



UNIVERSITY OF  
CAMBRIDGE

# Constraining reionization with Lyman-alpha emitters and the CMB

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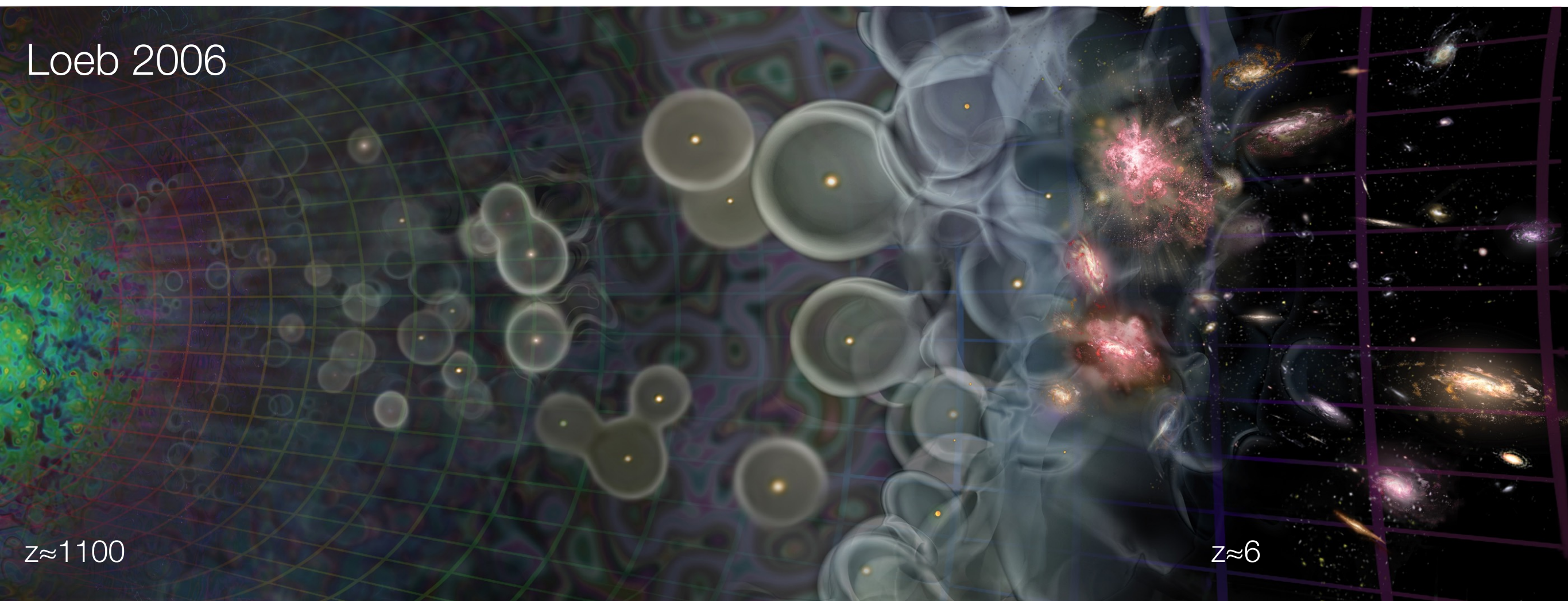
Ewald Puchwein

KICC / IoA, University of Cambridge

collaborators: Tirthankar Choudhury, Martin Haehnelt, James Bolton

# CMB and the epoch of reionization

time



CMB

cosmic  
dark ages

first stars  
and galaxies

reionization

today

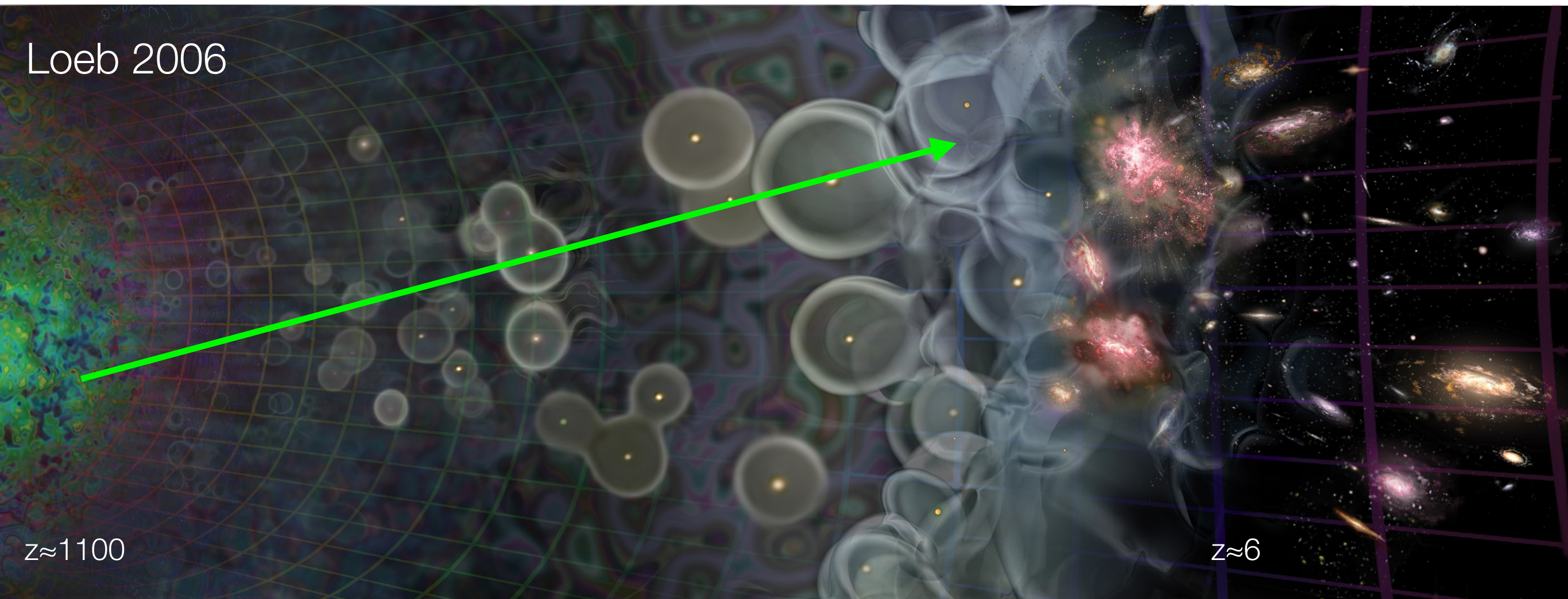


# CMB and the epoch of reionization

time



Loeb 2006



CMB

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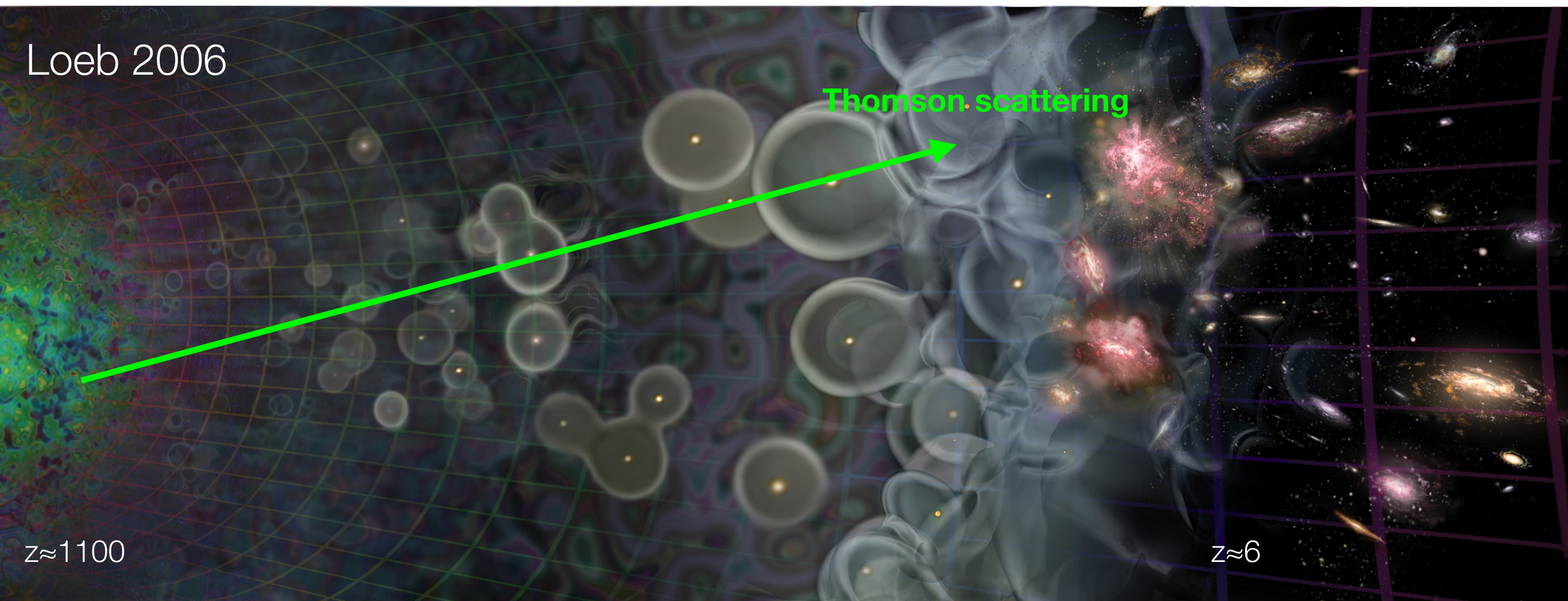
reionization

today



# CMB and the epoch of reionization

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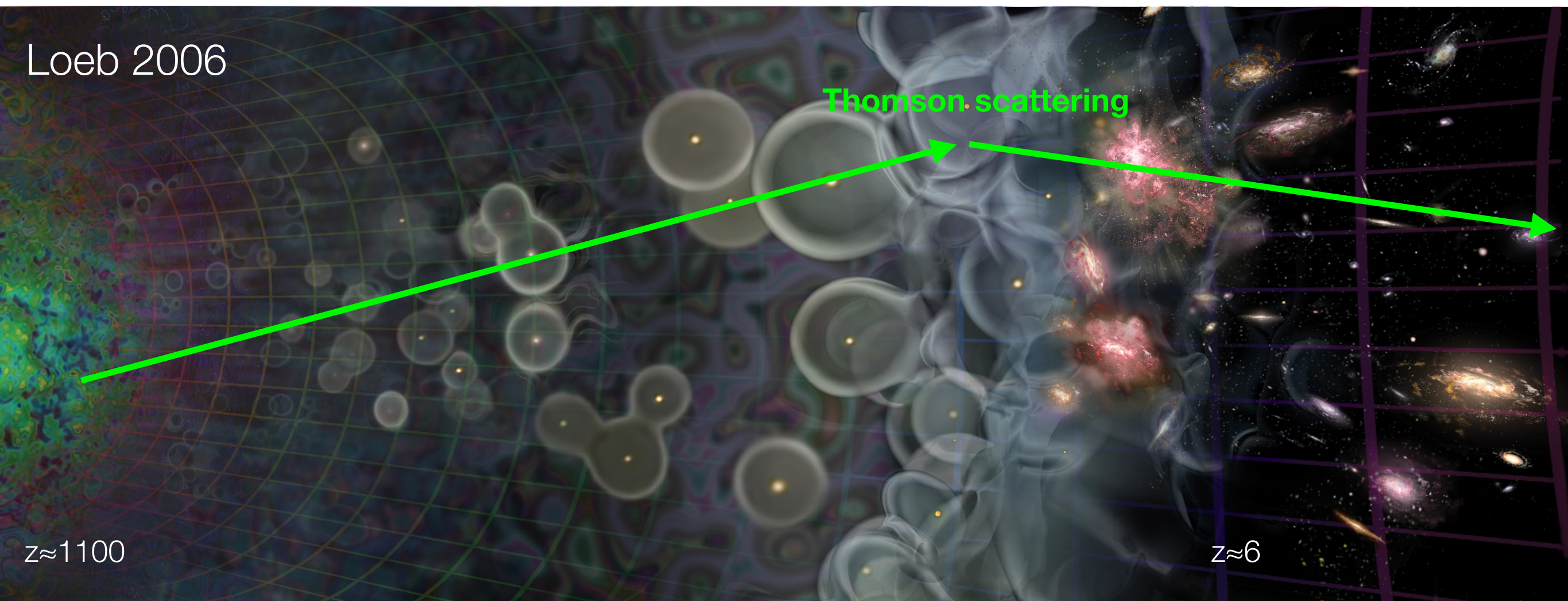


# CMB and the epoch of reionization

time



Loeb 2006



$z \approx 1100$

$z \approx 6$

CMB

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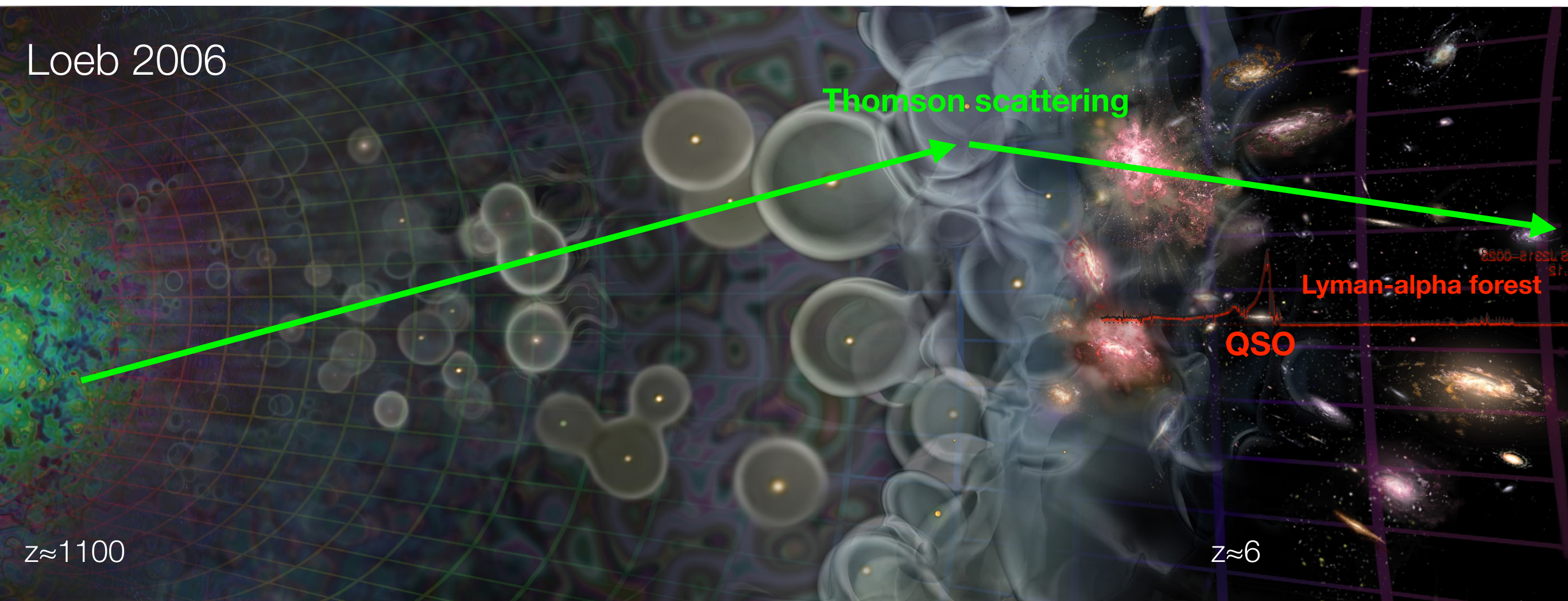


# CMB and the epoch of reionization

time



Loeb 2006



CMB

cosmic dark ages  
first stars and galaxies

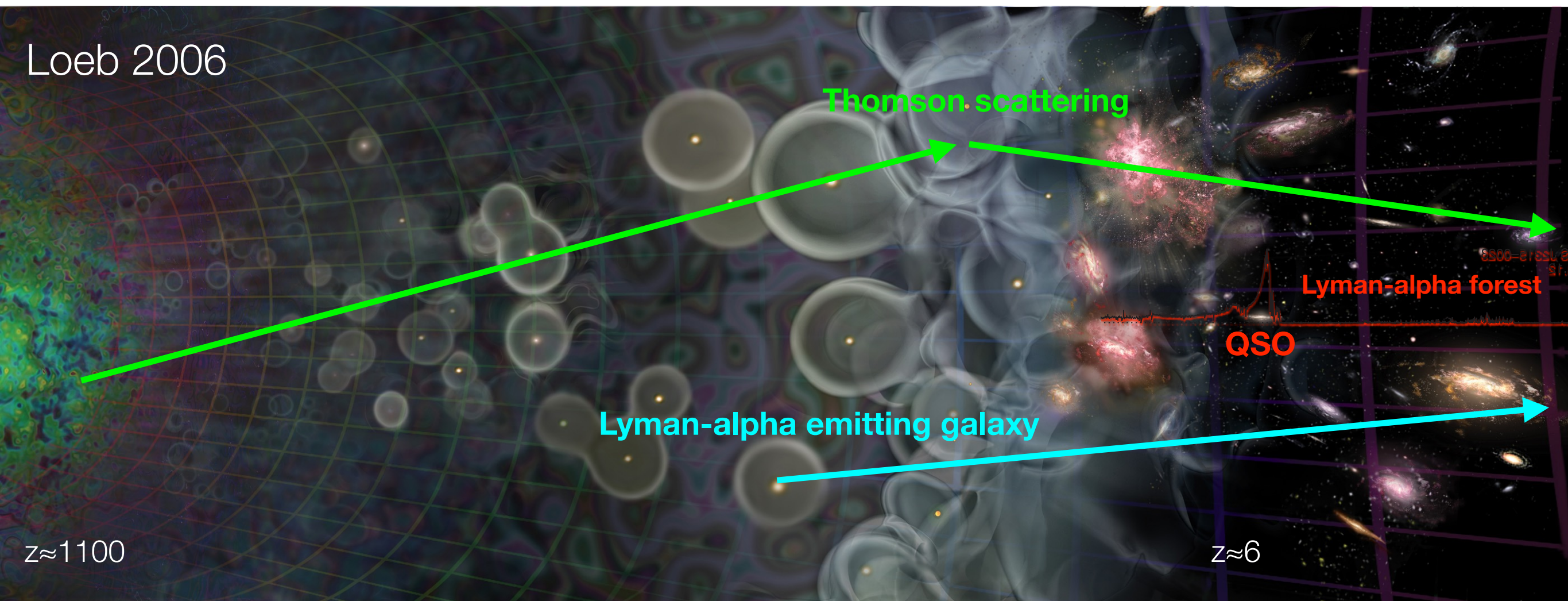
reionization

today



# CMB and the epoch of reionization

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CMB

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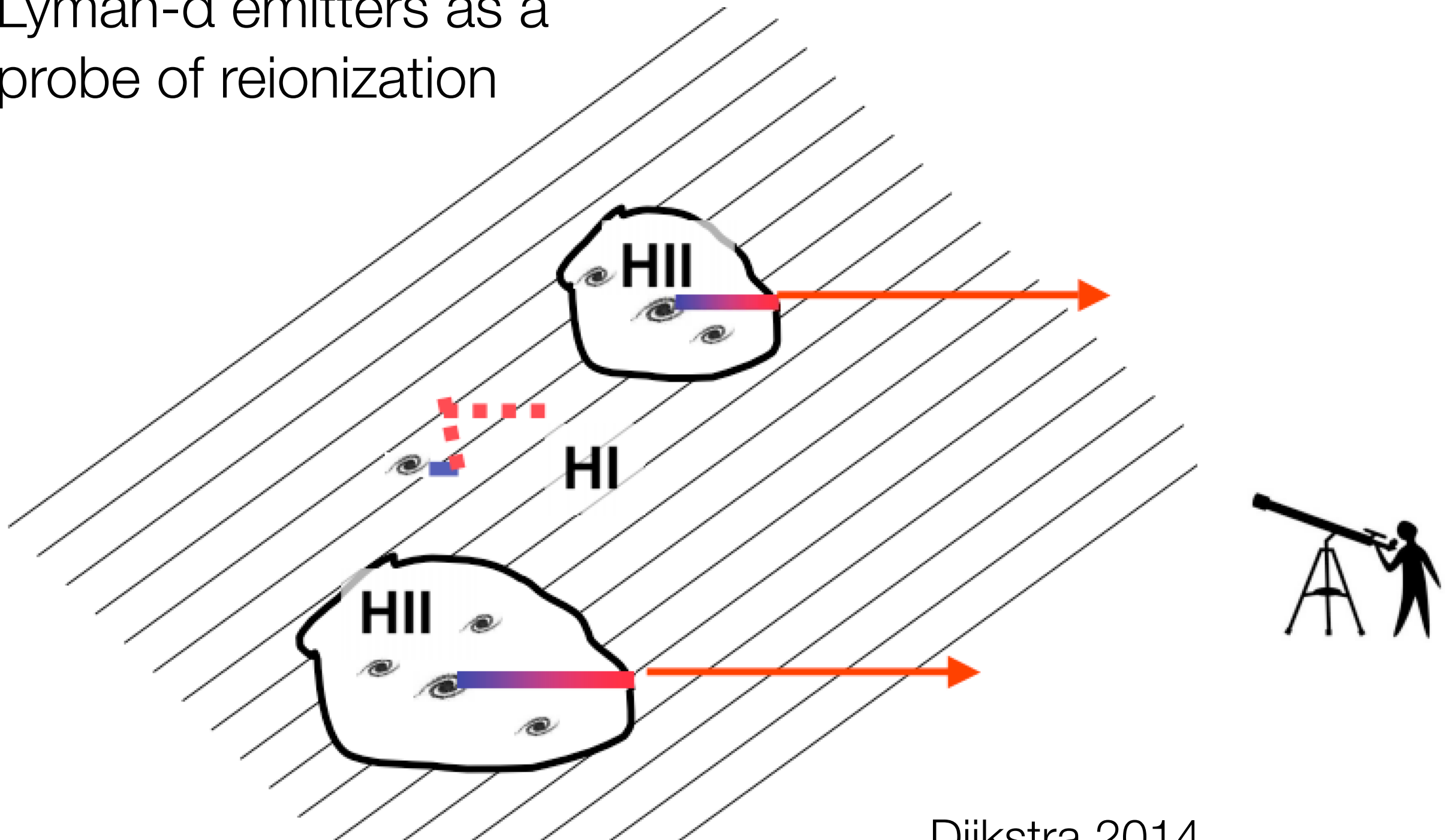
today



# Probing the ionization & thermal state of the IGM

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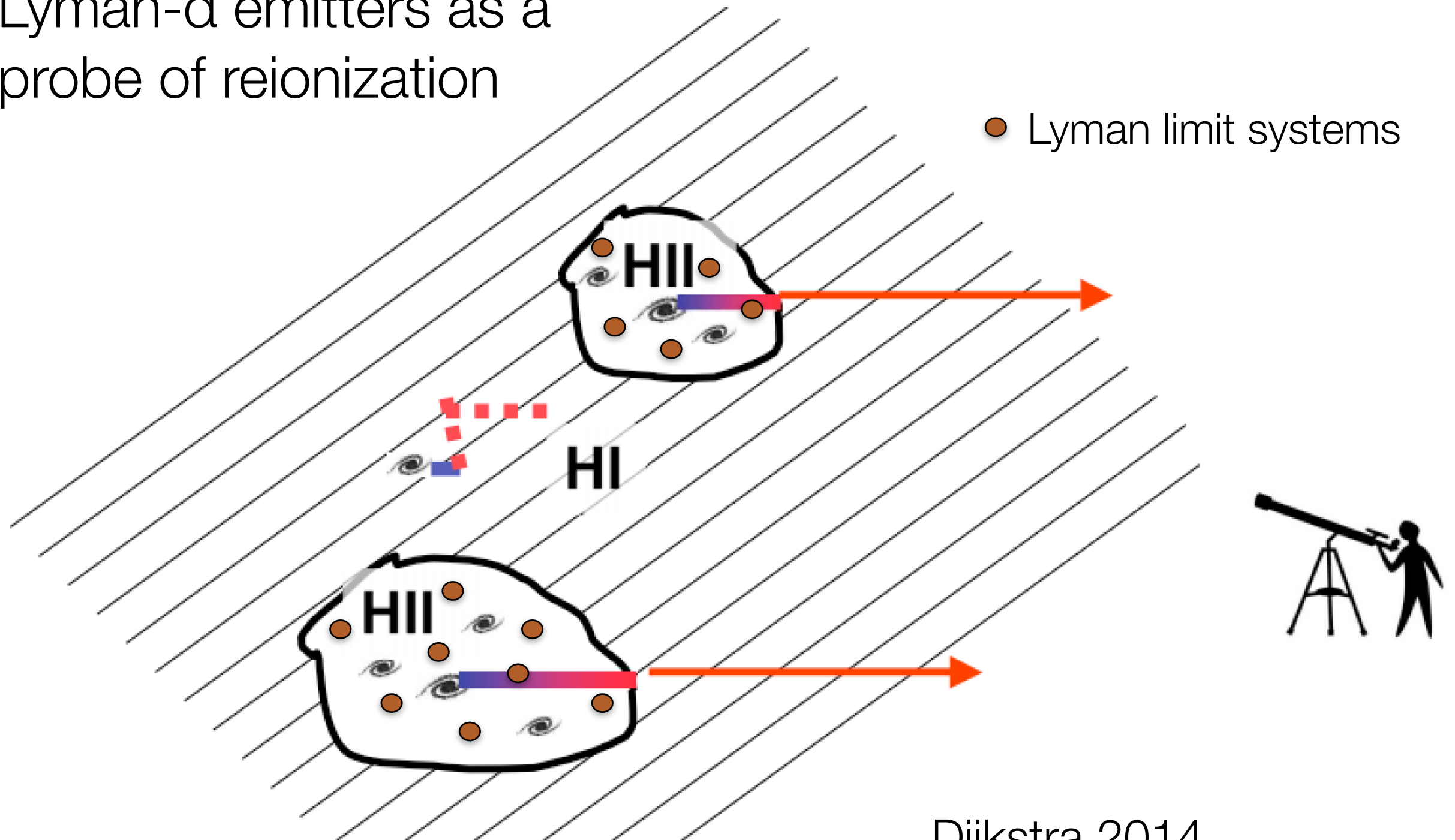
- Lyman- $\alpha$  emitters as a probe of reionization





# Probing the ionization & thermal state of the IGM

- Lyman- $\alpha$  emitters as a probe of reionization

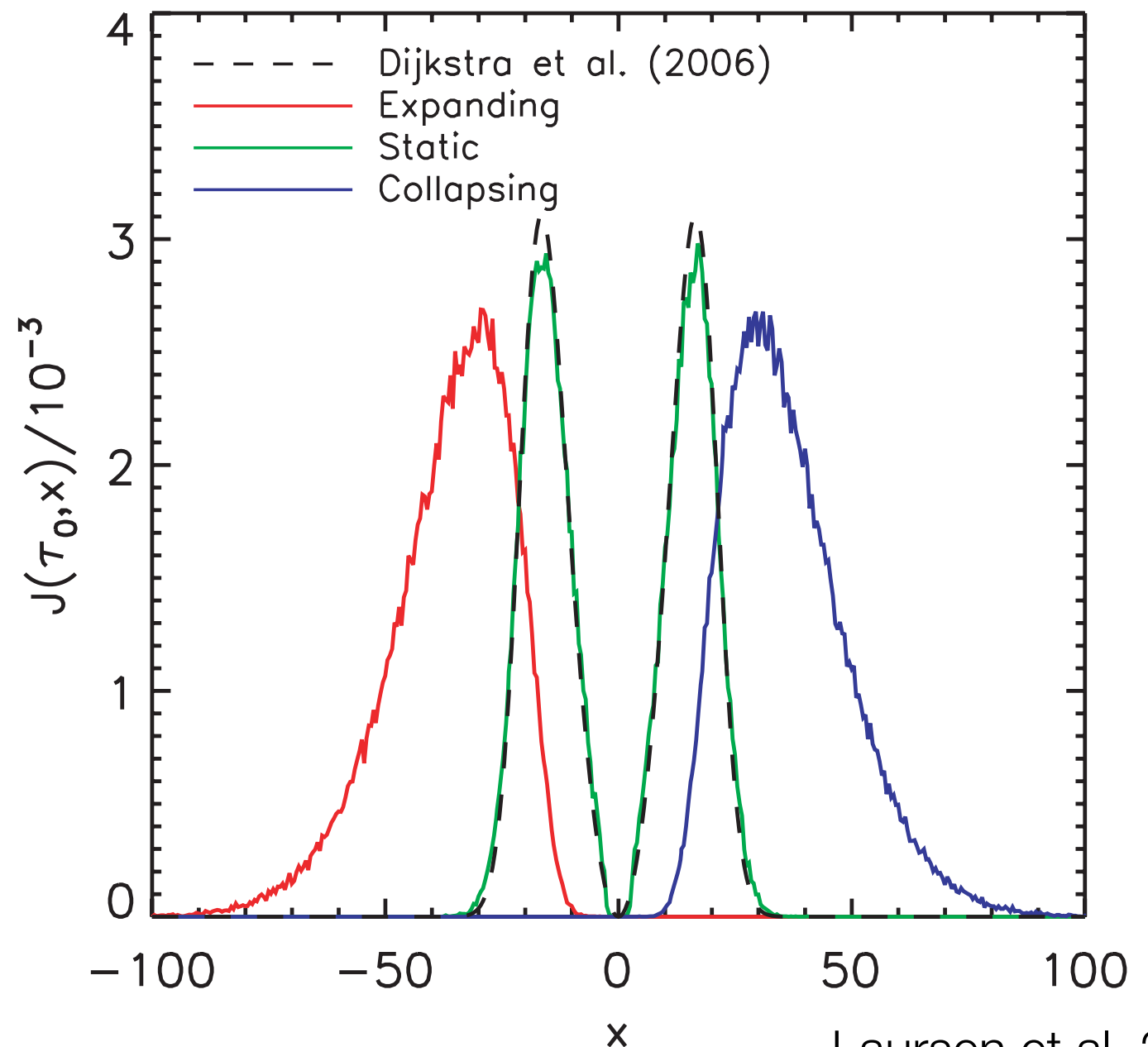




# Probing the ionization & thermal state of the IGM

frequency diffusion of Lyman- $\alpha$  photons

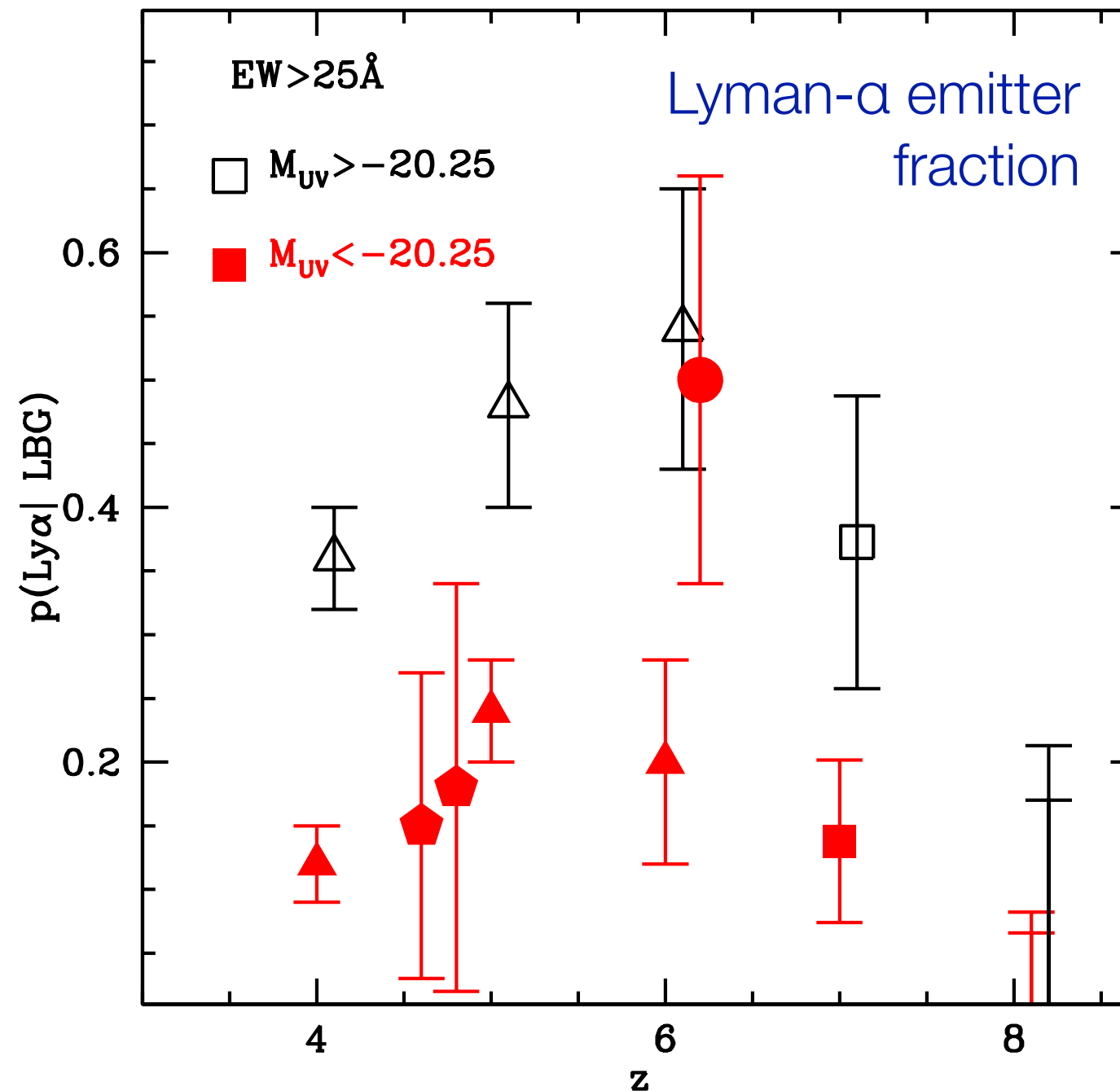
- Lyman- $\alpha$  emitters:
- line usually redshifted with respect to systemic velocity





# Probing the ionization & thermal state of the IGM

- Lyman- $\alpha$  emitters as a probe of reionization





# Semi-numerical models of reionization

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- semi-numerical models based on the ionizing photon budget, ionized if:

**ionizing photons - recombinations > number of hydrogen atoms**

- excursion set approach:

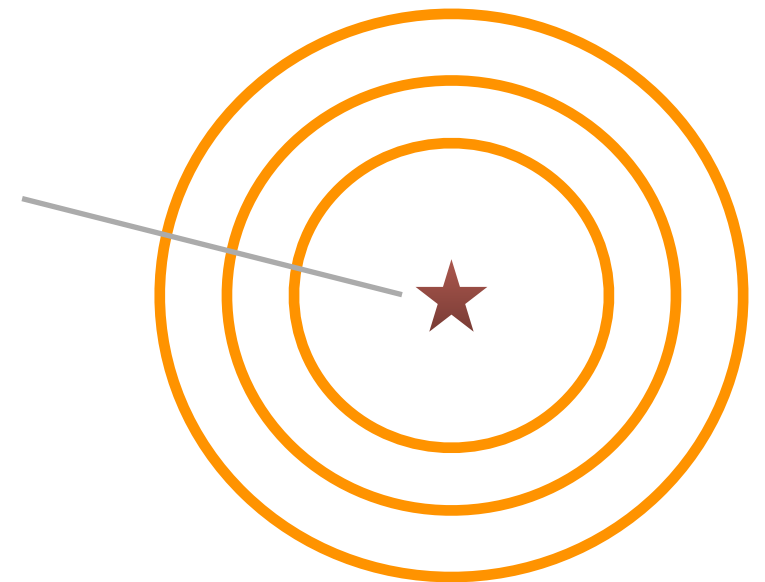
check for each point if there is any radius inside which the ionization condition is satisfied



not ionized



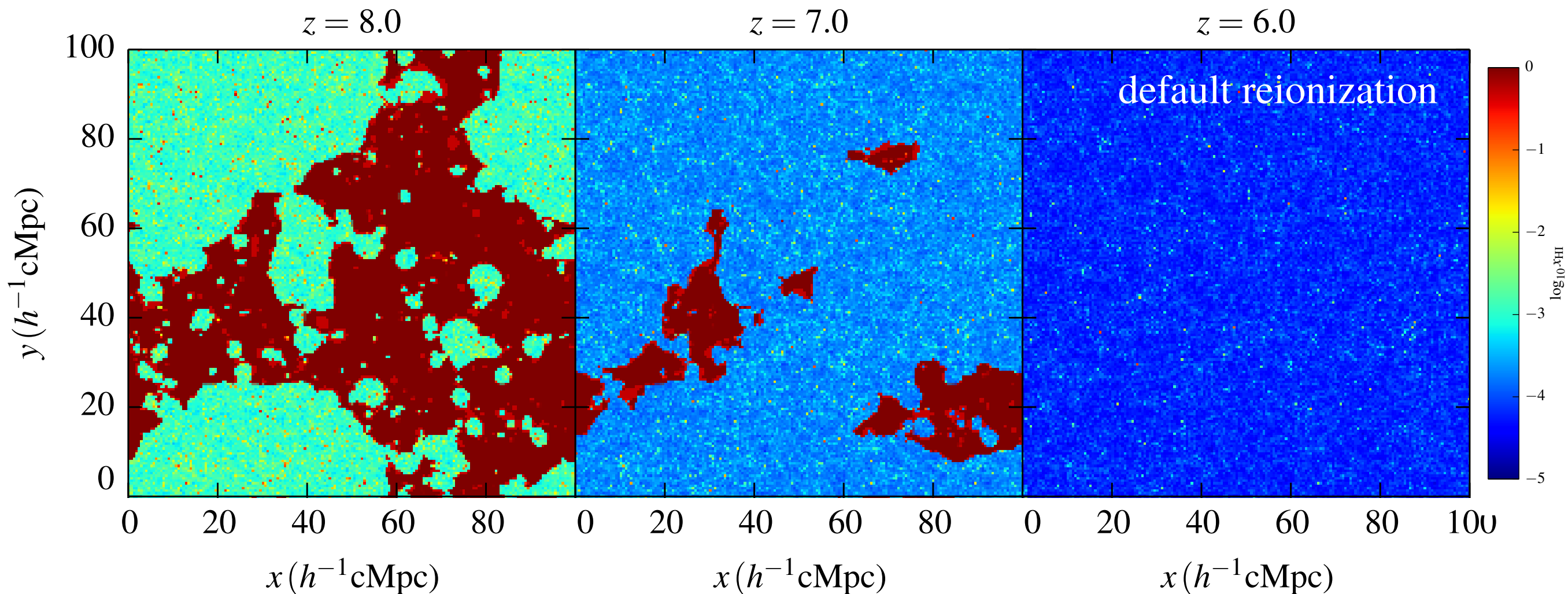
ionized





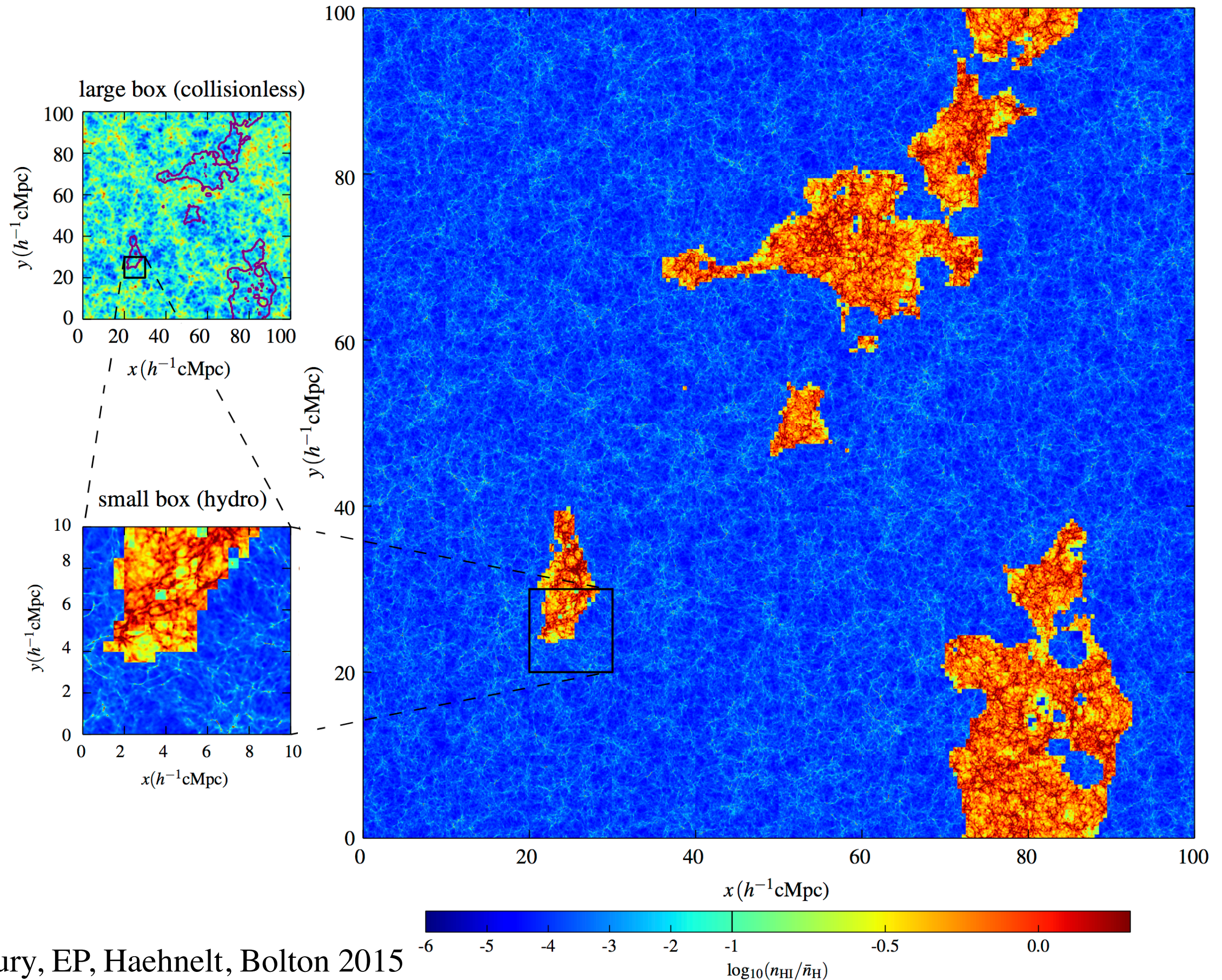
# Semi-numerical reionization in post-processing

- empirical assignment of ionizing luminosity to halos
  - > get ionized regions in post-processing from simulations



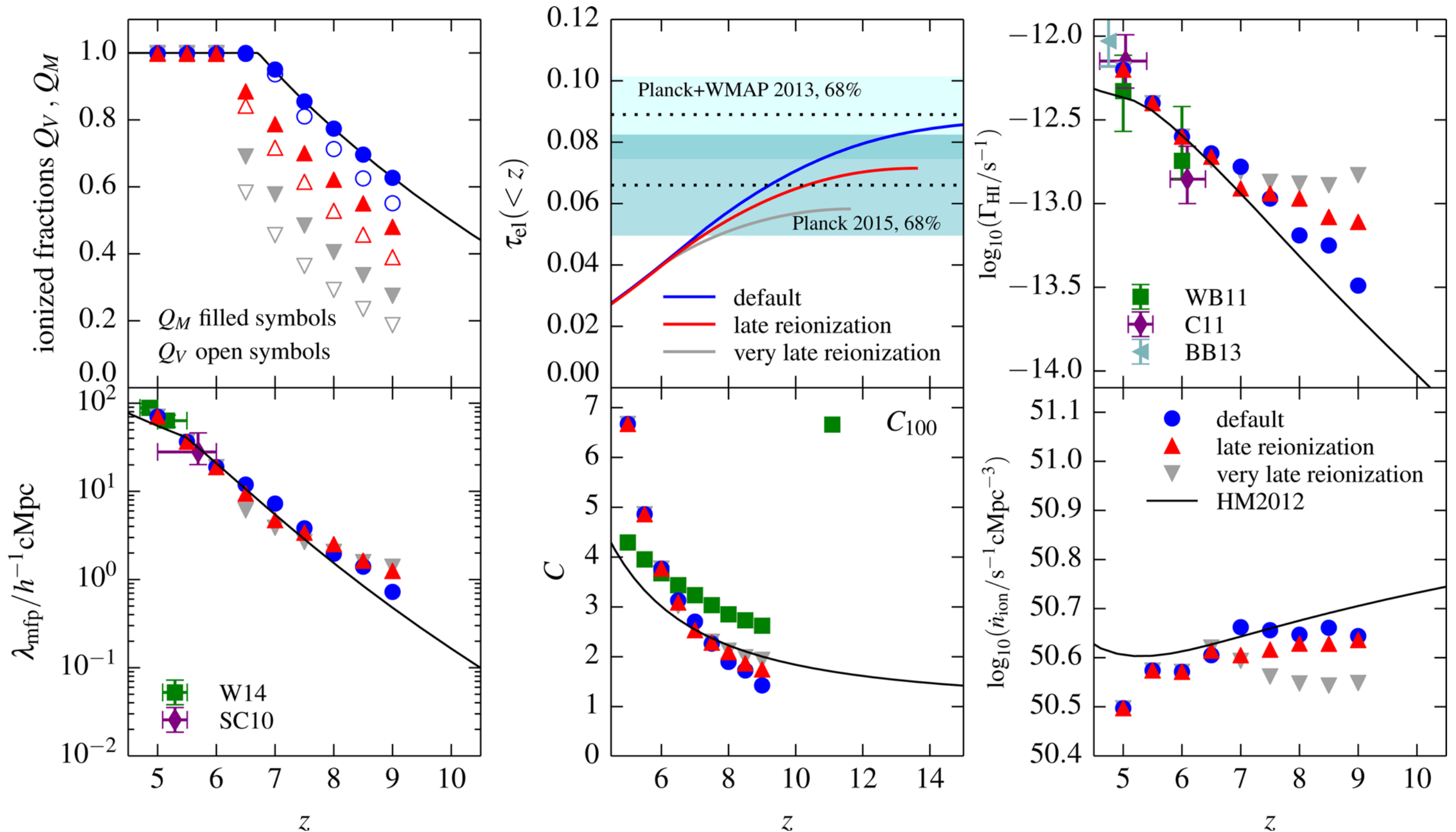


# hybrid technique combining large and small box hybrid box



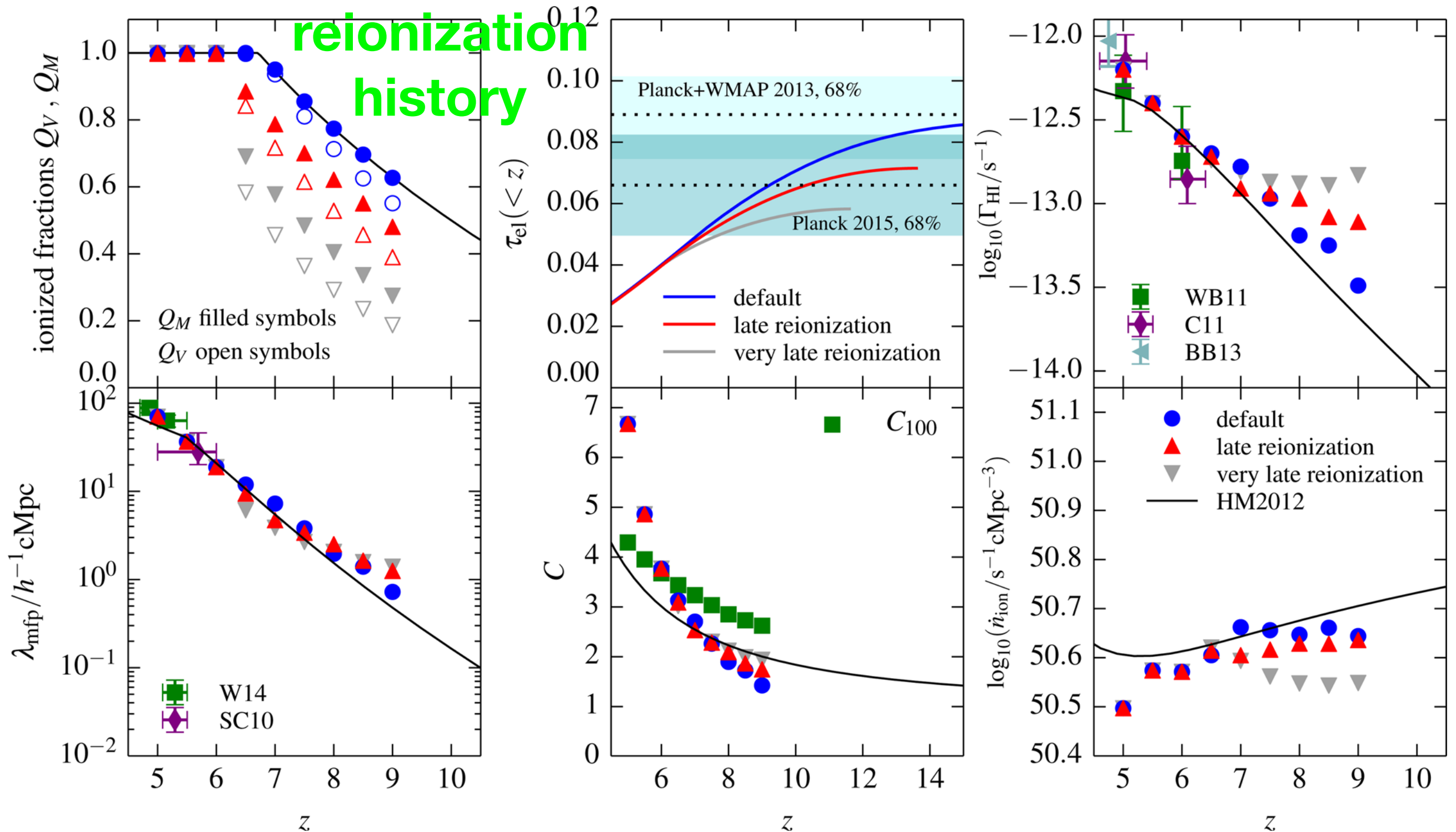


# Iterative calibration of the photoionization rate



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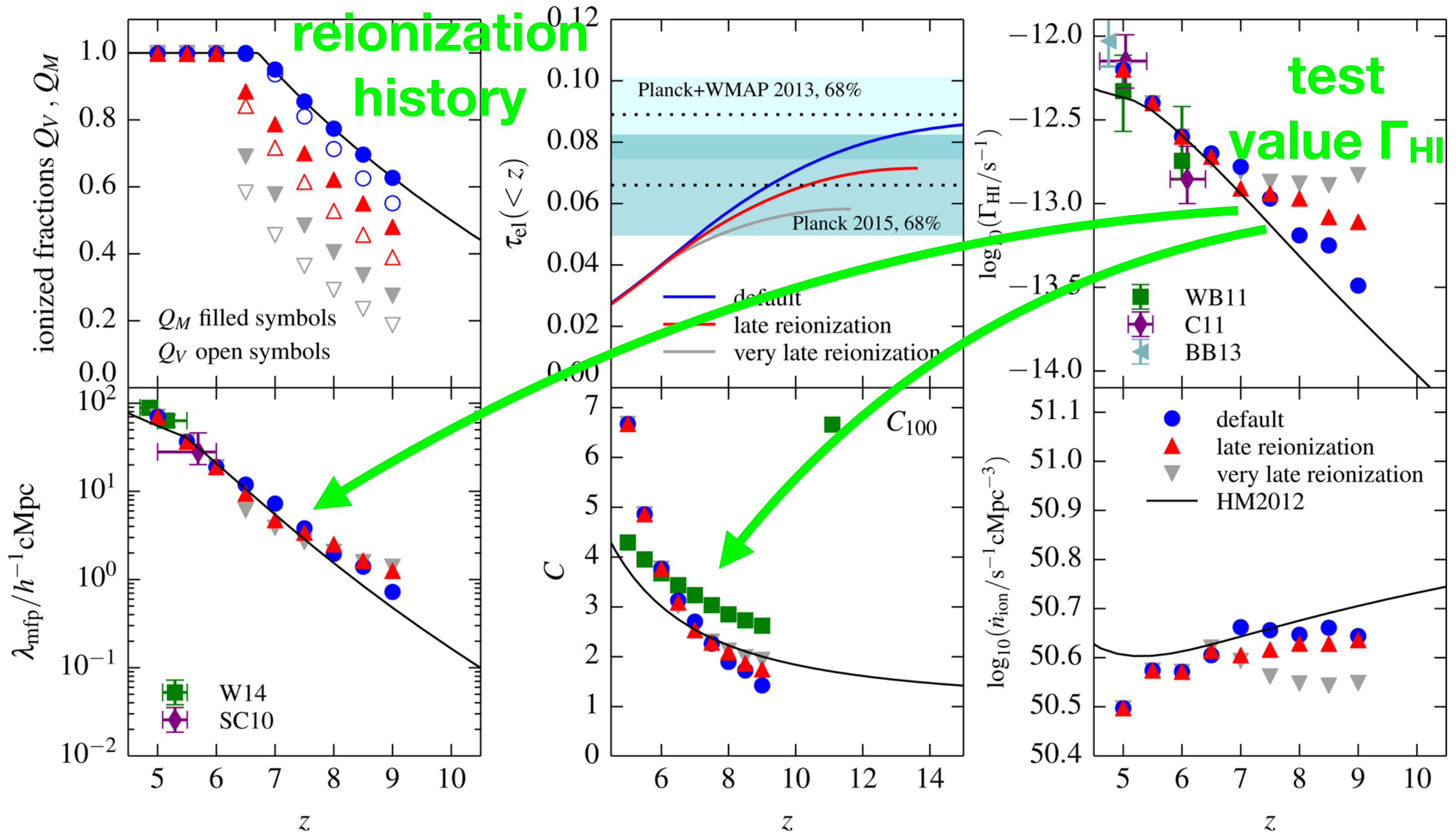
choose



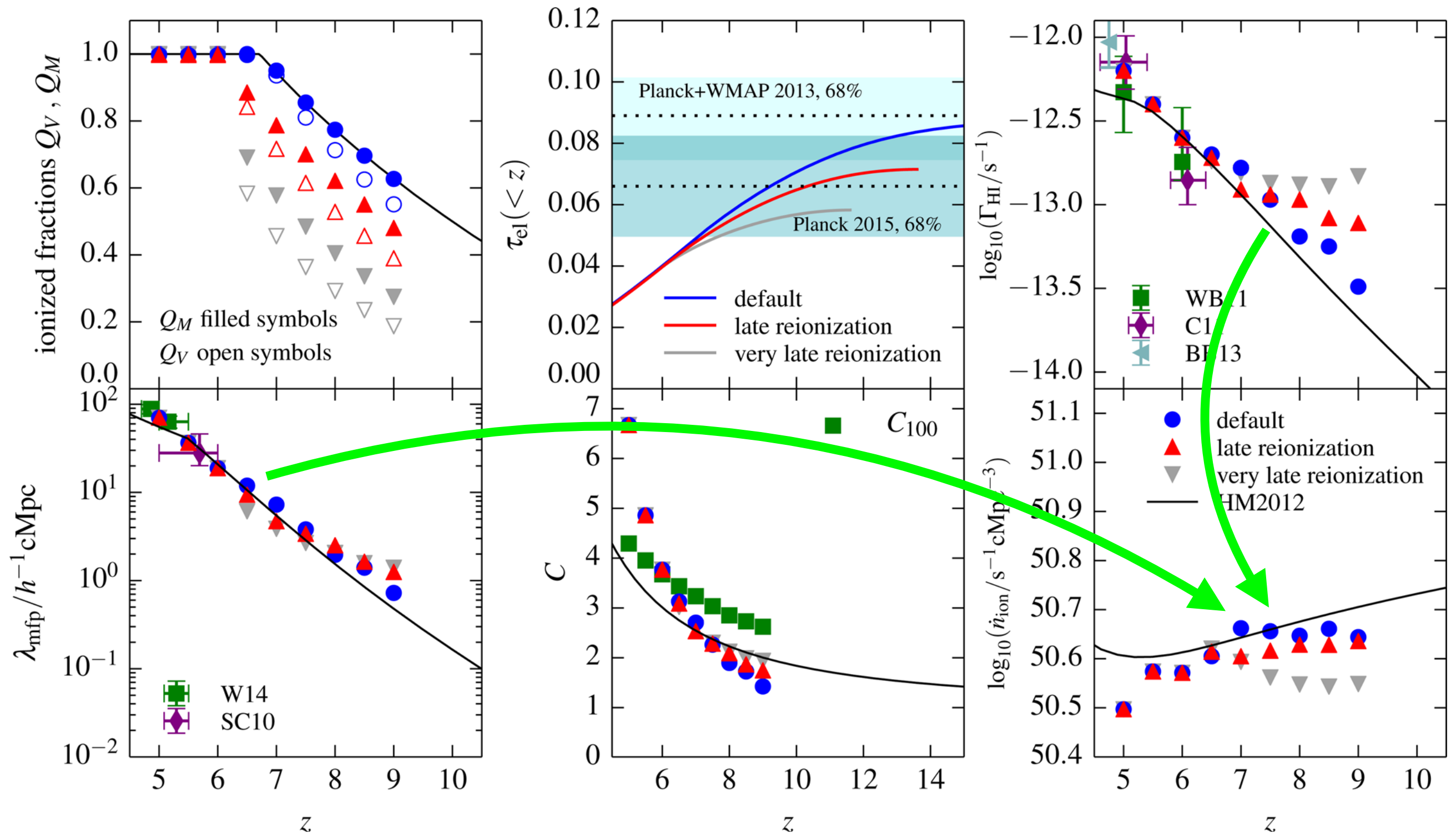


# Iterative calibration of the photoionization rate

choose

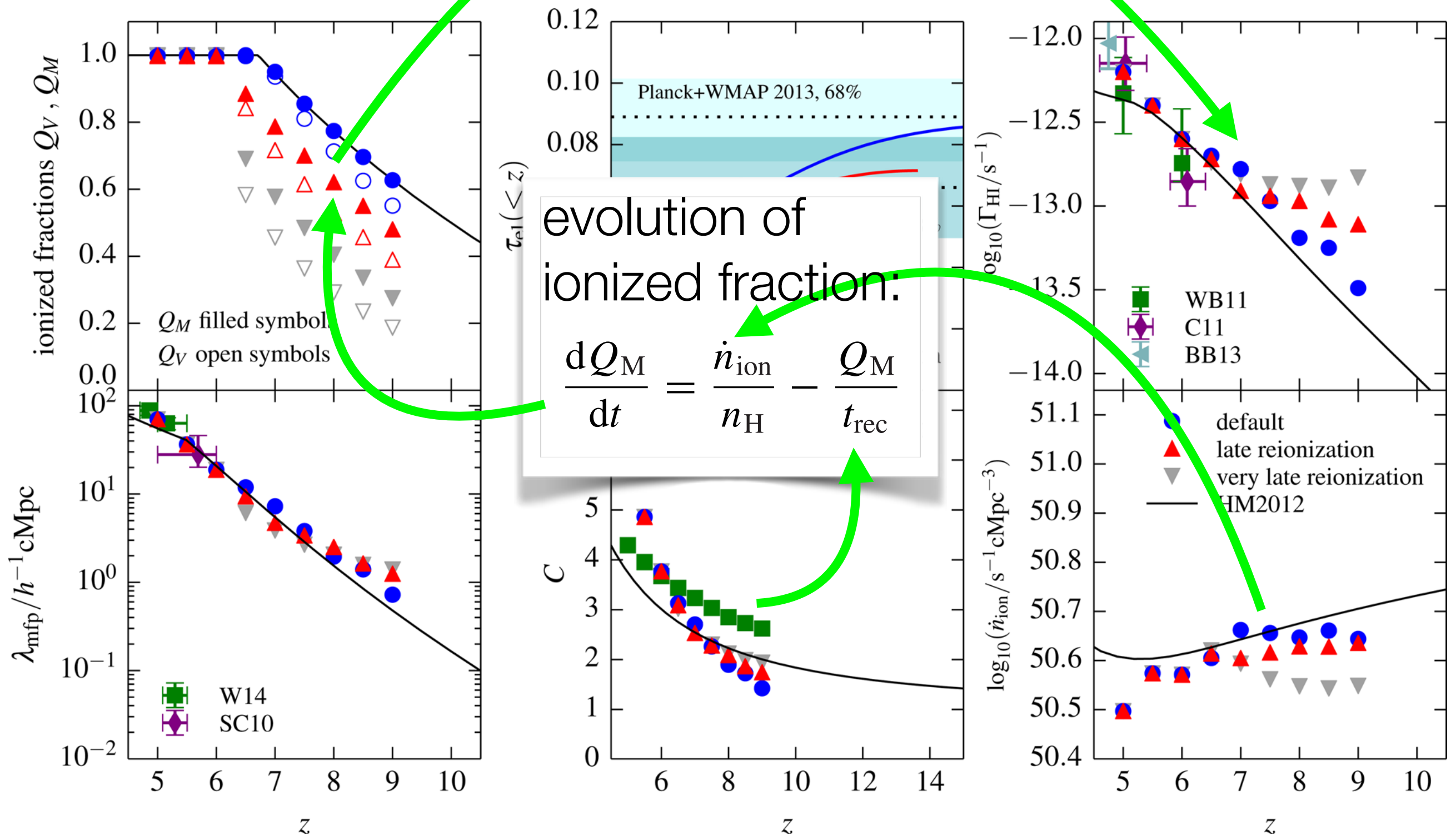


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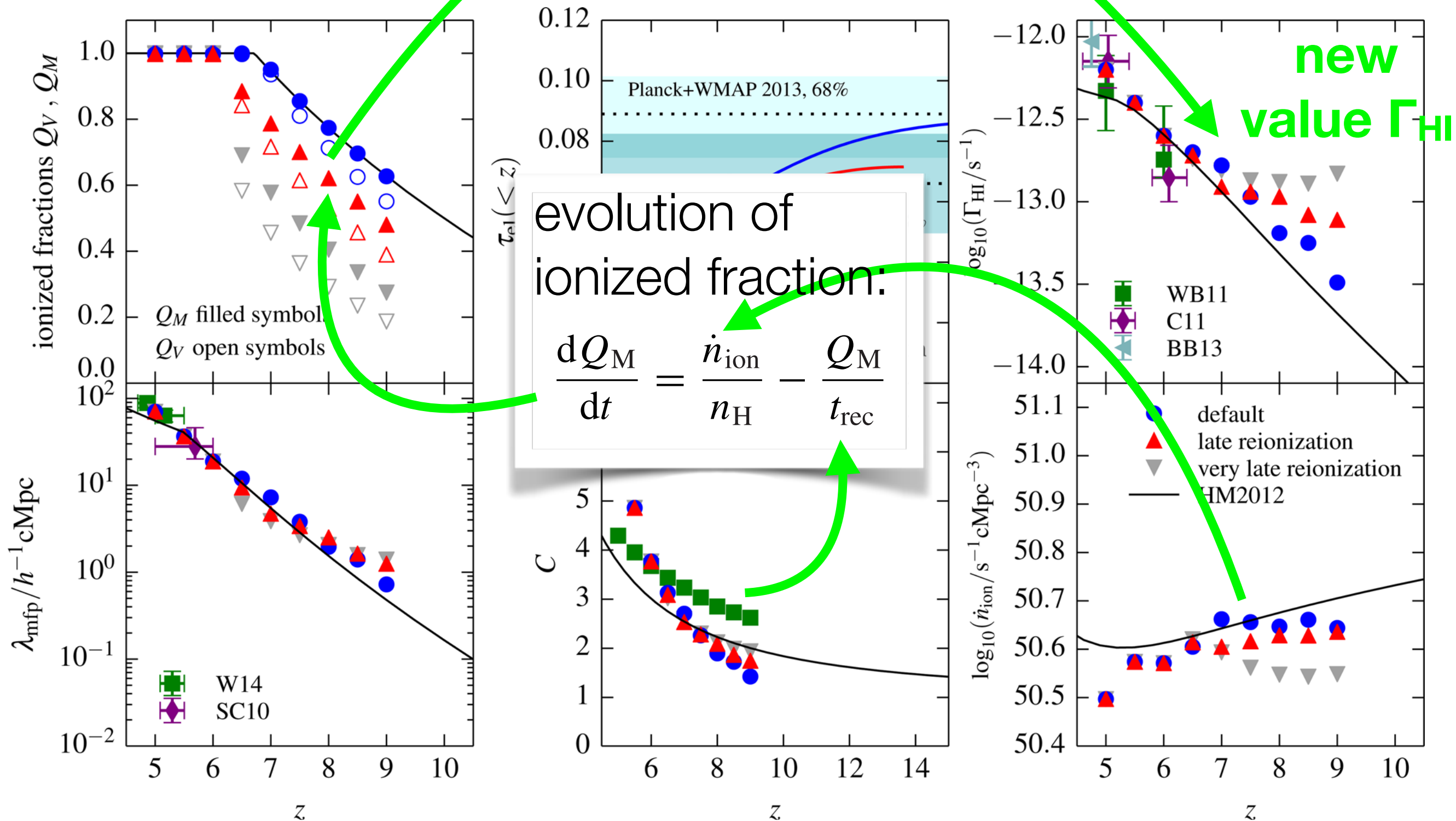




# Iterative calibration of the photoionization rate

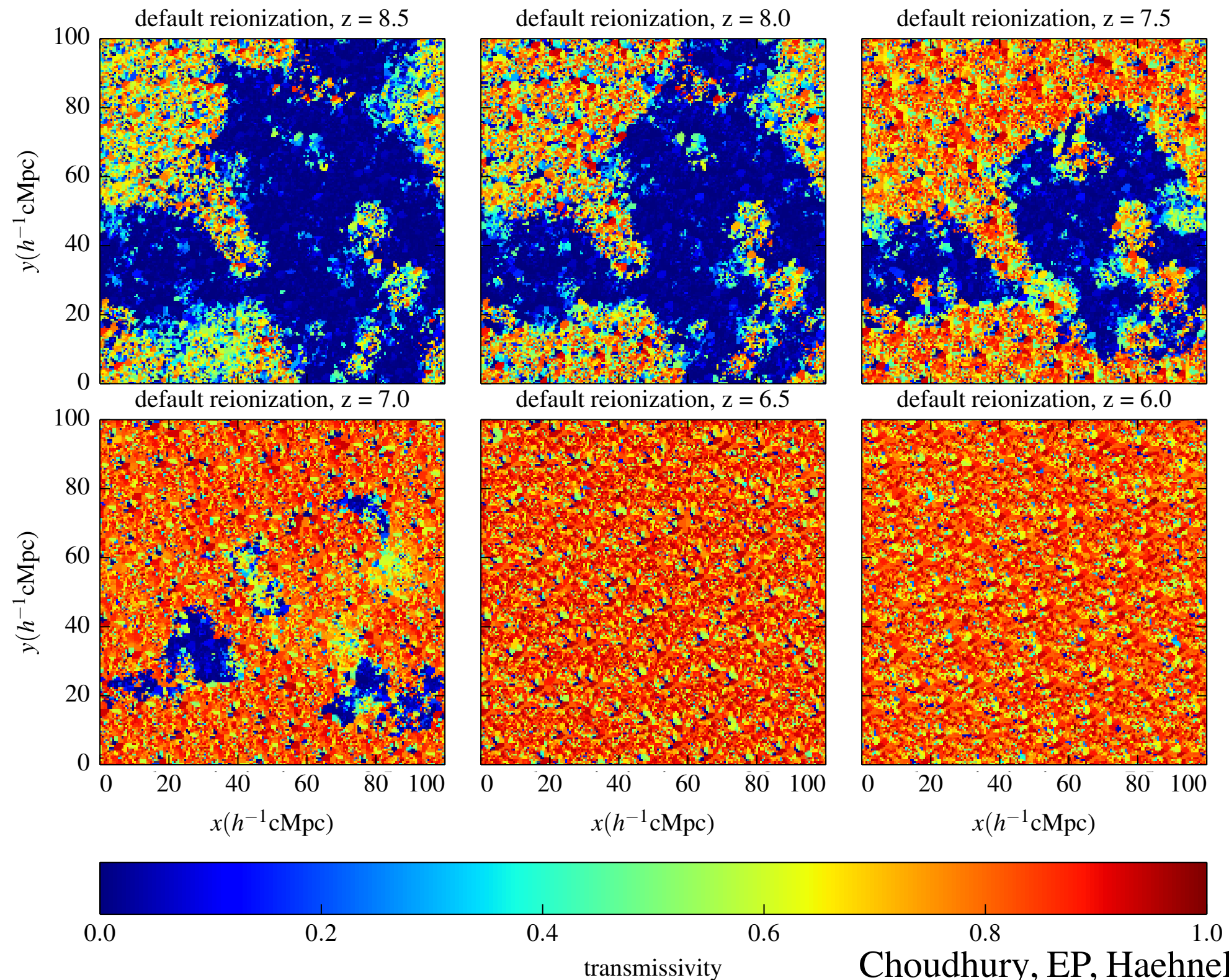


# Iterative calibration of the photoionization rate

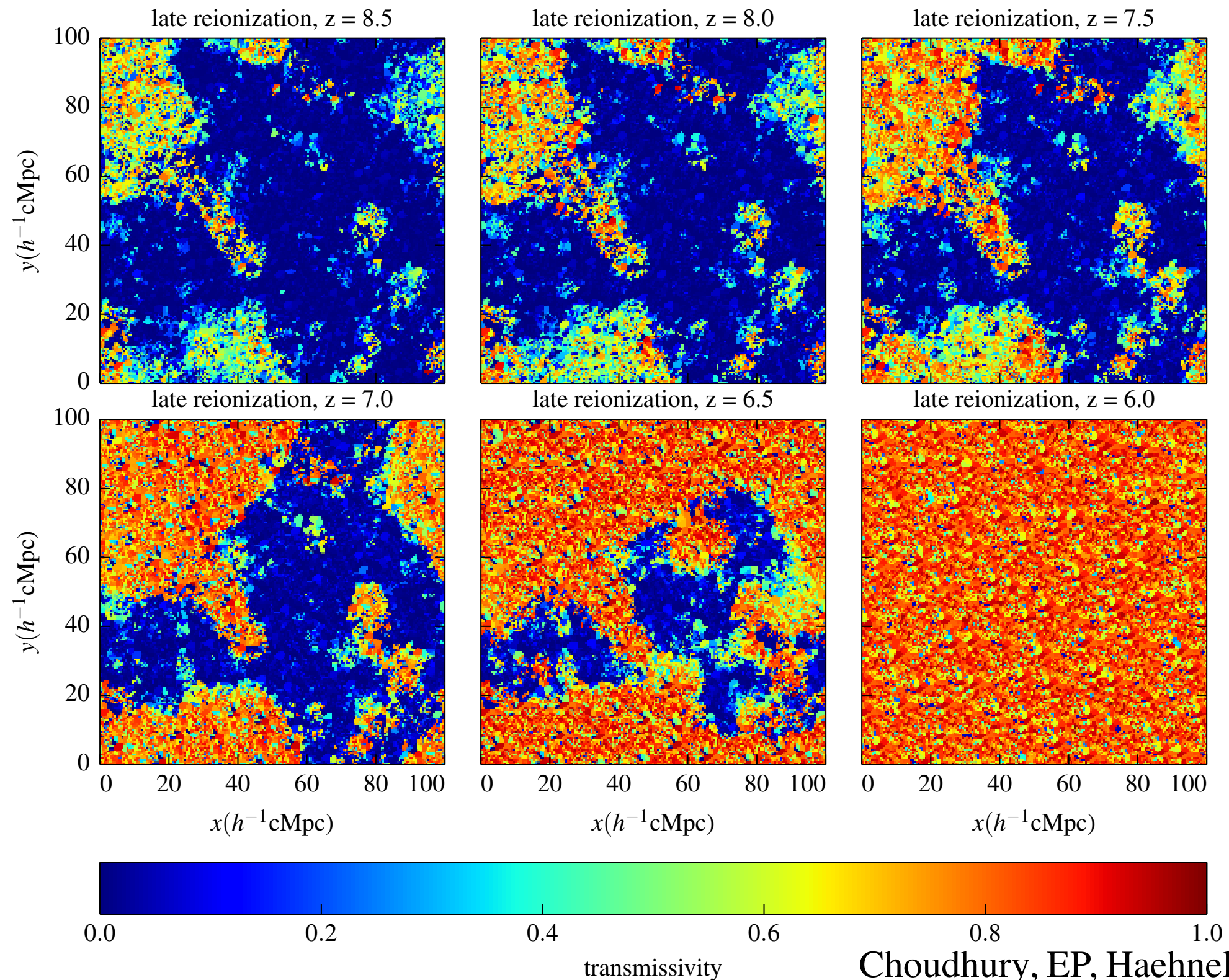




# Transmissivity for Lyman- $\alpha$ emission lines



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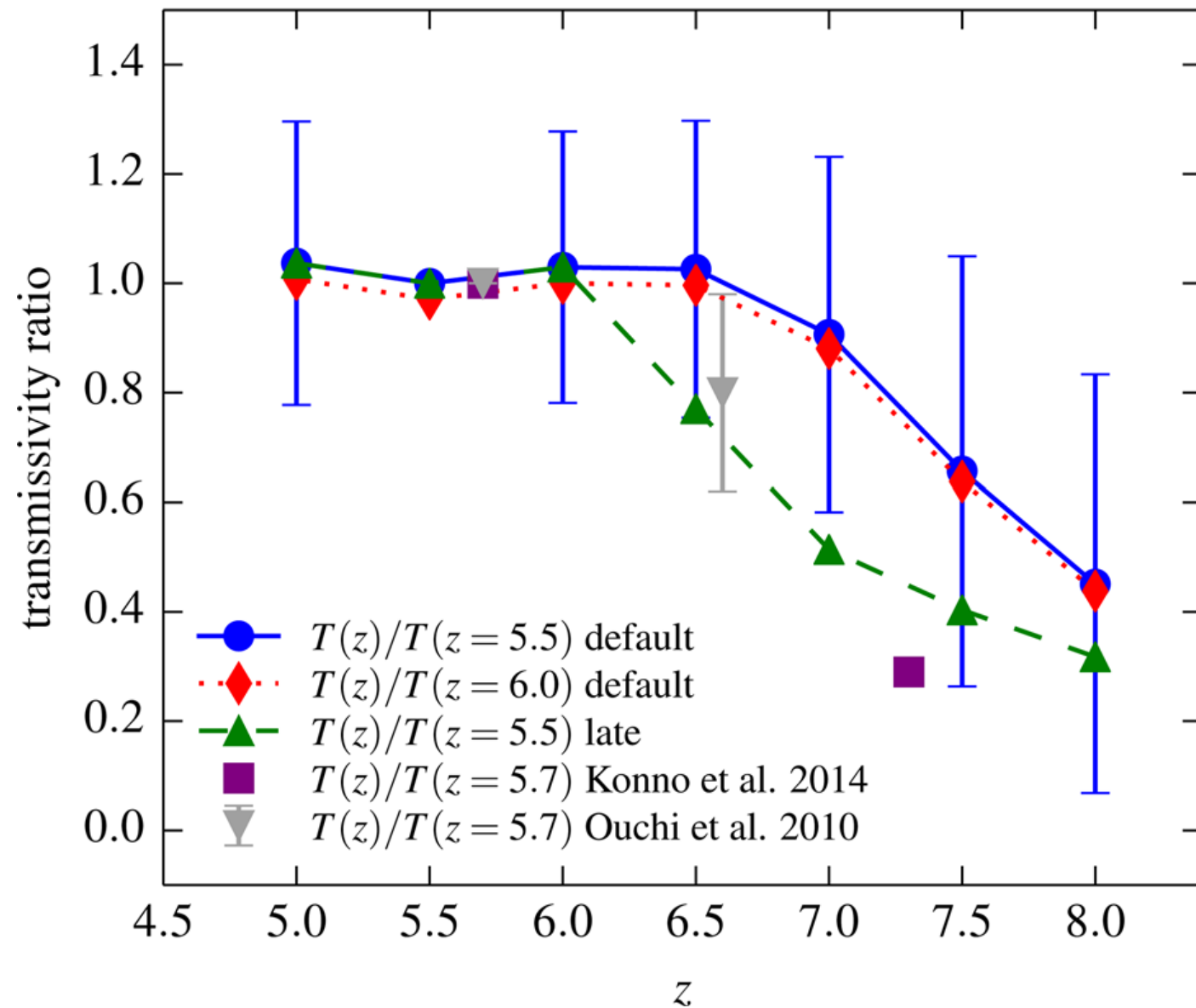




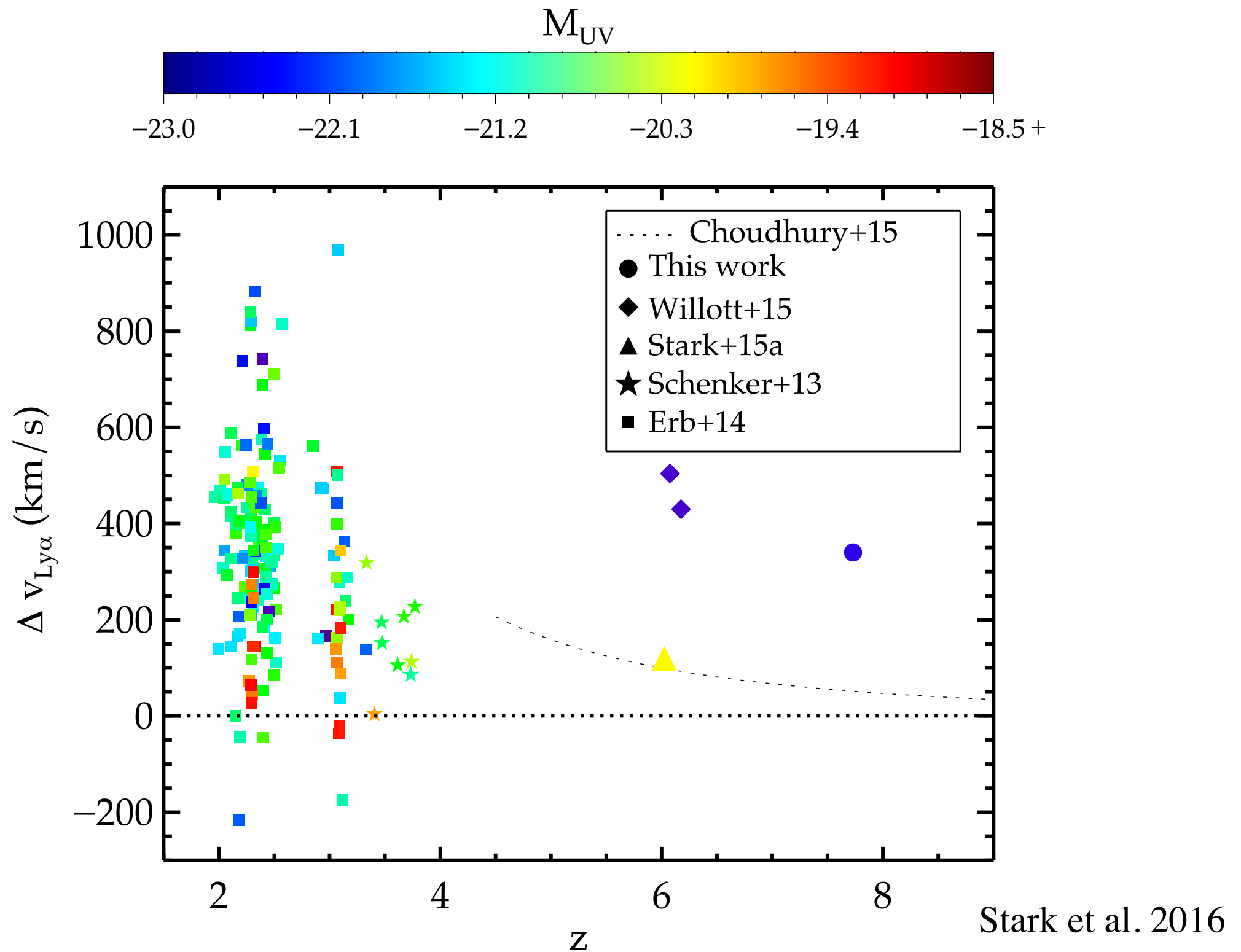
# Redshift evolution of the transmissivity

constant velocity  
shift

$$\Delta v_{\text{int}} = 100 \text{ km s}^{-1}$$



# Intrinsic velocity shift vs redshift and luminosity





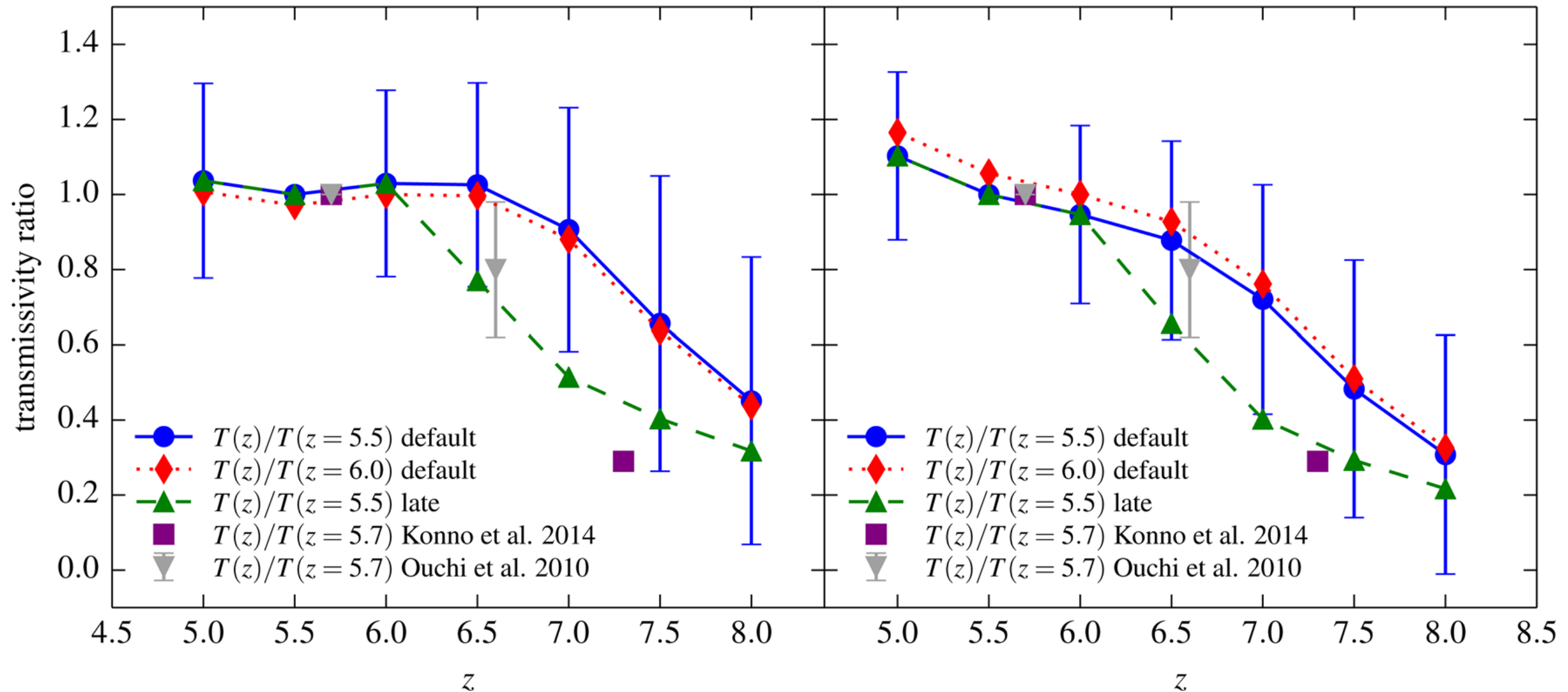
# Redshift evolution of the transmissivity

constant velocity  
shift

$$\Delta v_{\text{int}} = 100 \text{ km s}^{-1}$$

evolving velocity  
shift

$$\Delta v_{\text{int}} = 100[(1+z)/7]^{-3} \text{ km s}^{-1}$$

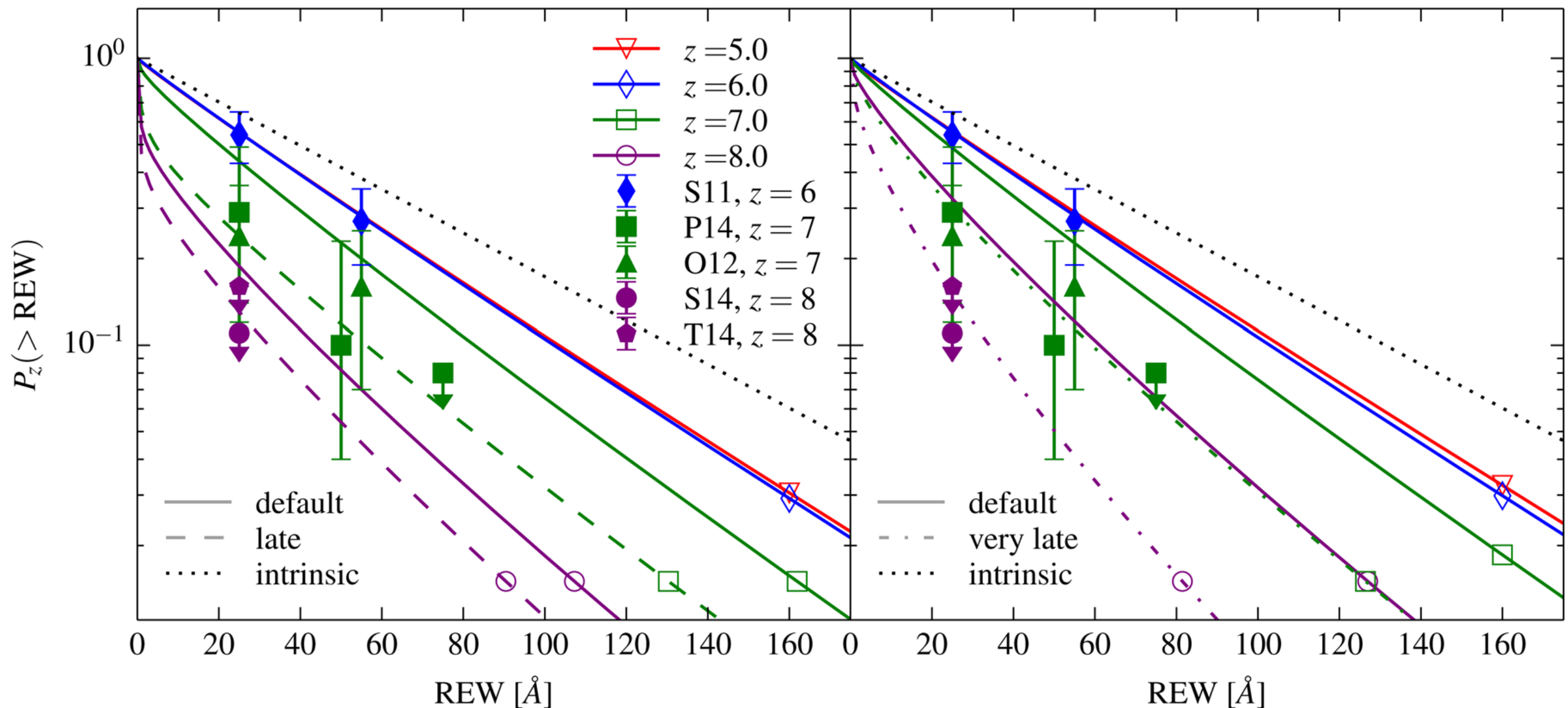


# Cumulative Lyman- $\alpha$ equivalent width distribution

evolving velocity shift

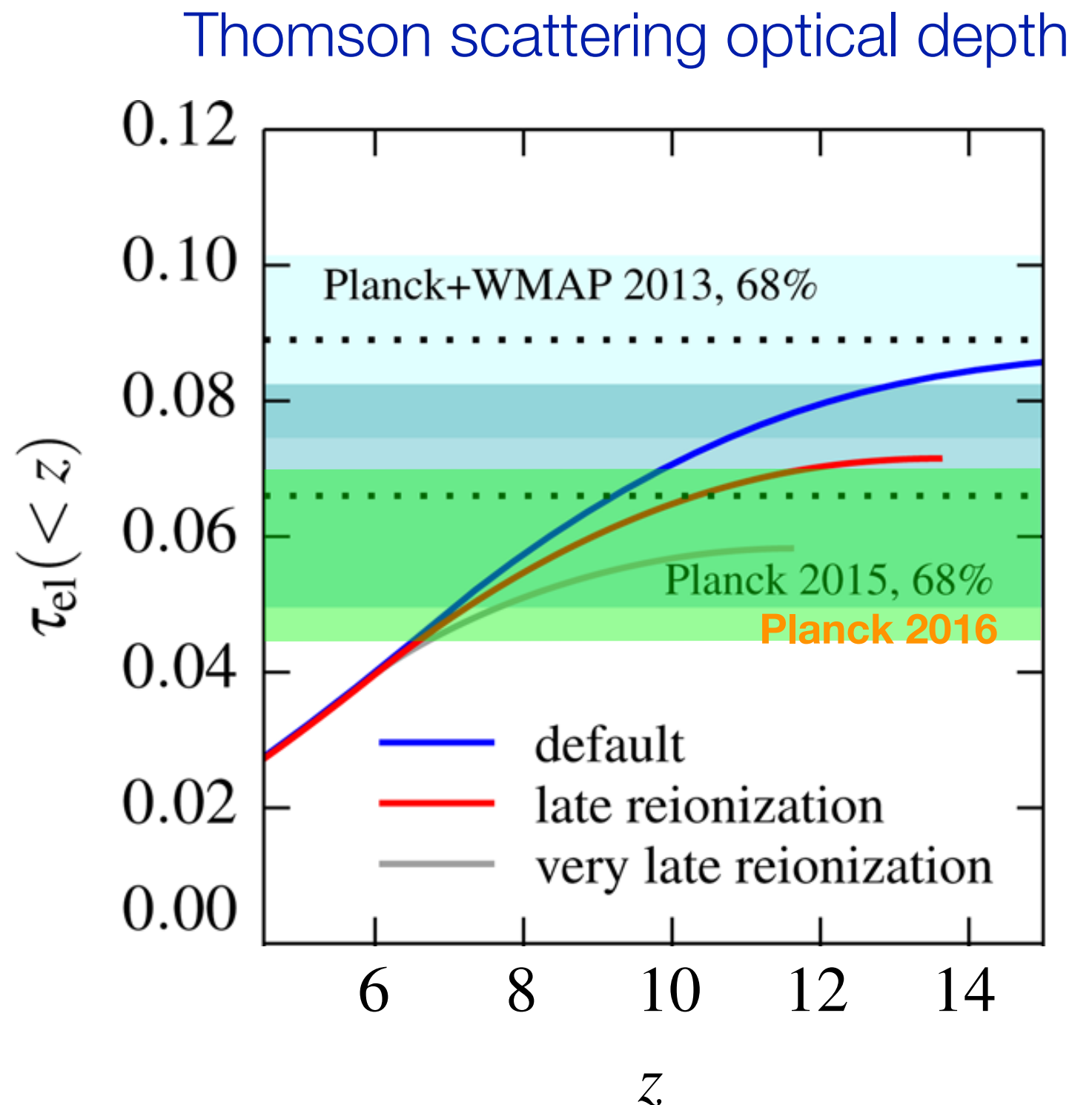
uncorrelated LAE positions  
 $\Delta v_{\text{int}} = 100[(1+z)/7]^{-3} \text{ km s}^{-1}$

strongly correlated LAE positions  
 $\Delta v_{\text{int}} = 100[(1+z)/7]^{-3} \text{ km s}^{-1}$



# Consistency with latest Planck constraints

- optical depth in Planck 2015/16 came down
- consistent with what we need for the Lyman- $\alpha$  emitters
  - > reionization ends around  $z \sim 6$  and is not too extended





# Summary

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- **Lyman- $\alpha$  emitters:**
  - favour a late and not too extended reionization history (finishing at  $z \sim 6$ )
  - evolution of intrinsic velocity offsets may be important
- **CMB:**
  - Planck 2015/16 find lower optical depths  $\rightarrow$  late reionization in agreement with LAEs