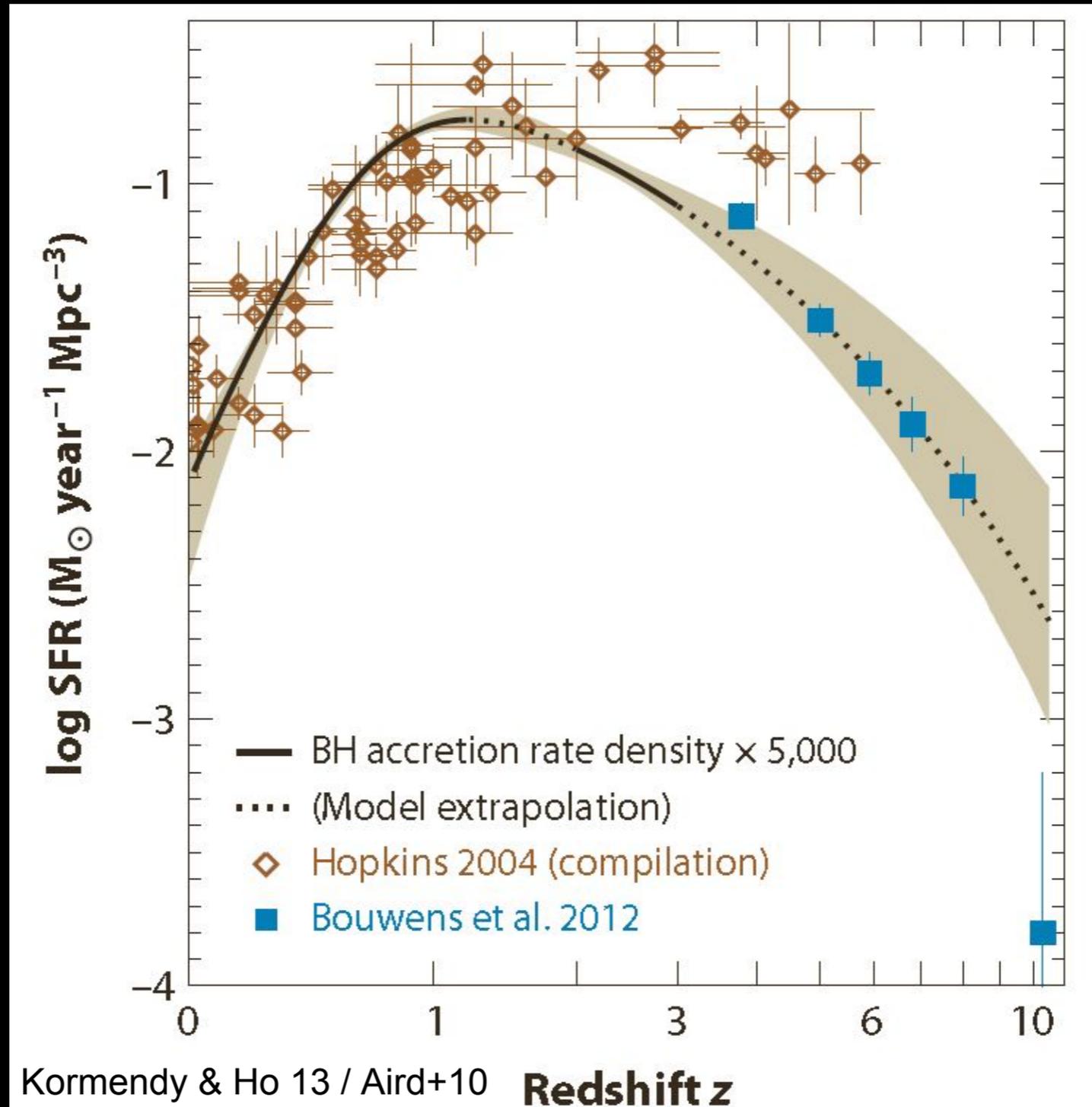


supermassive blackhole ↔ host symbiosis

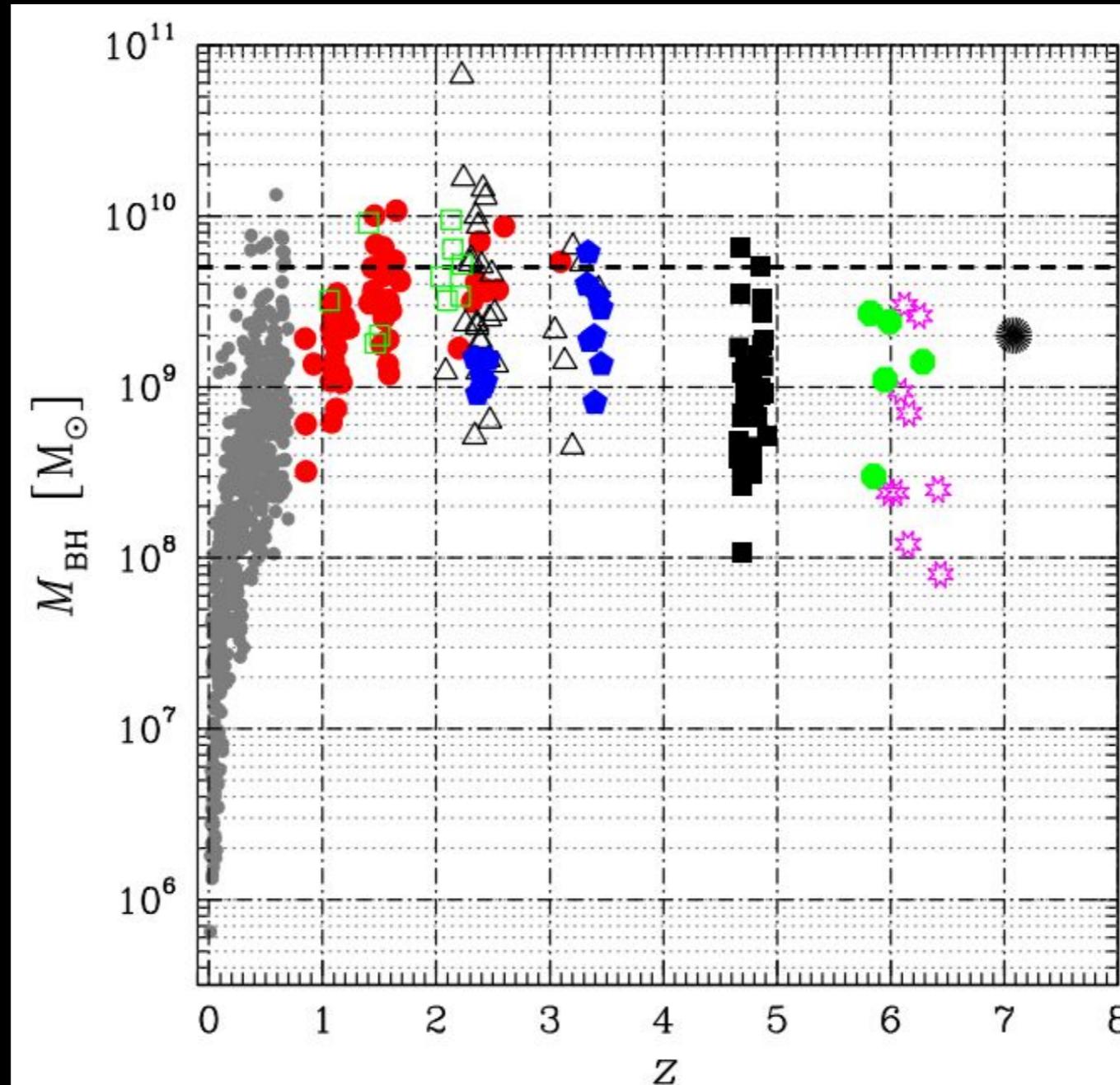


QSOs ideal:
SMBHs
hosts
IGM
feedback

Co-evolution on average

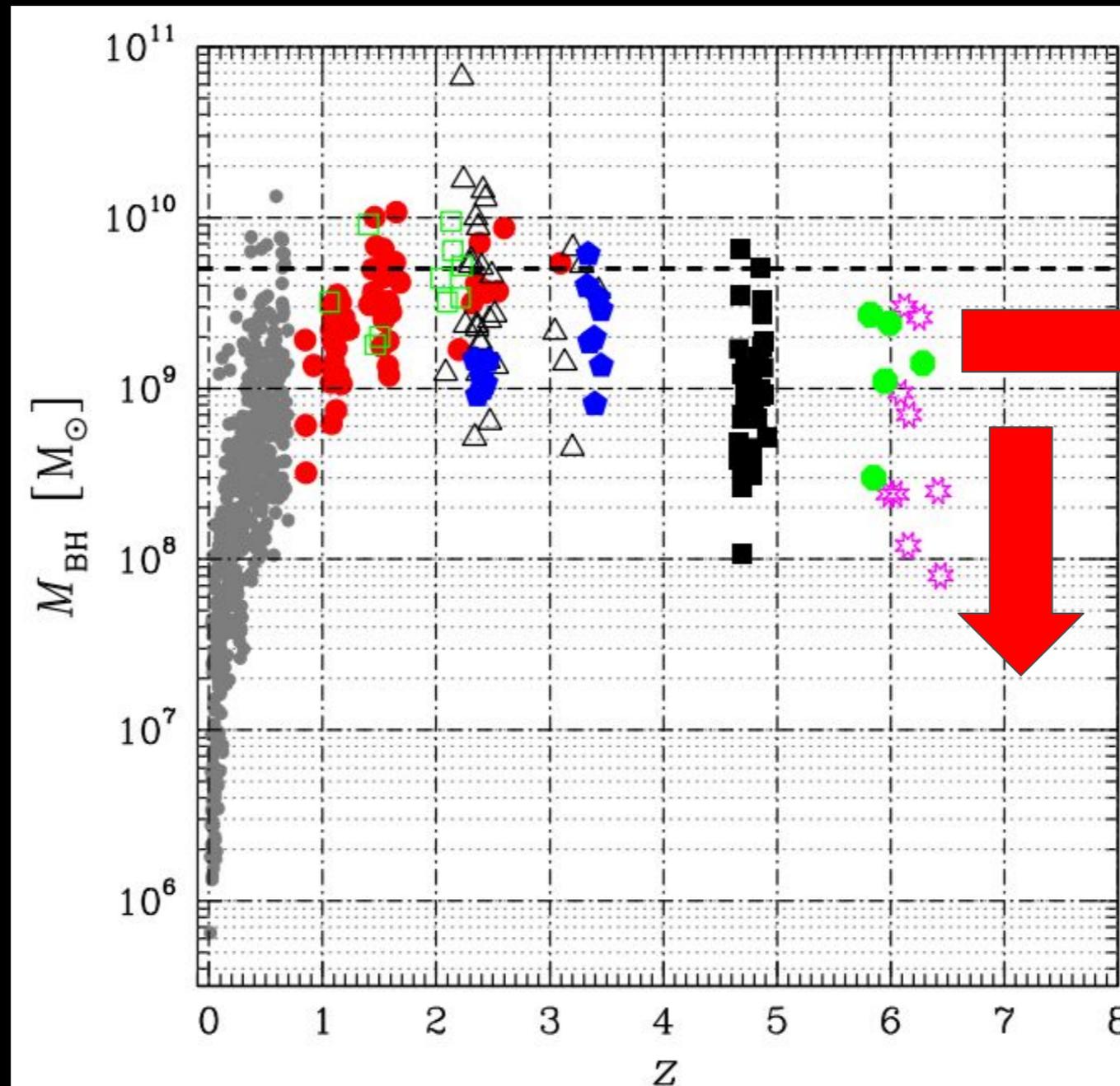


Varying “evolutionary clockspeeds” (aka downsizing)



Marziani, Sulentic 11

What theorists *and* observers want



Marziani, Sulentic 11



survey telescopes

VIRCAM@VIST

4m telescope

FoV 0.6 sq.deg, NIR
16 2kx2k, 0.35" pix

VIKING

passband	AB, 5 σ , 2"
Z _v	22.7
Y	22.0
J	21.8
H	21.1
K	21.2

Ω CAM@VST

2.6m telescope

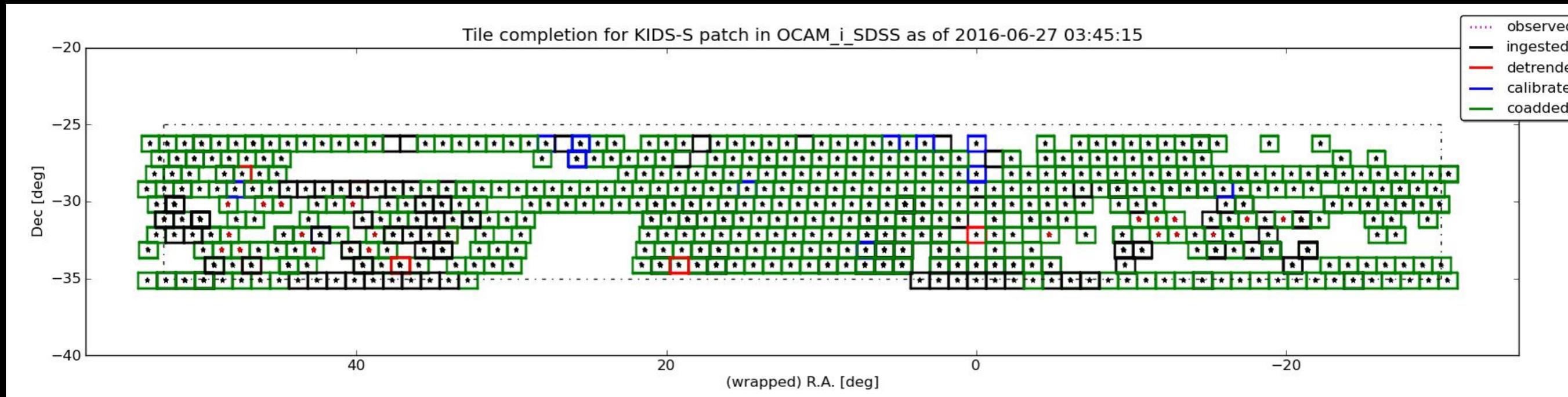
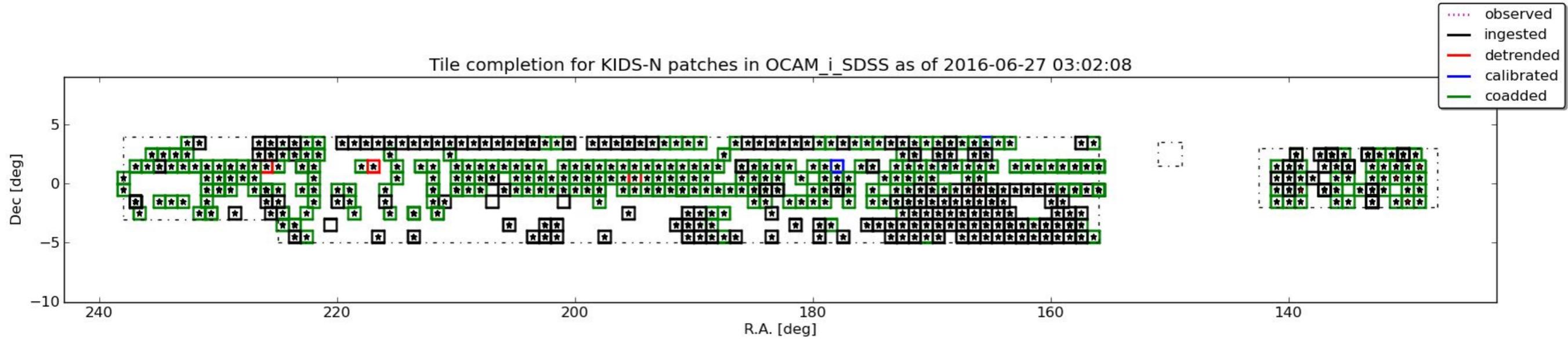
FoV 1 sq.deg, OPT
32 2kx4k, 0.21" pix

KiDS

passband	AB, 5 σ , 2"
u	24.2
g	25.2
r	25.0
i	23.8

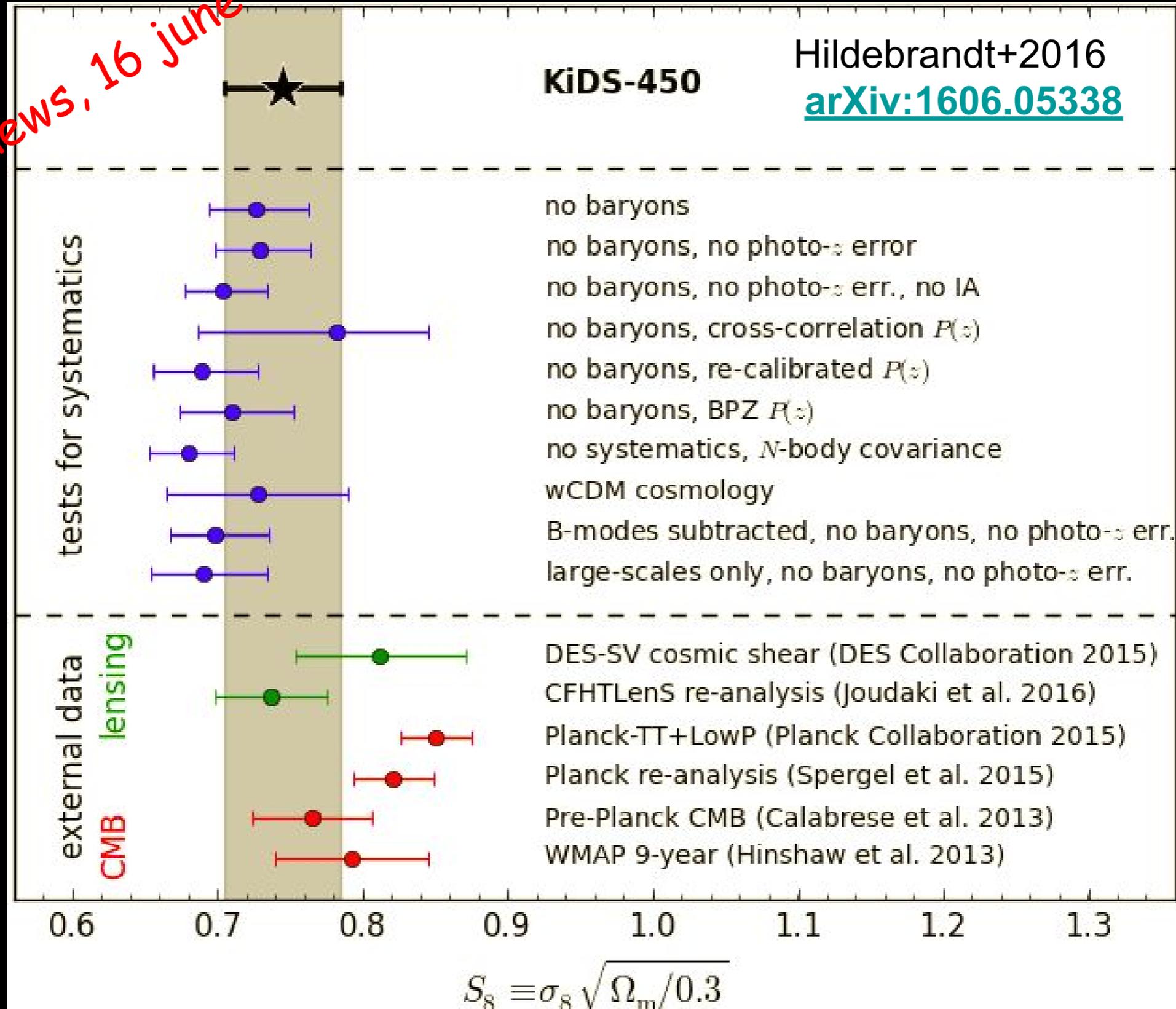


KIDS+VIKING iZYJHK ~1200sq.deg. today

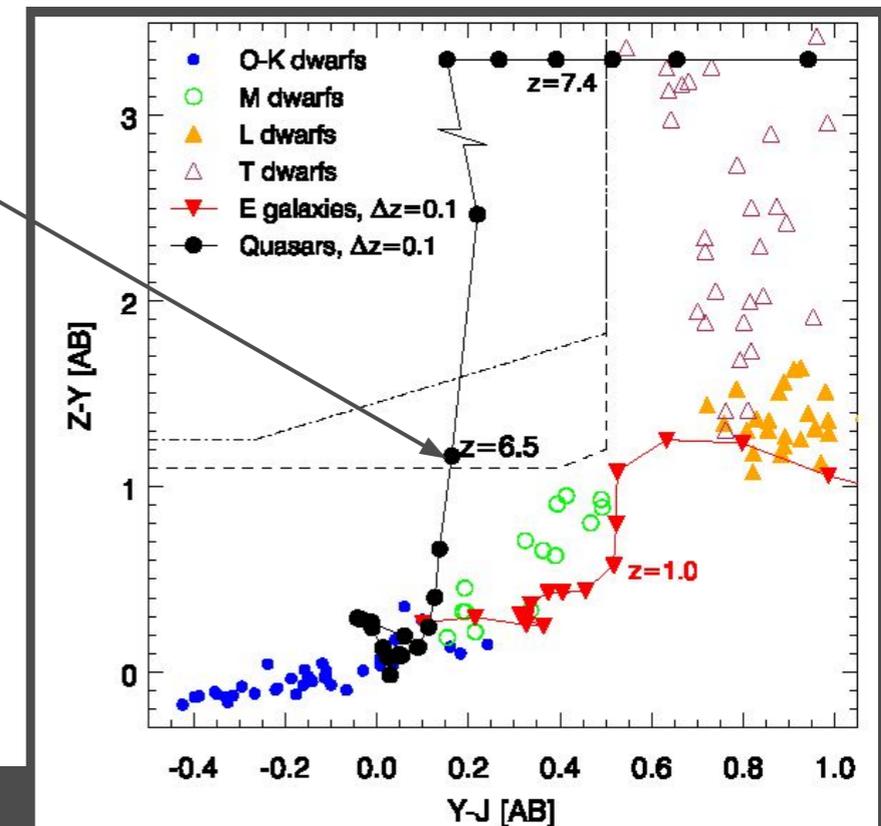
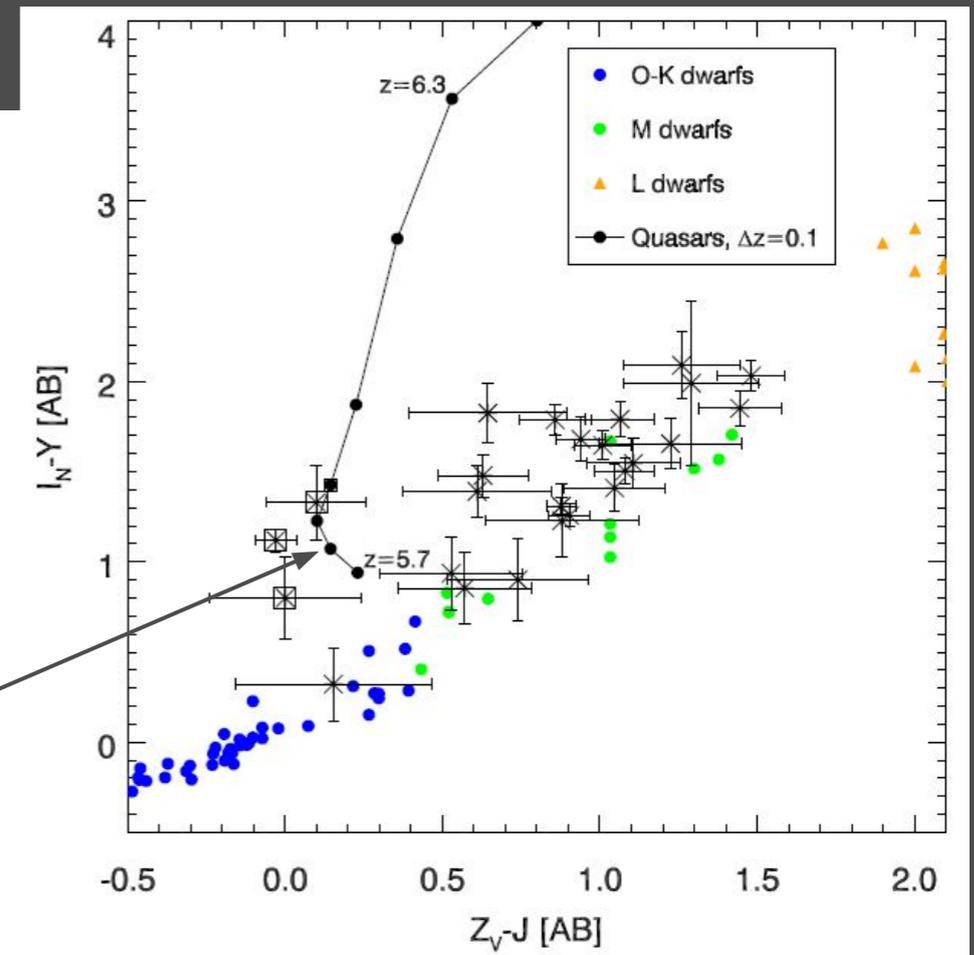
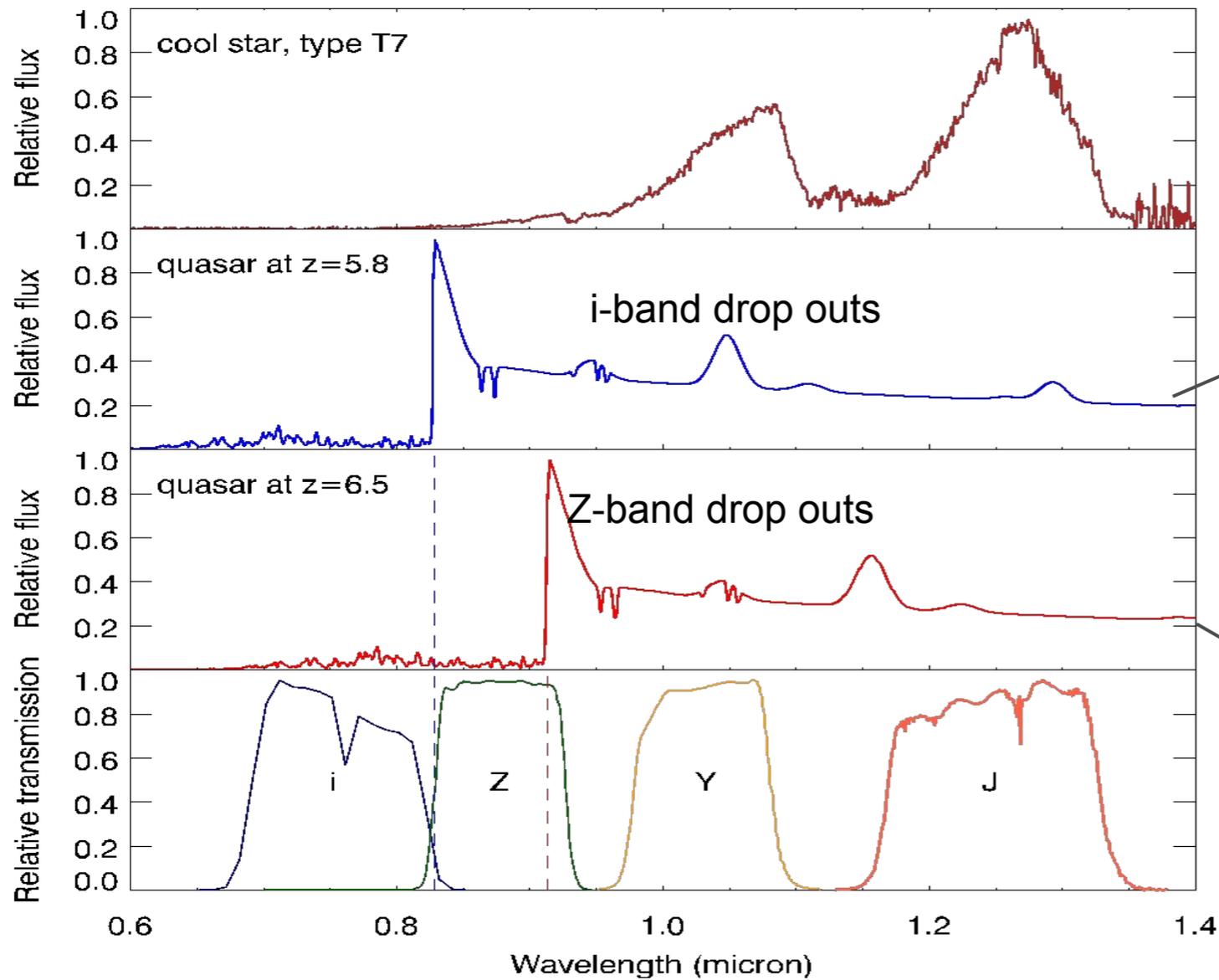


Tension in fundamental cosmology

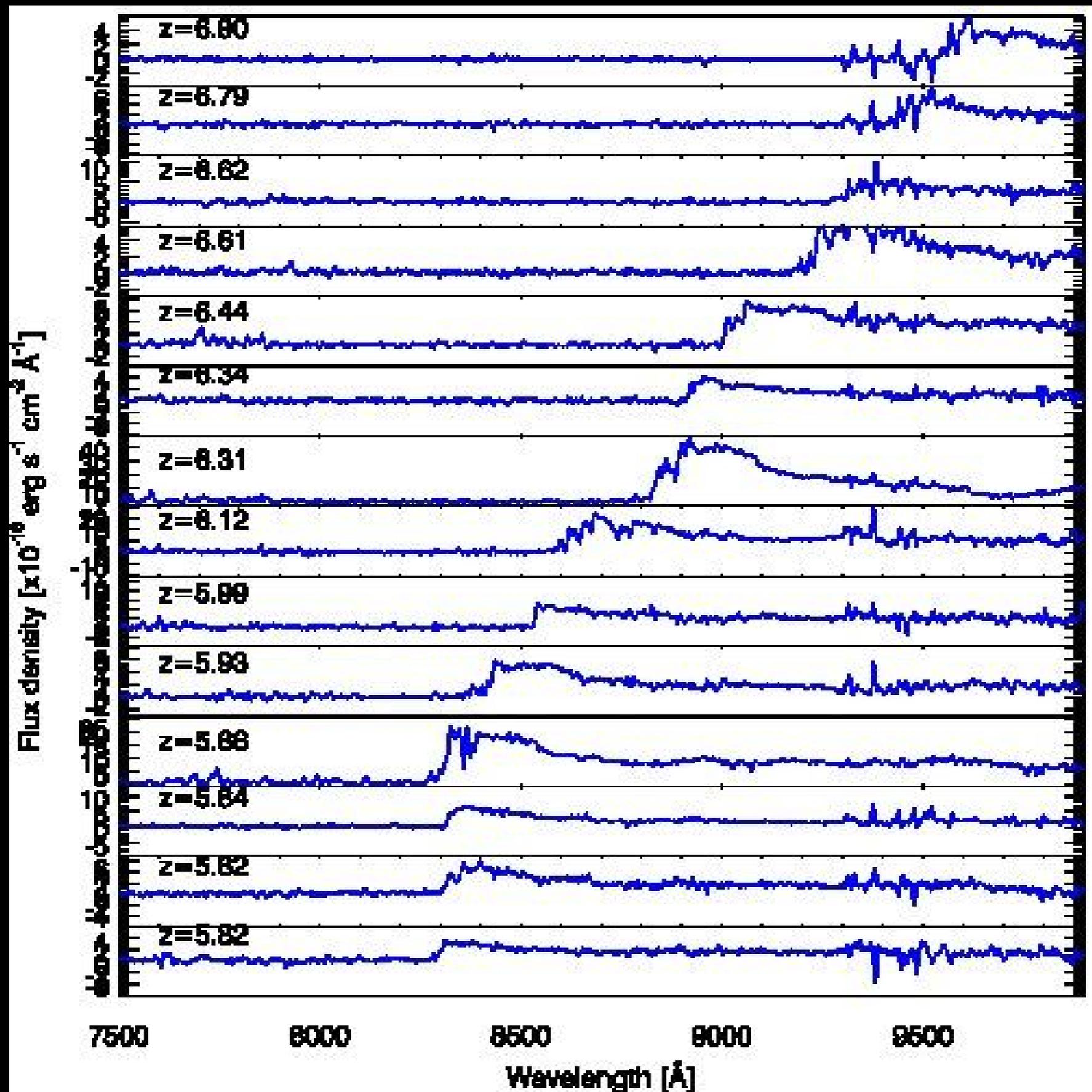
Latest KiDS news, 16 June



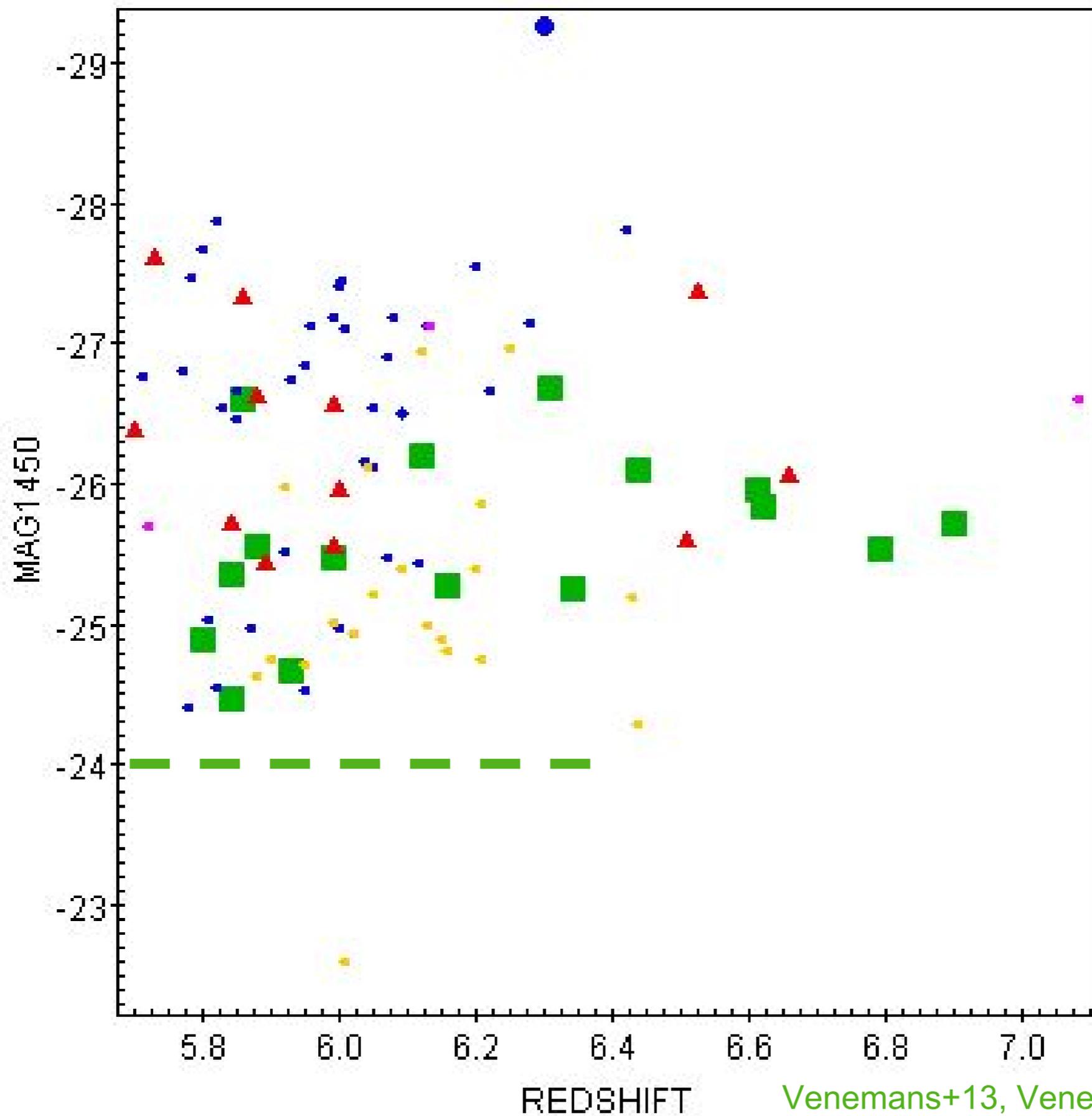
High-z QSO selection and follow-up



VIKING+KiDS QSO harvest 30Jun 2016 :16



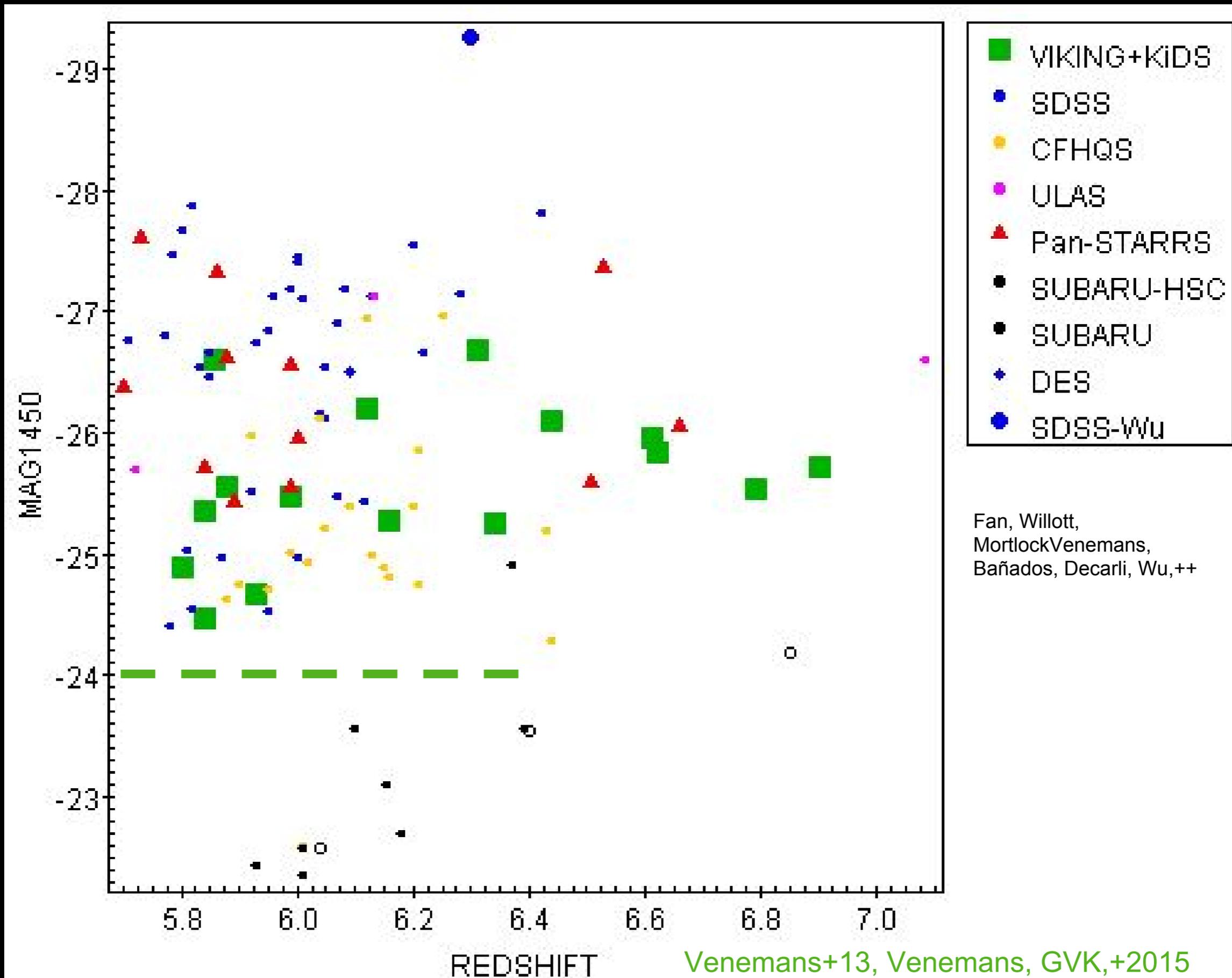
+2



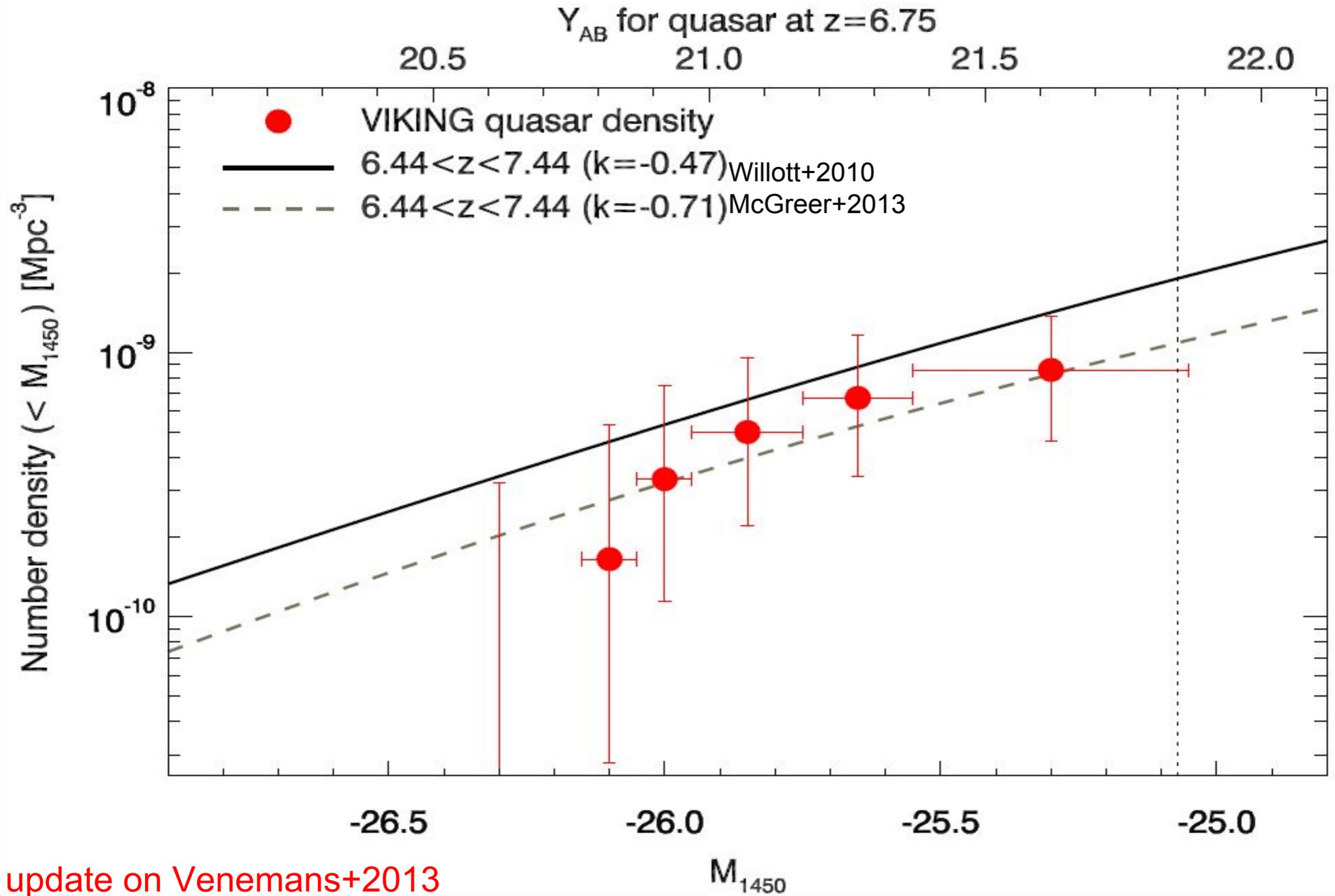
- VIKING+KIDS
- SDSS
- CFHQS
- ULAS
- ▲ Pan-STARRS
- ★ DES
- ★ SDSS-WU

Fan, Willott,
Mortlock Venemans,
Bañados, Decarli, Wu,++

Venemans+13, Venemans, GVK,+2015



QLF at $z \sim 7$, favoring steeper decline



update on Venemans+2013

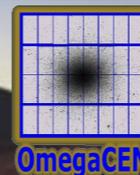


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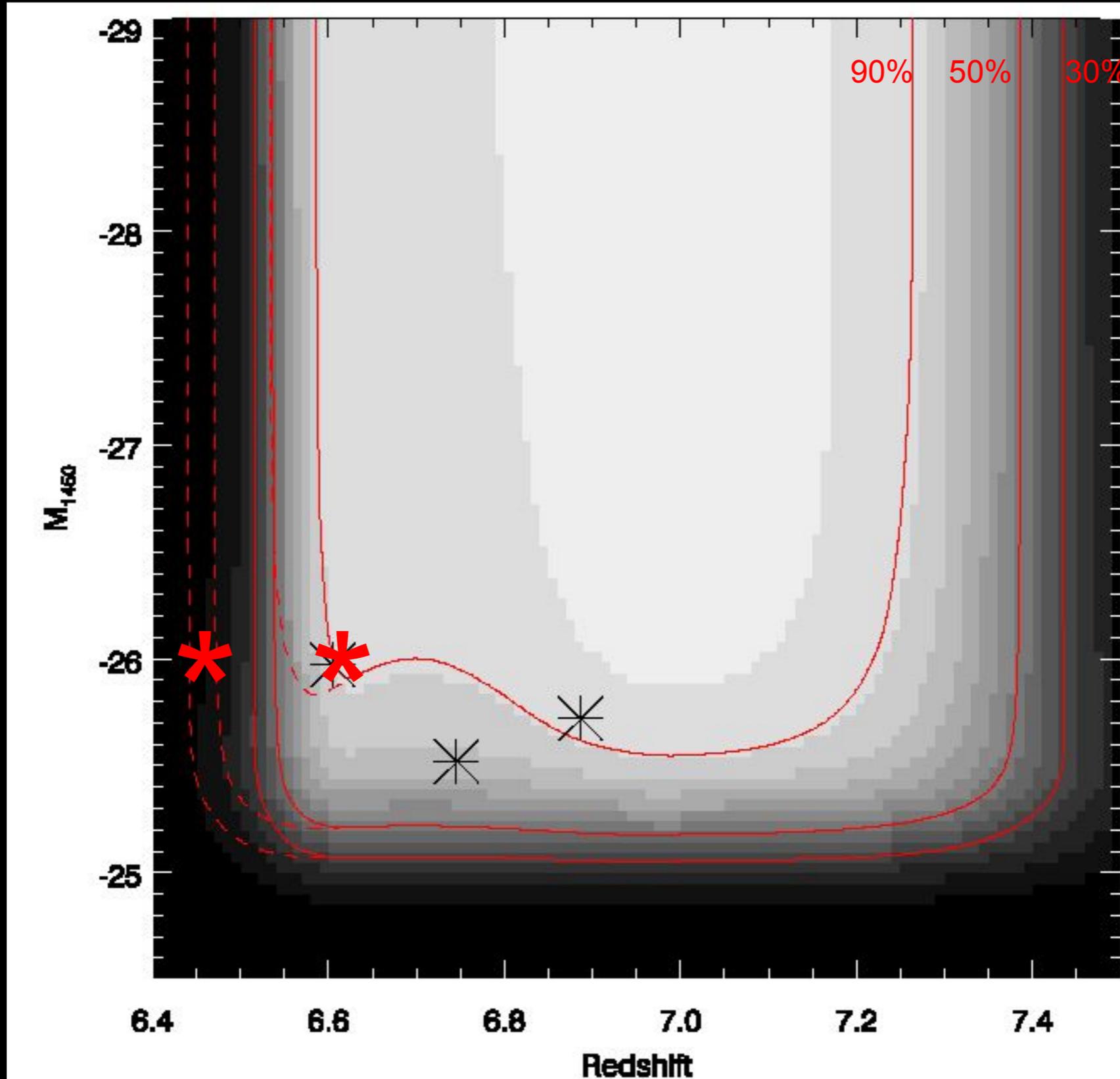
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98 99 100 101 102 103 104

Completeness

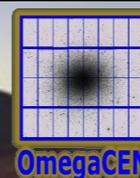


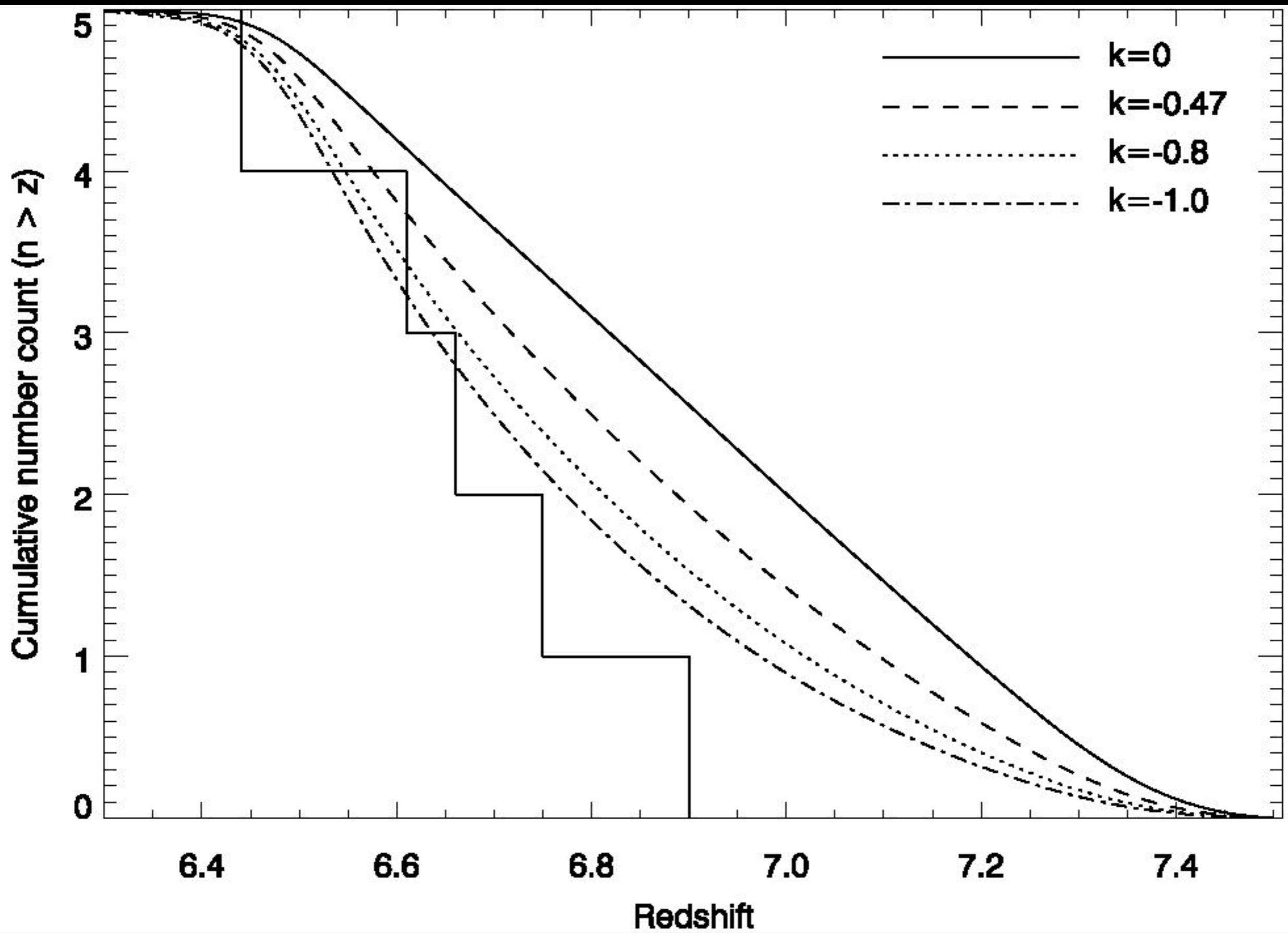
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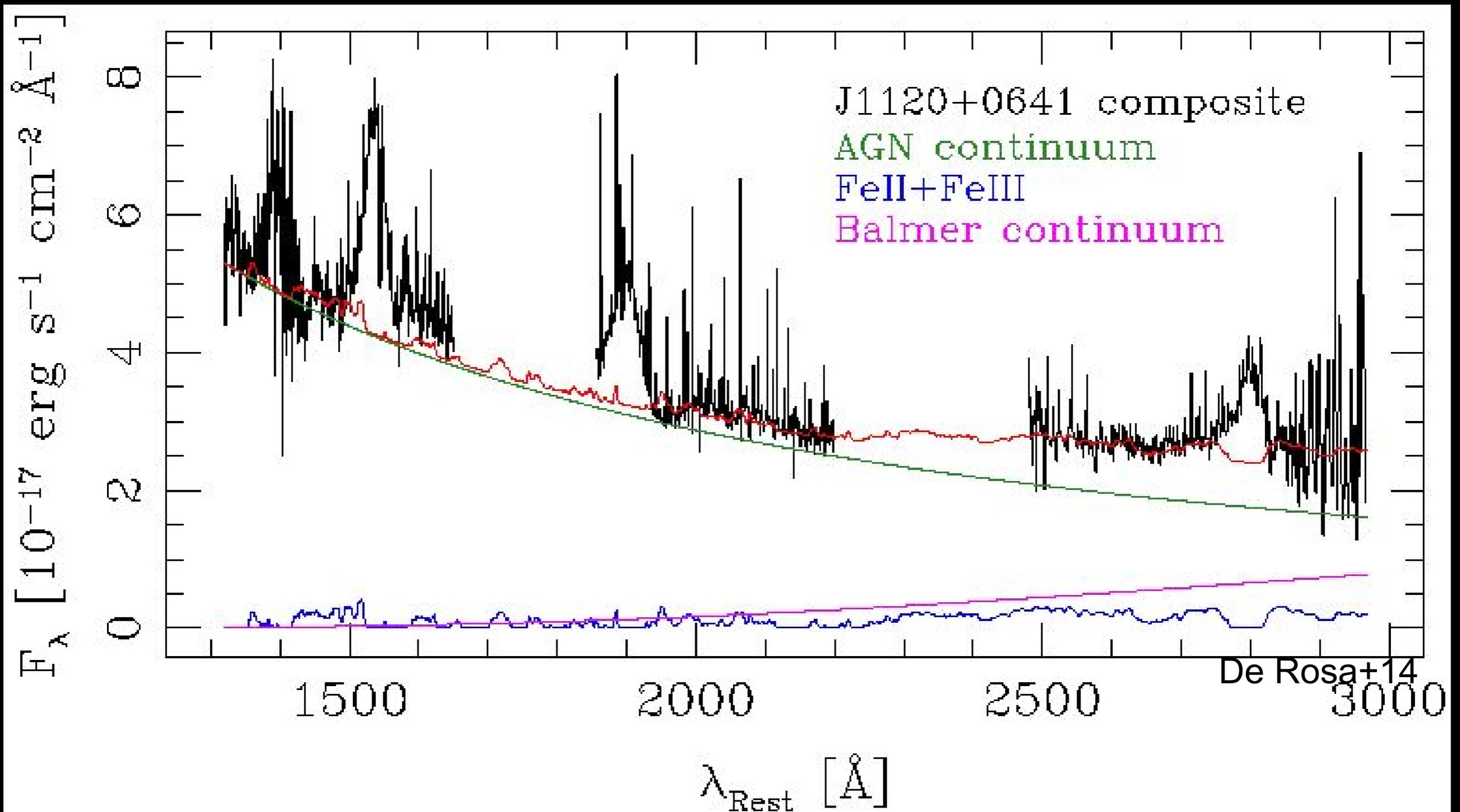
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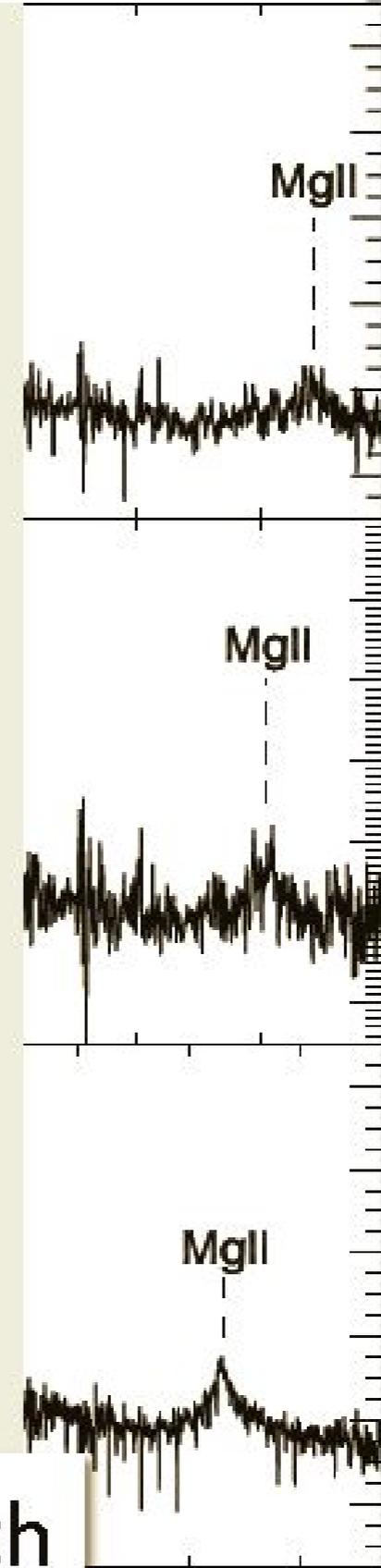
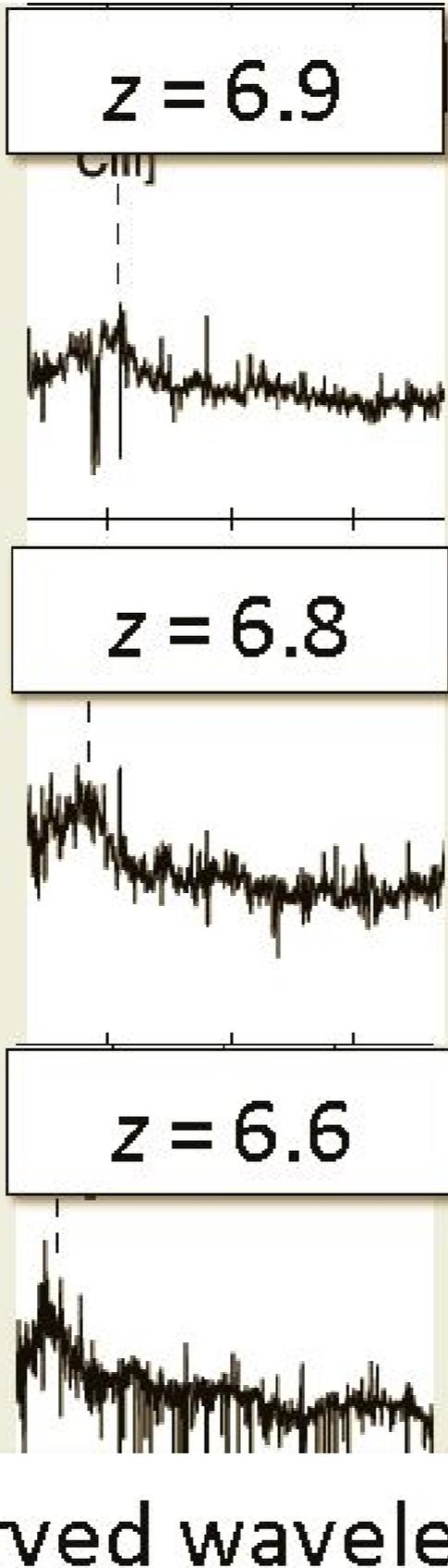
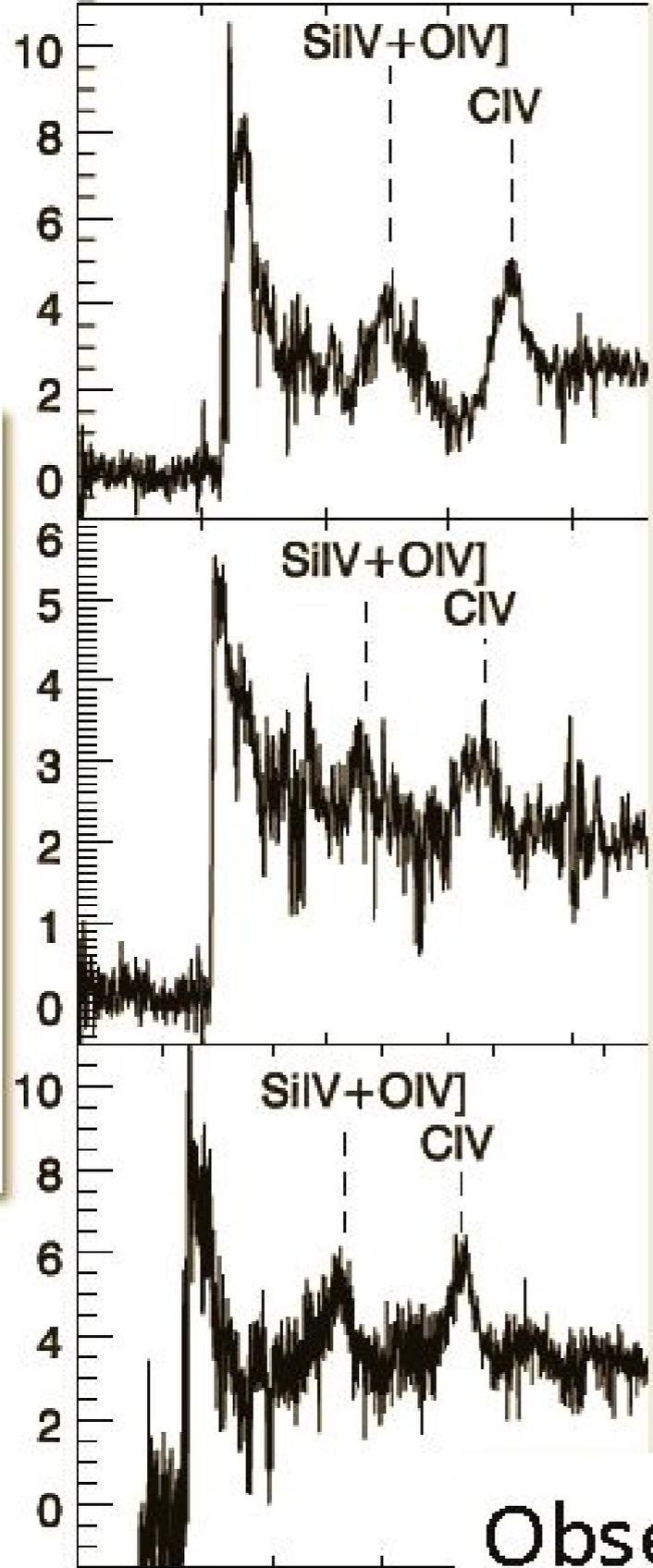




SMBH masses: careful spectral modelling



Flux density



Black hole mass:

$2.1 \times 10^9 M_{\odot}$

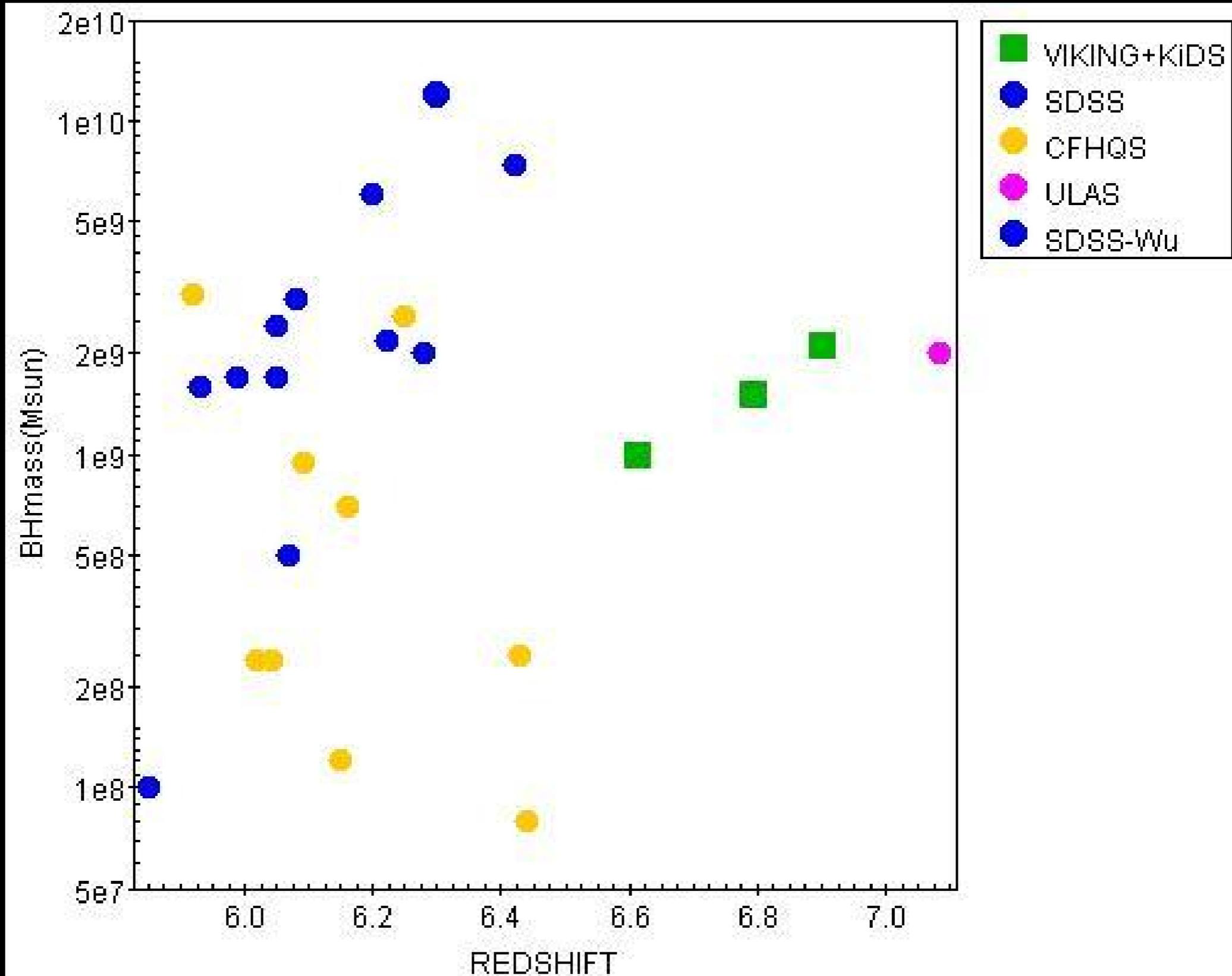
$1.5 \times 10^9 M_{\odot}$

$1.0 \times 10^9 M_{\odot}$

Observed wavelength

0.9 1.0 1.1 1.2 1.3 1.4 1.5 1.6 1.7 1.8 1.9 2.0
Observed wavelength [μm]

e.g., De Rosa+ 2014

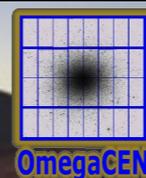


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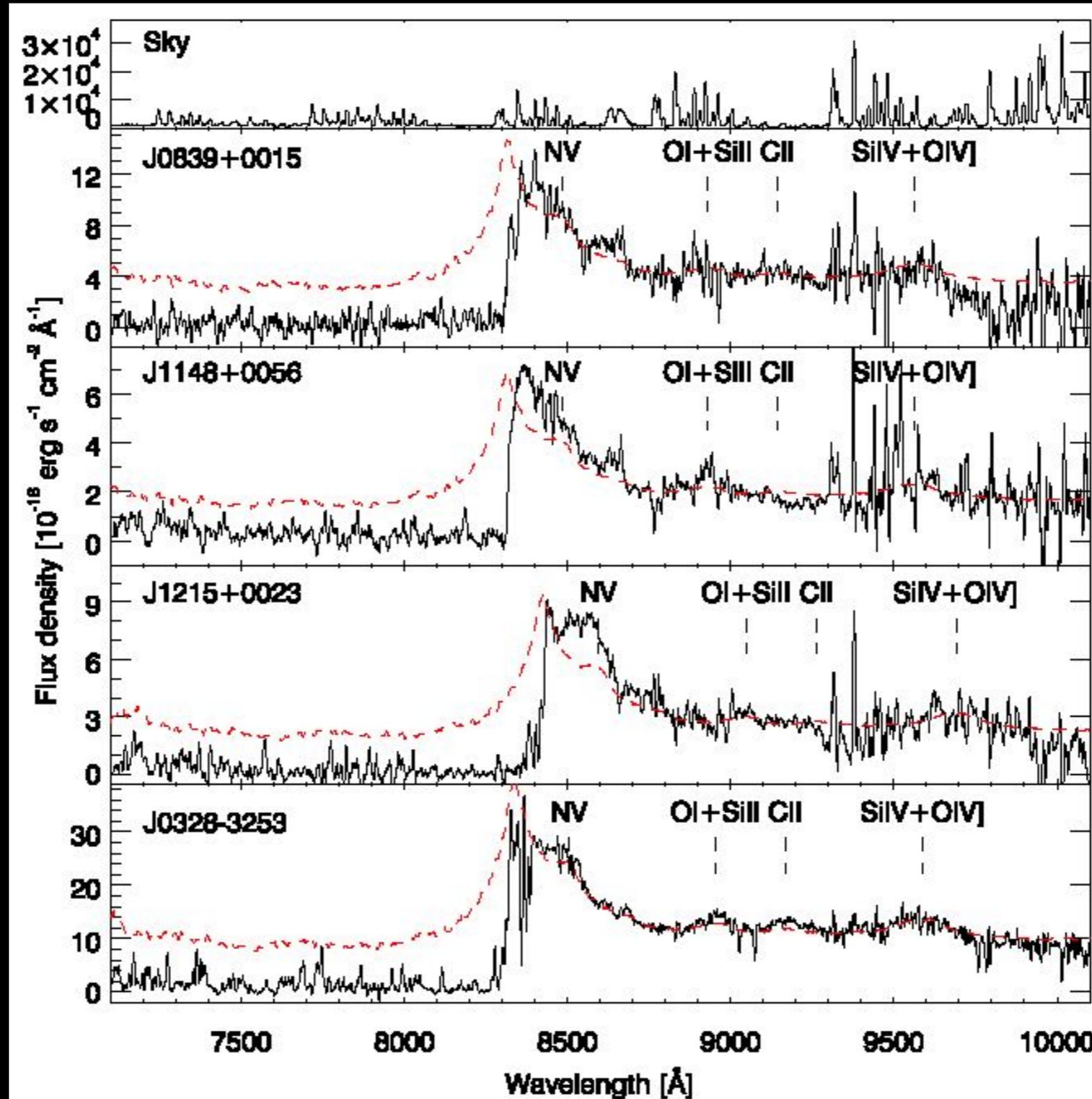
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Mostly weak Ly α QSOs at z~6



Venemans, GVK+2015

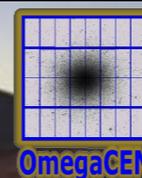


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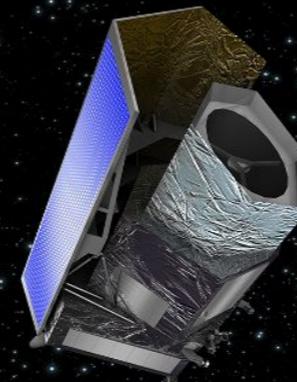
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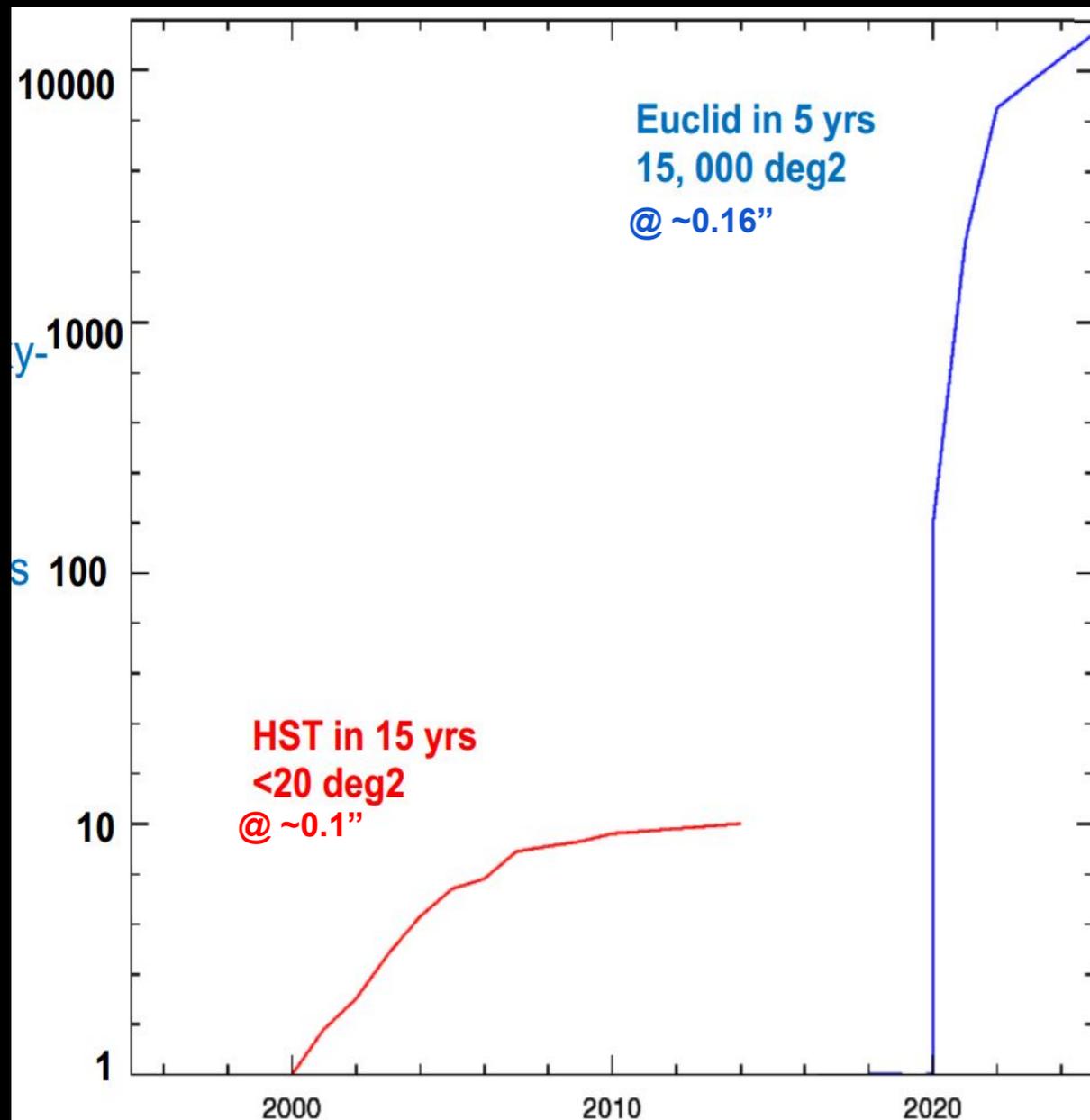


KiDS → Euclid

1 → 10



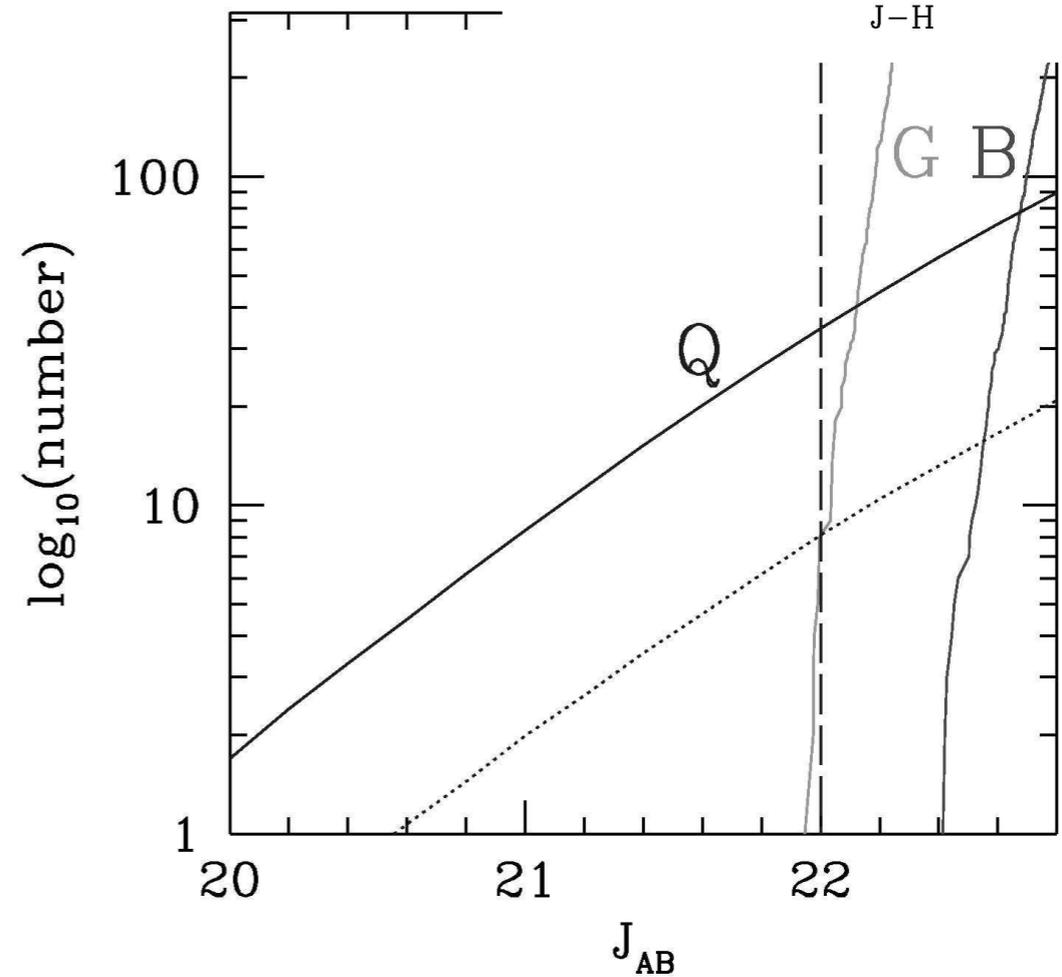
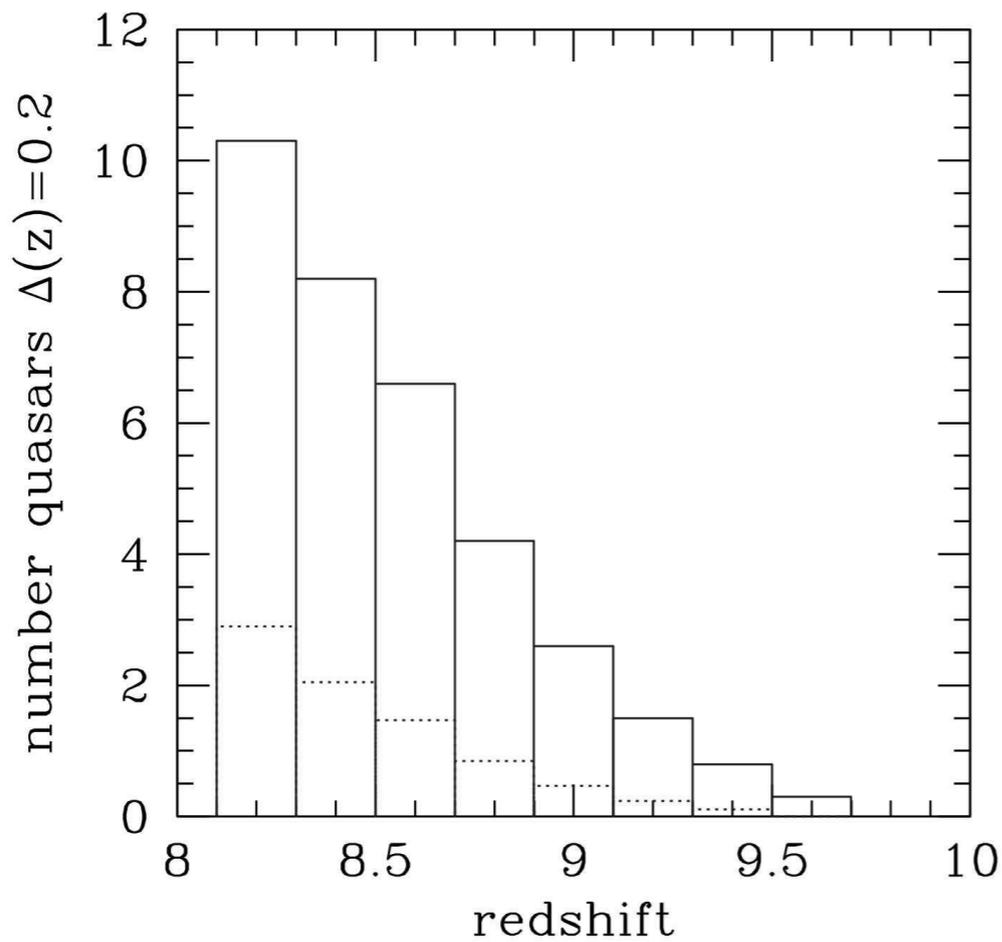
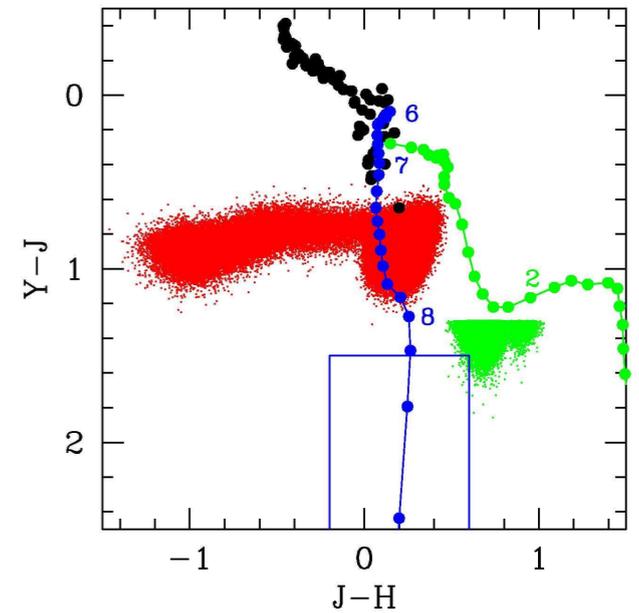
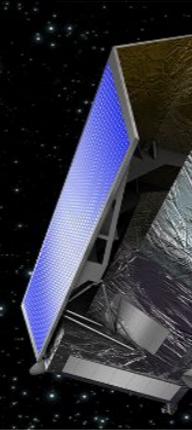
M1=1.2m
 VIS / Y,J,H, NIR-spectra
 5 sigma (AB) ~25 / 24.5
 Launch 2020



What	KiDS+VIKING	Euclid
FoM Dark Energy		~1500
Galaxies at $1 < z < 3$ with good mass estimates	$\sim 10^7$	$\sim 2 \times 10^8$
Massive galaxies ($1 < z < 3$) w/spectra	--	$\sim \text{few} \times 10^3$
H α emitters/metal abundance in $z \sim 2-3$	--	$\sim 4 \times 10^7 / 10^4$
Galaxies in massive clusters at $z > 1$	$\sim 10^3?$	$\sim 2 \times 10^4$
Type 2 AGN ($0.7 < z < 2$)	$< 10^3?$	$\sim 10^4$
Dwarf galaxies	?	$\sim 10^5$
$T_{\text{eff}} \sim 400\text{K}$ Y dwarfs	few	$\sim \text{few} \times 10^2$
Strongly lensed galaxy-scale lenses	~ 2000	$\sim 300,000$
$z > 6.5$ QSOs	~ 10	> 30

KiDS \rightarrow Euclid

1 \rightarrow 10



Courtesy Steve Warren and QSO Science Working Group

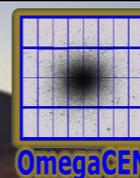


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Conclusions

VIKING + KiDS, a homogeneously selected sample,

- probes up to $z \sim 7.4$
- and down to $M_{1450} = -24$

Redshifts	#QSO	Brightnesses
$5.80 \leq z \leq 6.34$	11	$-26.7 \leq M_{1450} \leq -24.5$
$6.44 \leq z \leq 6.90$	5	$-26.1 \leq M_{1450} \leq -25.4$
TOTAL	16	~75% harvested

- found only weak Ly α QSOs at $z \sim 6$
- hint of drop in QSO density $z > 6.5$
- observed $\rho \sim 0.4 \text{ Gpc}^{-3}$ for $\text{SMBH} = 10^9 M_{\text{sun}}$ at $z > 6.5$
- The QSO future looks bright with Euclid: 2020
 - many LLQSO, many $z > 8$ expected

